



Cornell University  
Laboratory for Elementary-Particle Physics



USPAS course on  
*Recirculated and Energy Recovered Linacs*

*Ivan Bazarov, Cornell University*

*Geoff Krafft, JLAB*

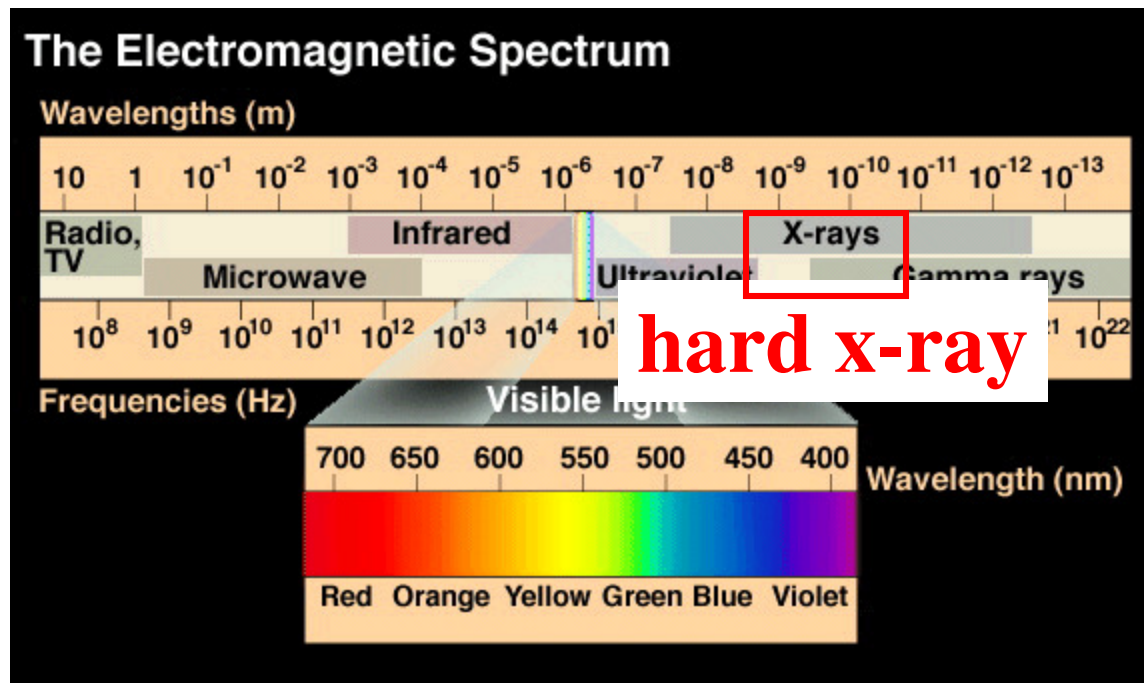
ERL as a X-ray Light Source





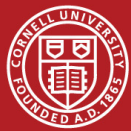
- Introduction
  - Light sources landscape
  - General motivation for a new light source
- Comparison with storage rings and XFELs
  - Storage ring basics
  - XFEL basics
- Science case examples
- Cornell ERL plan





- Relativistic free electrons – the only medium for tunable light production in widest spectral range
- Hard x-ray range is the subject of this talk





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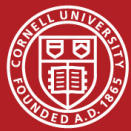
## ERL light source idea

Third generation light sources are storage ring based facilities optimized for production of high brilliance x-rays through spontaneous synchrotron radiation. The technology is mature, and while some improvement in the future is likely, one ought to ask whether an alternative approach exists.

Two orthogonal ideas (both linac based) are XFEL and ERL. XFEL will not be spontaneous synchrotron radiation source, but will deliver GW peak powers of transversely coherent radiation at very low duty factor. The source parameters are very interesting and at the same time very different from any existing light source.

ERL aspires to do better what storage rings are very good at: to provide radiation in quasi-continuous fashion with superior brilliance, monochromaticity and shorter pulses.



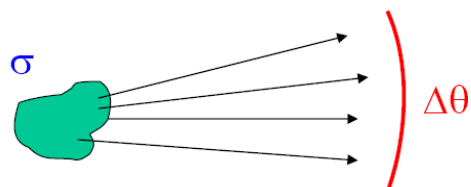


## Brilliance / Spectral Brightness

FLUX OF PHOTONS IN UNIT SPECTRAL RANGE

(SOURCE AREA) X (BEAM DIVERGENCE)

Units: Photons/s/mm<sup>2</sup>/mrad<sup>2</sup>/0.1% bandwidth



$$B = \frac{\Phi}{\sigma \Delta\theta_h \Delta\theta_v}$$

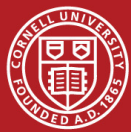
- **Average brightness:** measure of transversely coherent flux

