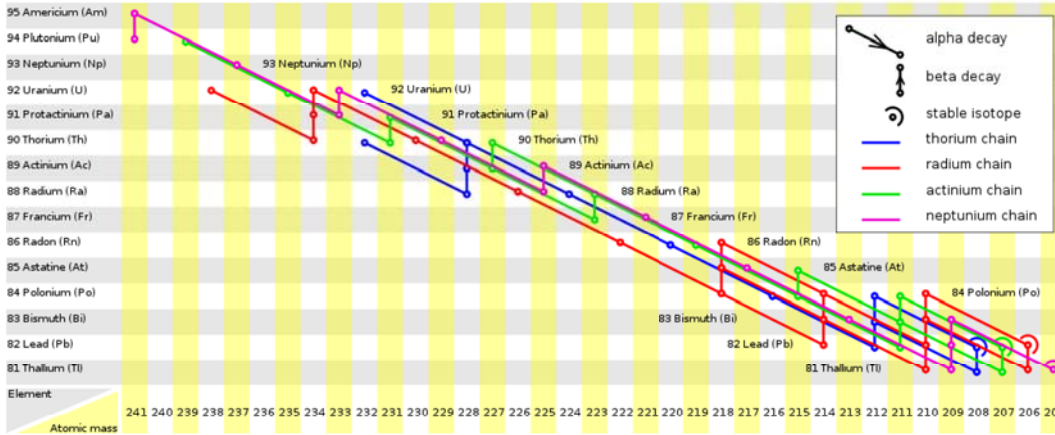
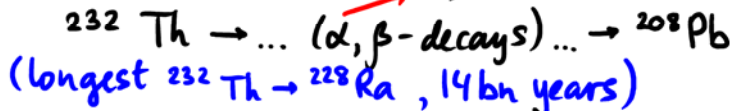


# Nuclear dating

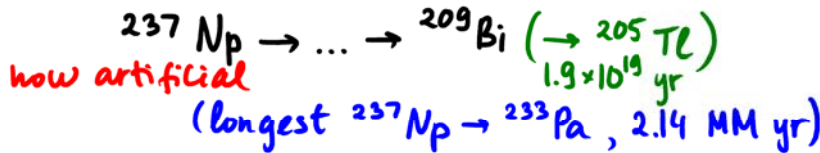
## ① Natural decay chains



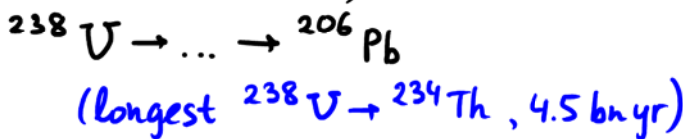
a) Thorium series ( $A=4n$ )



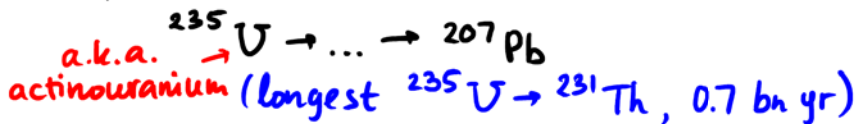
b) Neptunium series ( $A=4n+1$ )



c) Uranium (or radium) series ( $A=4n+2$ )



d) Actinium series ( $A=4n+3$ )



## ② Radiometric dating

E.g. U/Pb, Sm/Nd, Rb/Sr, ...

Earth age  $\rightarrow$  4.5 bn yr old (universe  $\sim$  14 bn yr old)

Early estimate:

Kelvin - the earth starts as molten rock,  
Now temperature gradient  $1^\circ\text{F}$  every 50 ft down.  
 Assume surface  $T_s \sim 0^\circ\text{C}$ , how long does it  
 take to establish  $1/50^\circ\text{F}/\text{ft}$  gradient?

Answer: 25 million years (< Kelvin's initial number)

Why? Radioactivity heats up the earth, roughly cancels the heat flow from the earth.