$$F_c = B_{avg} \left( \frac{\lambda}{2} \right)^2$$

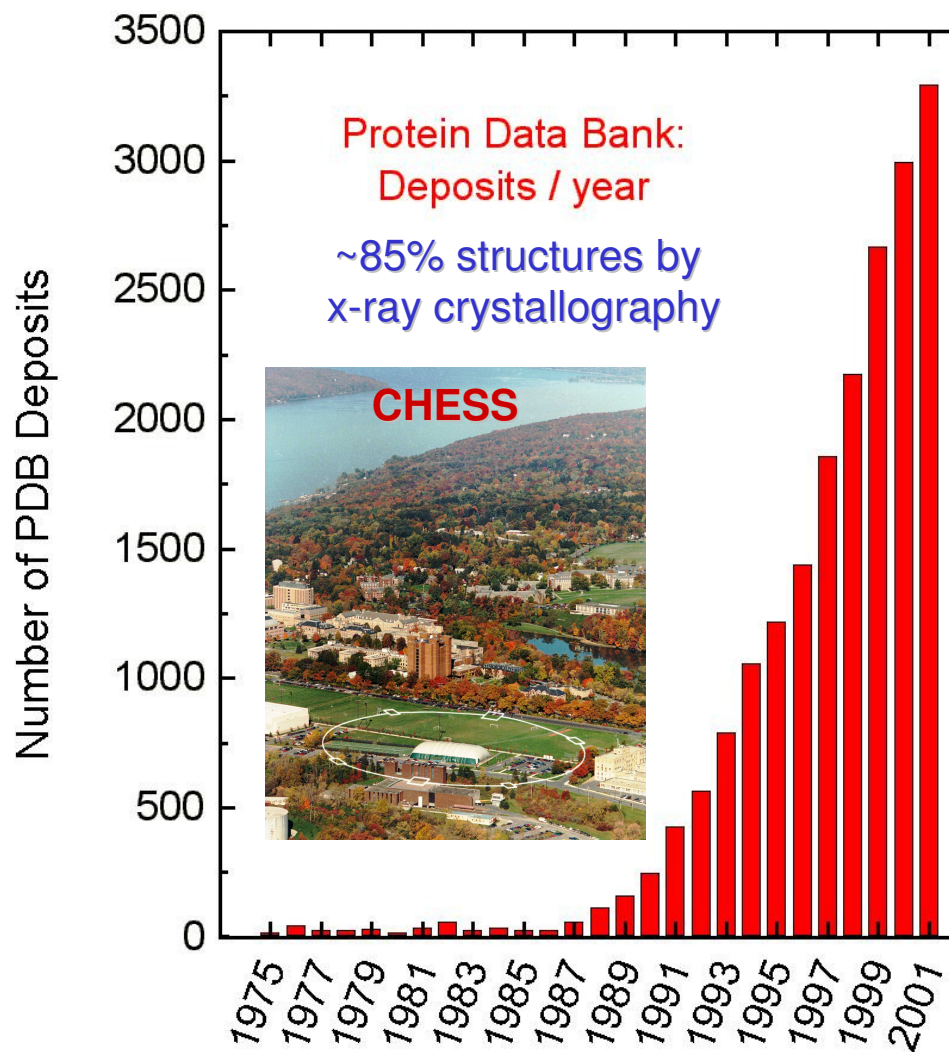
- **Peak brightness:** proportional to the number of photons per coherence volume in 6D phase space  $\equiv$  the photon degeneracy

$$\Delta_c = B_{peak} \left( \frac{\lambda}{2} \right)^3 \frac{\Delta\lambda}{\lambda} \frac{1}{c}$$

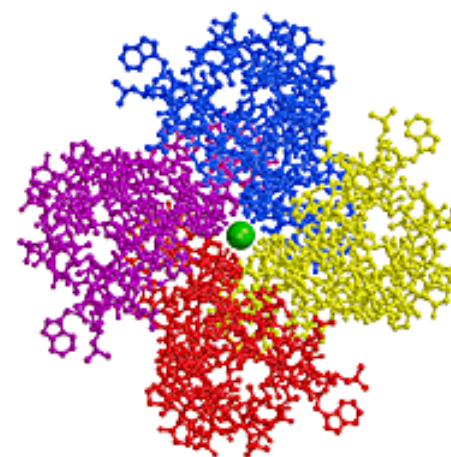




# Demand for X-rays

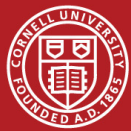


## Ion channel protein



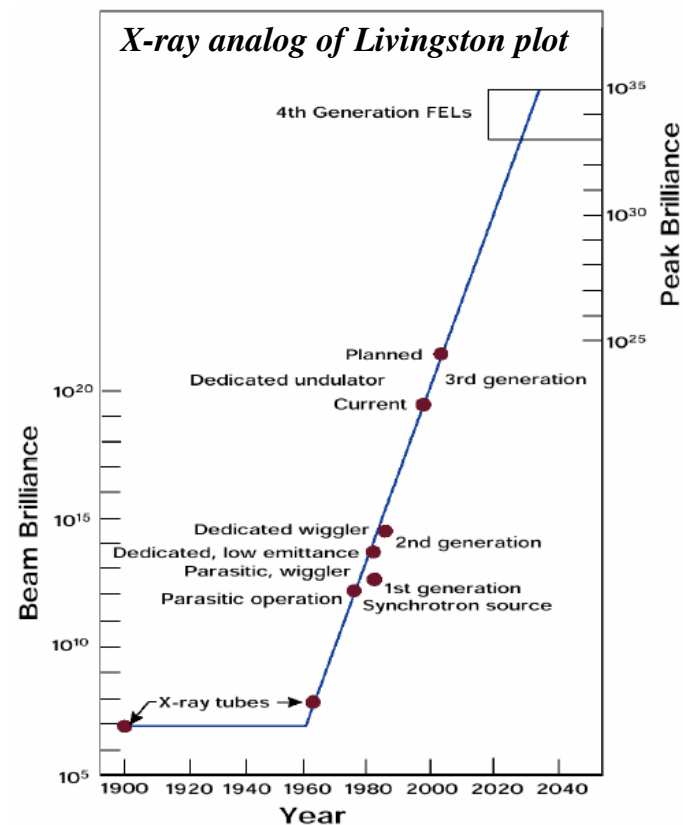
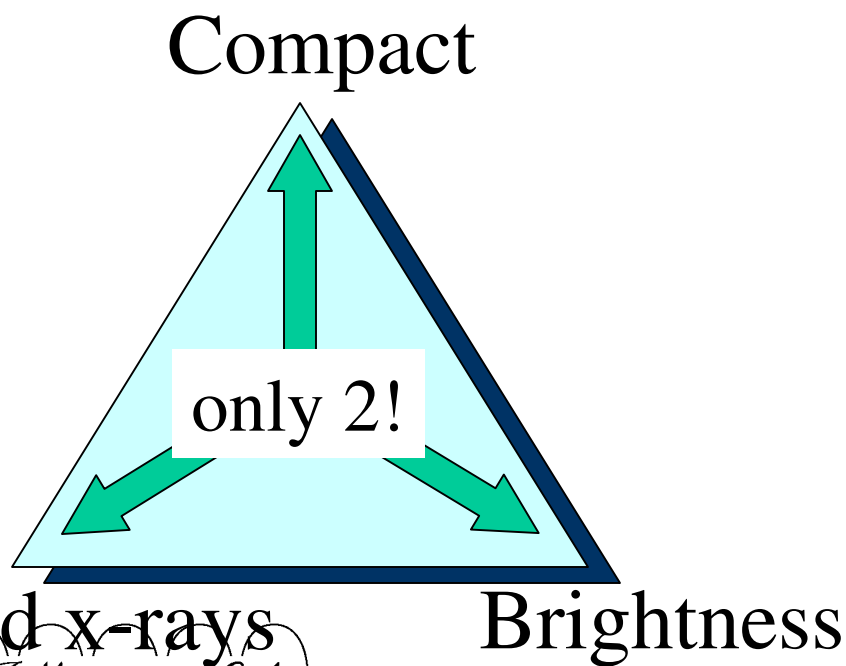
2003 Nobel Prize  
in Chemistry:

Roderick MacKinnon  
(Rockefeller Univ.)  
1<sup>st</sup> K<sup>+</sup> channel structure  
by x-ray crystallography  
based on CHES data (1998)



# Three frontiers in source dev.

- Avg. brightness
- Short pulses ( $< \text{ps}$ ) & peak brightness
- Compactness

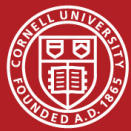






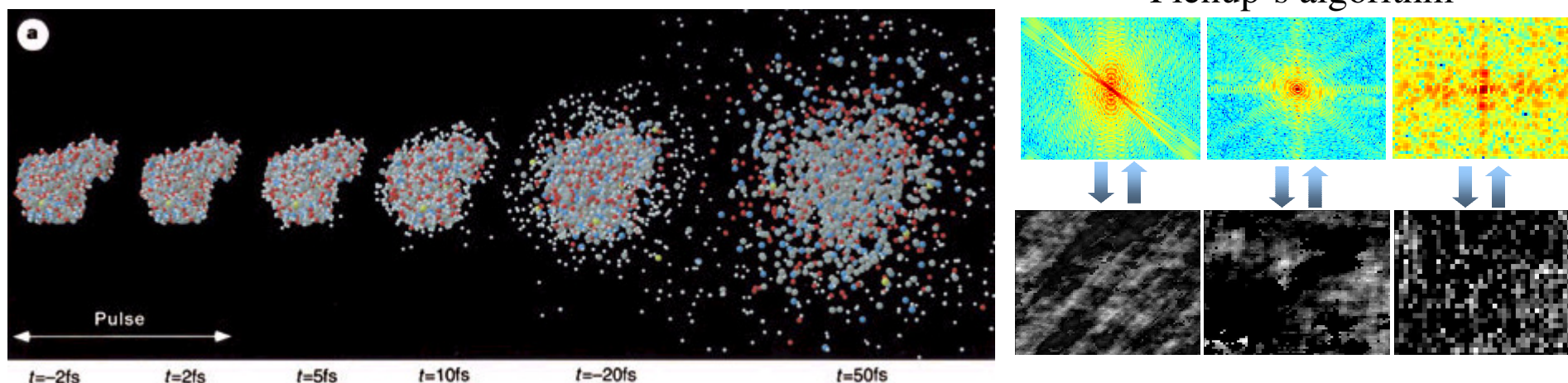
- About 70 light sources worldwide based on **storage ring** technology (**VUV to hard X-rays**), new ones are being built / designed
- 22 **FELs** operational, some as scientific research instruments (**far IR to VUV**)
- 3 **XFELs** in construction / committed to, plus half a dozen in CDR or earlier stages (**soft to hard X-rays**)
- 3 labs seriously consider building **ERL** as a **hard X-ray** light source





# Exp1: 'explosive' proteins

R. Neutze, et al., *Nature*, **406**, 752



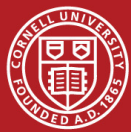
Briefly:

calculations were done for T4 lysozyme (diameter 32 Å,  $N_C \sim 1000$ );  
flux  $4 \times 10^6$  X-rays/Å<sup>2</sup> with  $\sim 2000$  primary ionization events;  
elastically scattered  $\sim 200$  photons.