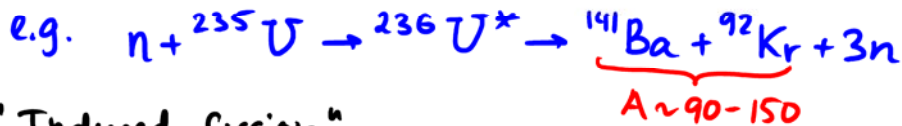
- ③  $^{14}\text{C}$  dating [human-history-relevant timescale]
- \* used for [once] living substances
  - \* cosmic rays in upper atmosphere  $\rightarrow n$   
 $^{14}\text{N}(n,p)^{14}\text{C}$ , mostly in form of  $\text{CO}_2$
  - \*  $^{14}\text{C} \rightarrow ^{14}\text{N} + e^- + \bar{\nu}_e$ ,  $\beta$ -decay with  $5370\text{yr}$
  - \* atmosphere: contains  $\sim 600\text{bn/mole}$  of  $^{14}\text{C}$
  - \* once dead,  $^{14}\text{C}$  decays

Shroud of Turin  $\sim 1300\text{ A.D.}$

Dead Sea scrolls  $\sim 0-200\text{ A.D.}$

## Nuclear energy

① Fission ( $^{235}\text{U}$ ,  $^{233}\text{U}$ ,  $^{239}\text{Pu}$ )



"Induced fission":

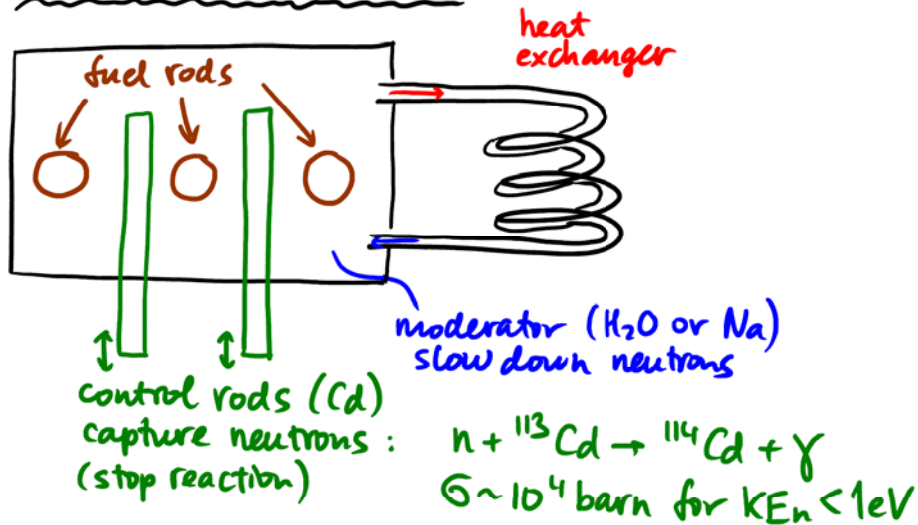


cross-section for thermal  $n$ :  $\sim 500\text{ barns}$

Q: How to thermalize  $n$ ?

A: use elastic collision with same mass (H).  
(e.g. paraffin, water)

## Nuclear fission reactor

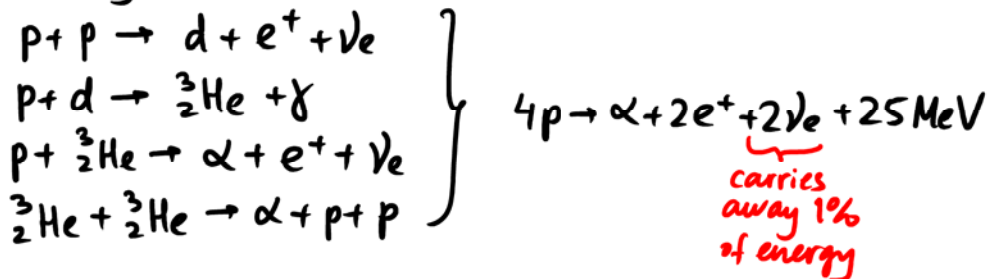


Fuel: typically enriched  ${}^{235}\text{U}$  (from 0.7 to few%)  
natural  ${}^{238}\text{U}$  99.3%