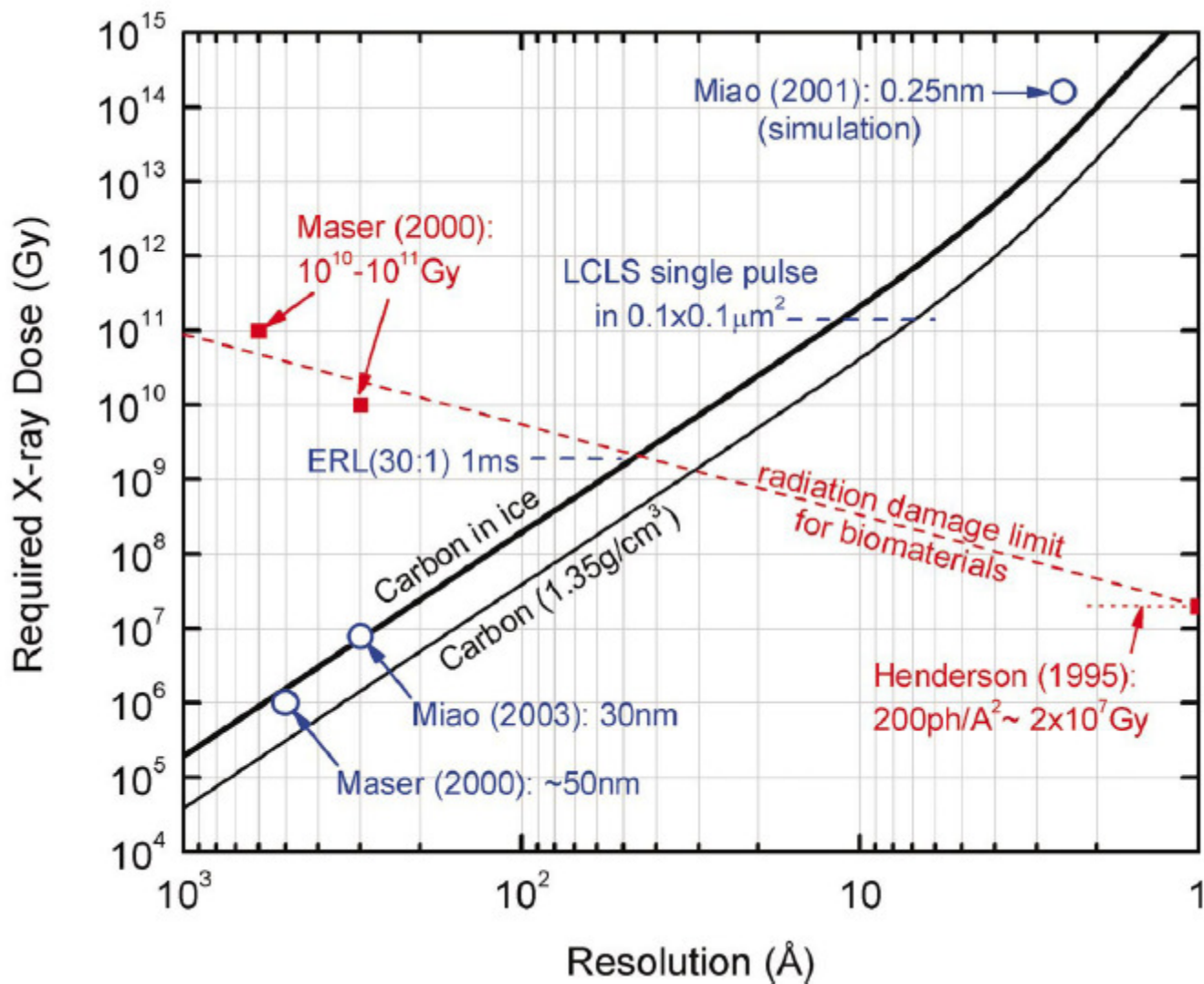
If pulse is sufficiently short ( $< 10$  fs),  $5 \times 5 \times 5$  lysozyme nanocrystal will scatter to  $< 2$  Å resolution.

*Key feature: sufficiently short X-ray pulse can beat Henderson's limit of radiation damage (200 x-ray photons /Å<sup>2</sup>)*



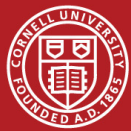


# Radiation damage: biomaterials



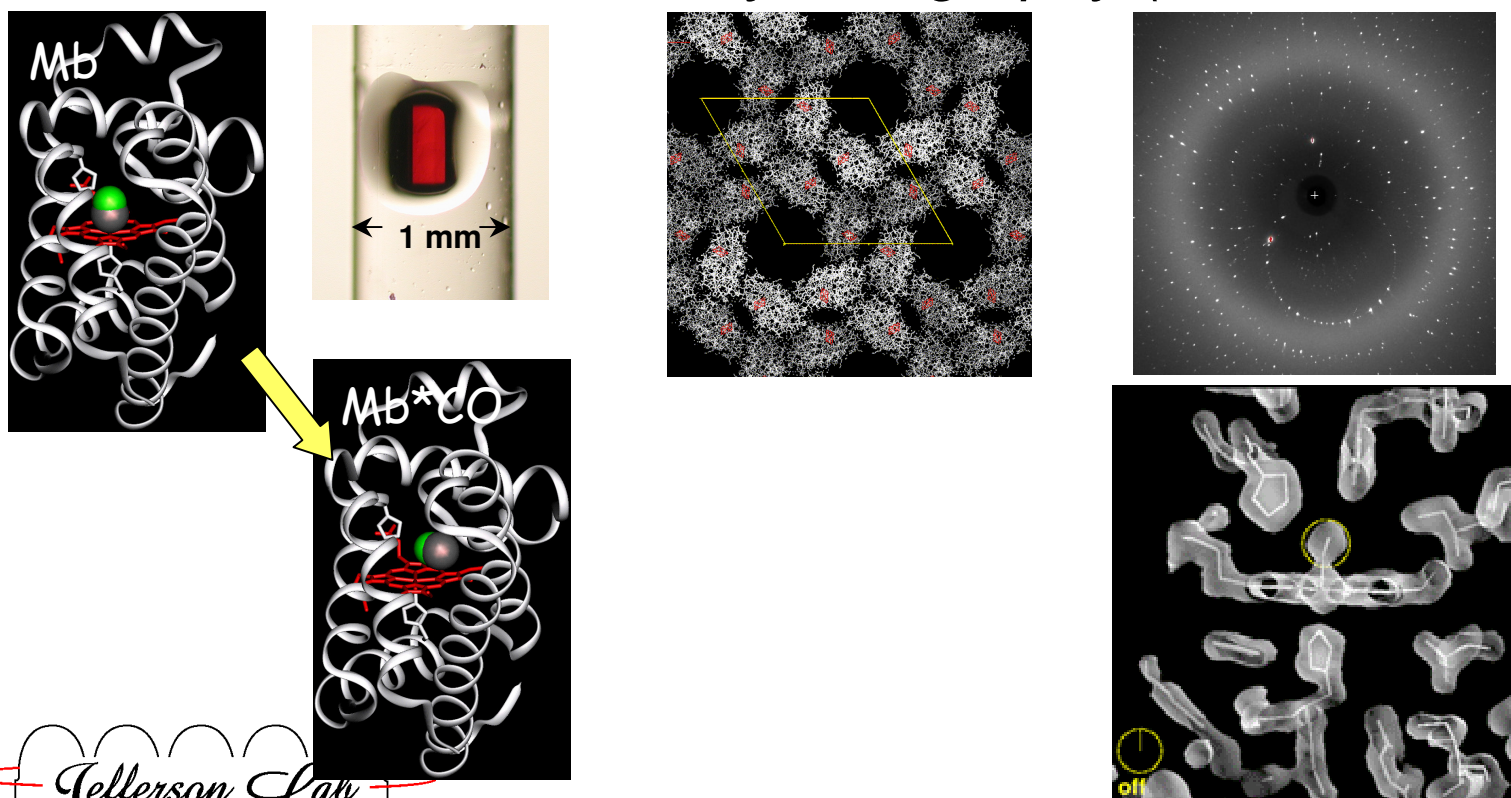
Shen, Bazarov, Thibault, J. *Synch. Rad.*, Vol. 11 (2004) 432



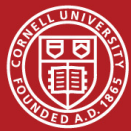


*Broad class of pump-probe experiments providing structural (core  $e^-$ 's) conformational changes in the initial stages (mol. vibrational timescale 10's fs) of photo-induced reactions*

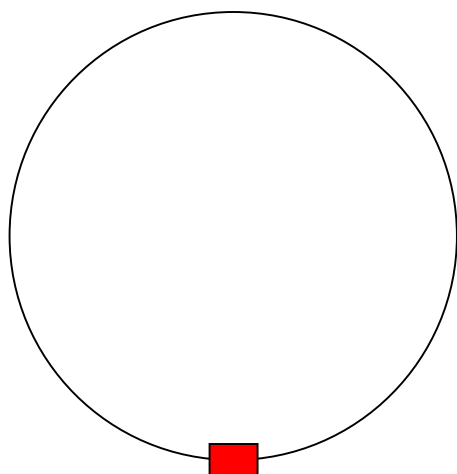
## Time-resolved Laue Crystallography (Phil Anfinrud)







## SR



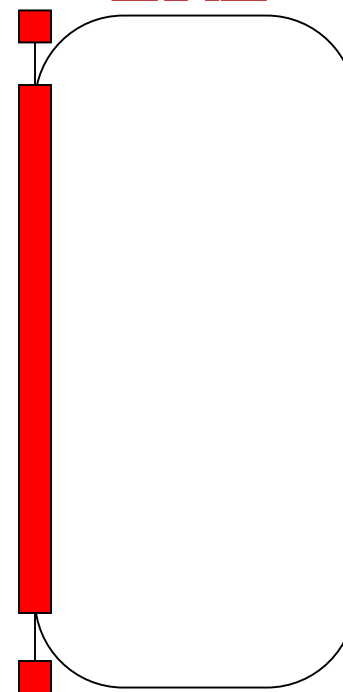
- efficient
- avg. brightness
- many beamlines
- workhorse technology

## XFEL



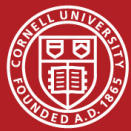
- peak brightness
- short pulse
- few beamlines
- new user-base

## ERL

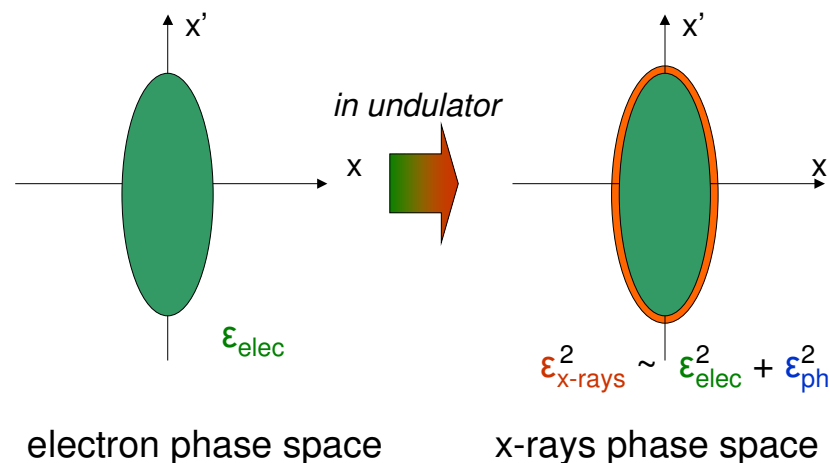


- avg. brightness
- short pulse
- many beamlines
- existing user-base





- In properly tuned undulator x-ray phase space is convolution of e-beam with diff. limit



- Goal:** for 1 Angstrom  $\rightarrow \epsilon_x \sim \lambda/4\pi = 8 \text{ pm}$  geometric, or  $\epsilon_{nx} = 0.08 \text{ }\mu\text{m}$  if energy is 5GeV
- E.g. best storage ring performance as of today:  
 $\epsilon_x / \epsilon_y = 3000 / 15 \text{ pm}$

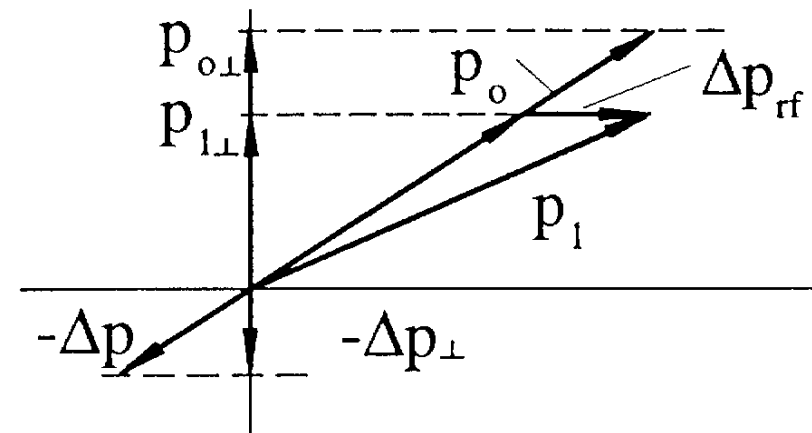
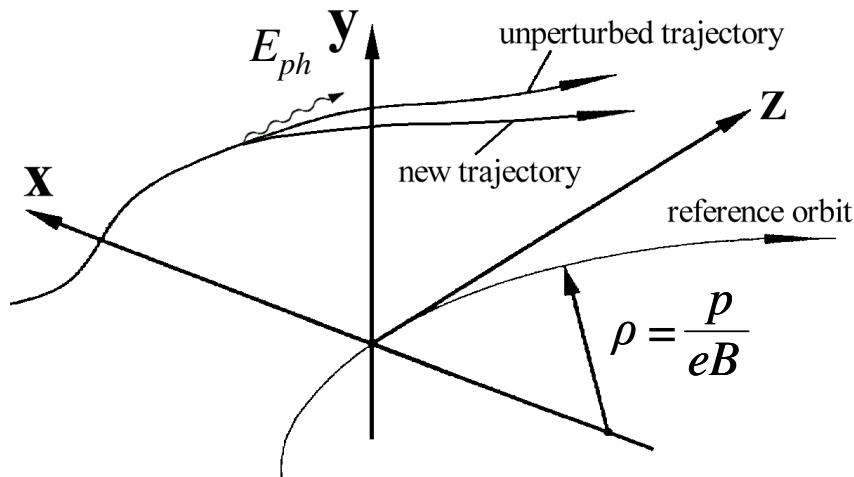


## Equilibrium

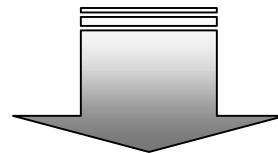
Quantum Excitation

vs.

Radiative Damping



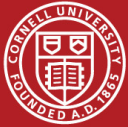
$$\frac{d\sigma_E^2}{dt} \sim \dot{N}_{ph} E_{ph}^2$$



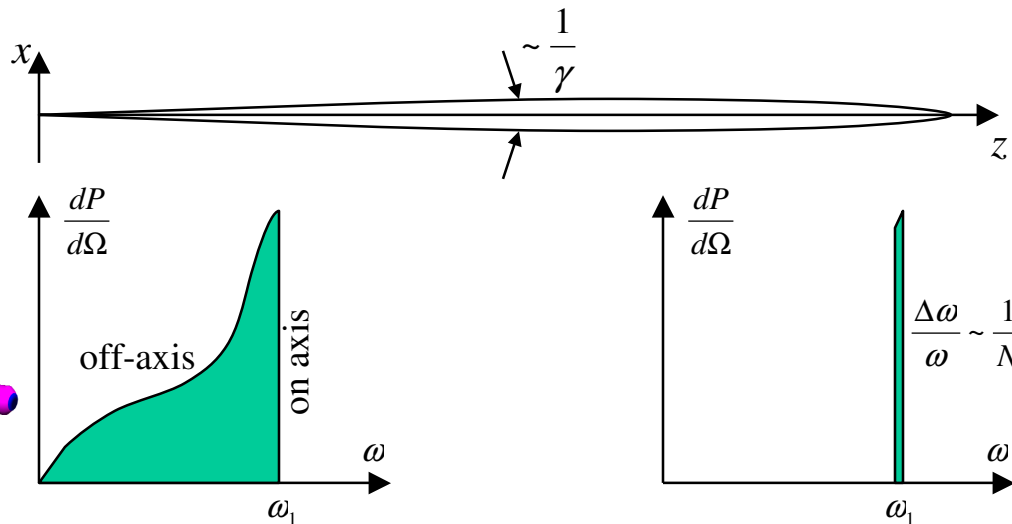
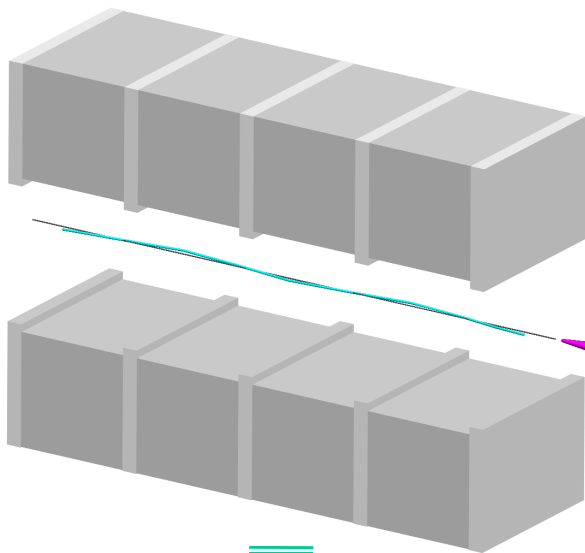
Emittance (hor.), Energy Spread, Bunch Length

*Tighter focusing (higher tune) → stronger 6-poles for chromaticity correction → smaller dynamic aperture & lifetime*



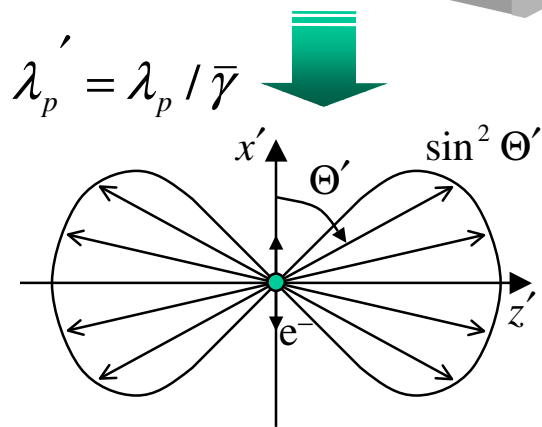


# Basics of sync. rad. production

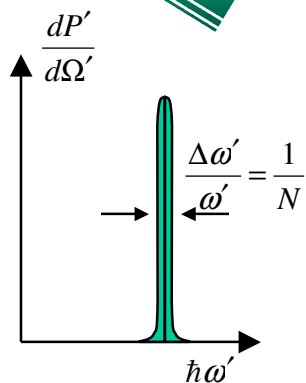


back to lab frame

after pin-hole aperture



in e<sup>-</sup> frame



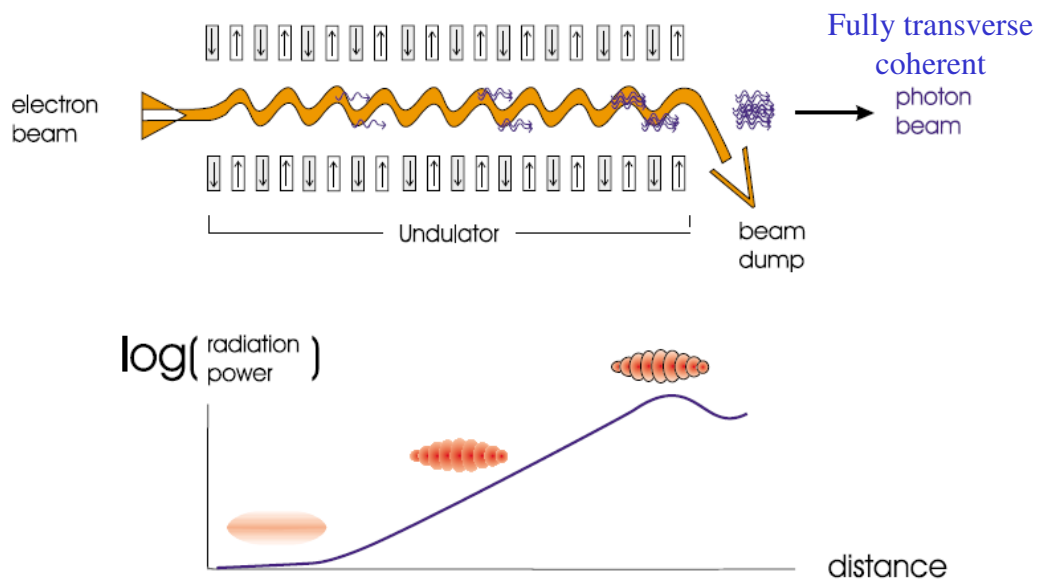
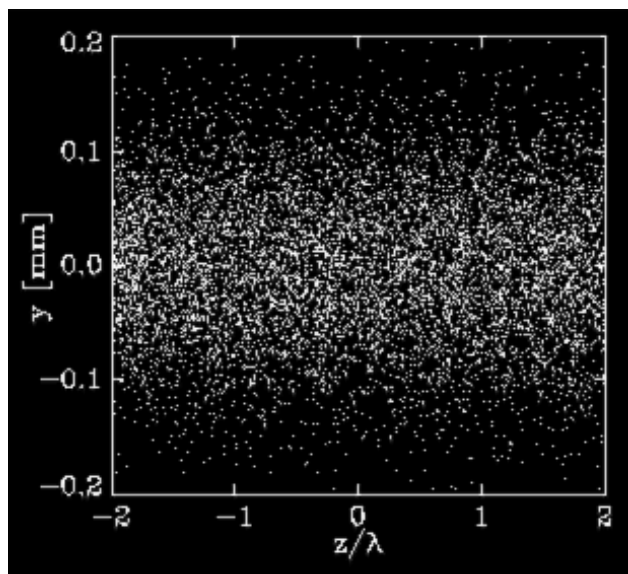
$$\lambda_n = \frac{\lambda_p}{2\gamma^2 n} \left(1 + \frac{1}{2} K^2 + \gamma^2 \theta^2\right)$$

$$\frac{\Delta\lambda}{\lambda_n} \sim \frac{1}{nN_p}$$

(for fixed  $\theta$  only!)







## Prerequisites for $e^-$ -bunch:

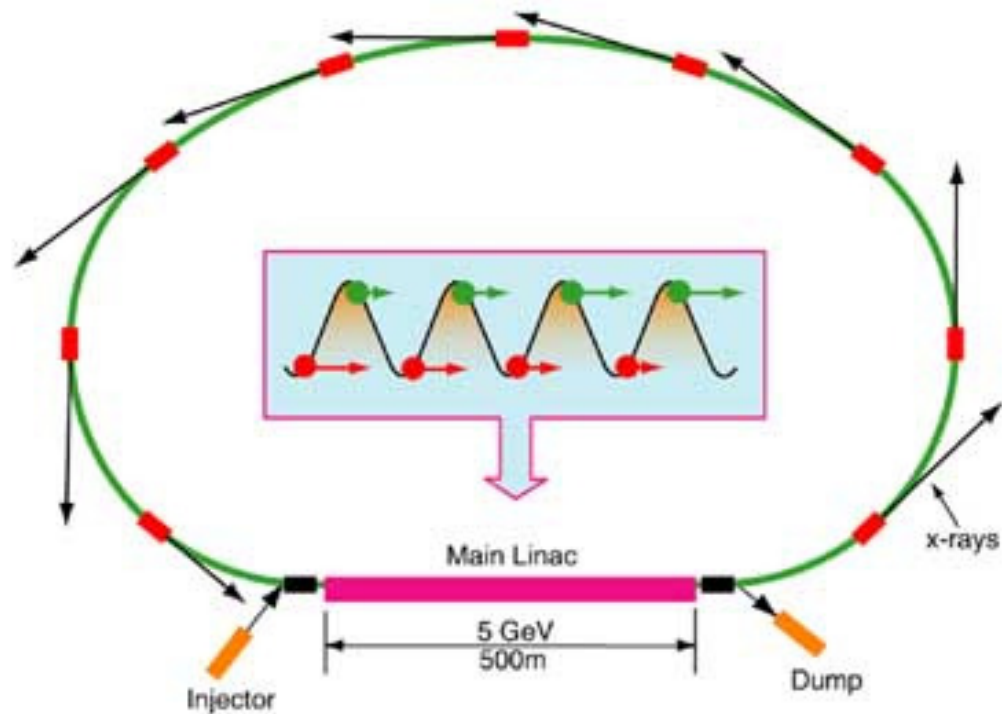
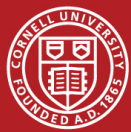
diffraction-limited emittance

peak current 3-5 kA

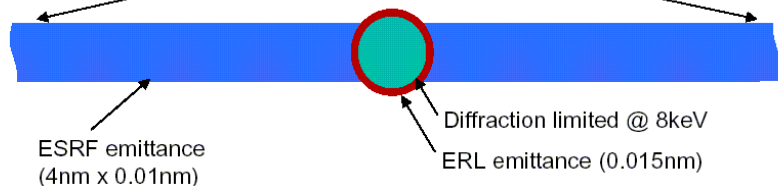
energy spread  $10^{-4}$

*Intense relativistic electron bunch  
becomes effective gain medium  
(e.g. use seed / amplifier setup)*

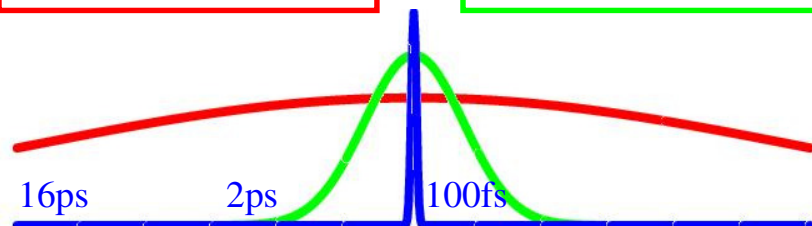




ESRF 5GeV@100mA → ERL 5GeV@100mA



ESRF 5GeV@100mA → ERL 5GeV@100mA



**Much smaller (×100) horizontal emittance**

**Much shorter (×100) pulses**





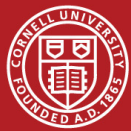
- Brightness figure of merit (FOM)  $\frac{I}{(\epsilon_x + \lambda/4\pi)(\epsilon_y + \lambda/4\pi)}$  for 1 Å

Light source	I (A)	$\epsilon_x$ (nm-rad)	$\epsilon_y$ (nm-rad)	FOM (A/nm <sup>2</sup> /rad <sup>2</sup> )
ESRF	0.2	3.7	0.010	3.0
Petra-III	0.1	1.0	0.010	5.5
NLSL-II*	0.5	1.54	0.008	20.24
UHXS(ESRF)	0.5	0.2	0.005	185.6

- 5 GeV ERL to achieve the same brightness per m of ID as Petra-III / NLSL-II / UHXS(ESRF) needs **1.3 / 0.6 / 0.15** μm rms **normalized** emittance for **80 pC** bunch (0.1 A average current at 1.3 GHz bunch rep rate) assuming no emittance degradation downstream
- Comparison: ILC norm. emit.  $\sqrt{\epsilon_{nx}\epsilon_{ny}} = \mathbf{0.6}$  μm for **3.2 nC**



\* without use of damping wigglers



# Cornell plans