Ex. 1kg of enriched  $\text{U} \sim 24 \times 10^9 \text{ W-h}$   
USA electric. consumption:  $\sim 4 \times 10^{15} \text{ W-h/year}$   
world  $\sim 17 \times 10^{15} \text{ W-h/year}$   
"economic"  ${}^{235}\text{U}$  (\$60/lb), to last  $\sim 100$  yr  
with present consumption rate

## ② Fusion

\* primary fusion reaction in the Sun:



\* controlled fission easy; fusion not

Q: Why?

A: neutron  $\rightarrow$  fission; Coulomb repulsion (fusion)

\* need  $T \sim 10^7 - 10^8 \text{ K}$ ,  $\Rightarrow$  plasma

dd fusion:  $d + d \rightarrow {}^3_2\text{He} + \text{n}$ , Q-val: 3.3 MeV

$d + d \rightarrow t + p$  escape from plasma 4 MeV

dt fusion:  $d + t \rightarrow \alpha + \text{n}$  17.6 MeV

## Conditions for fusion

### a) critical ignition temperature

$$\text{power from fusion} = \text{cooling}$$

neutrons      x-rays

$$dd: 4 \times 10^8 \text{ K (35 keV)}$$

$$dt: 4 \times 10^7 \text{ K (4 keV)}$$

### b) Lawson criterion

$$\overset{\text{density}}{\rightarrow} n \tau > C$$

cooling or  
confinement time

$$dd: C = 10^{16} \frac{\text{s}}{\text{cm}^3}, \quad dt: 10^{14} \frac{\text{s}}{\text{cm}^3}$$

Two approaches at the moment:

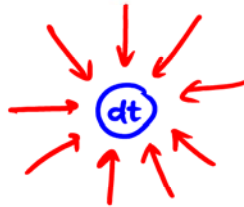
\* tokamak (Russian for toroidal chamber with magnetic coils)

contain plasma in magn. field,  
heat up by RF induction

e.g.  $\tau \sim 1 \text{ s}$ ,  $\Rightarrow n > 10^{14} \text{ cm}^{-3}$  for dt fusion  
(5 orders of magn. less dense than gas at STP)

ITER - international thermonuclear experimental reactor

\* inertial confinement  
200 high power lasers  
(~500 TW, ~2 MJ total)



compress few mg dt target  
to  $n \sim 10^{25} \text{ cm}^{-3}$   
(only  $\tau \sim 10^{-11} \text{ s}$  is needed)

NIF - national ignition facility  
(Lawrence Livermore Nat'l Lab)

## Elementary particles

Search for basic building blocks of matter

~100 chemical elements: periodic table  
explained by QM. All elements are made of  
3 basic constituents:  $e^-$ ,  $p$ ,  $n$

~1960's: powerful accelerators

$e^-e^+$  colliders (lepton) } main HEP  
 $p\bar{p}$  colliders (hadron) } discovery tools

>400 particles discovered. Half-lives  $10^{-6}$  to  $10^{-23}$  s.  
"particle zoo", no seeming order (cf. chemical  
elements before QM)

~1970: quark.  $p, n$  are composite

QM of quarks  $\rightarrow$  quantum chromodynamics

Standard Model: explains "zoo". Includes:

\* strong force

\* electroweak

\* elementary particles: spin  $\frac{1}{2}$  - quarks and leptons  
integer spin - "field" particles  
(gauge bosons)

Issues with SM:

\* gravity not included

\* needs 25 numerical const (masses of particles,  
coupling const)

\* astrophysics (last 3 decades):

baryonic matter ~ 5%

(non EM inter.) dark matter ~ 23%

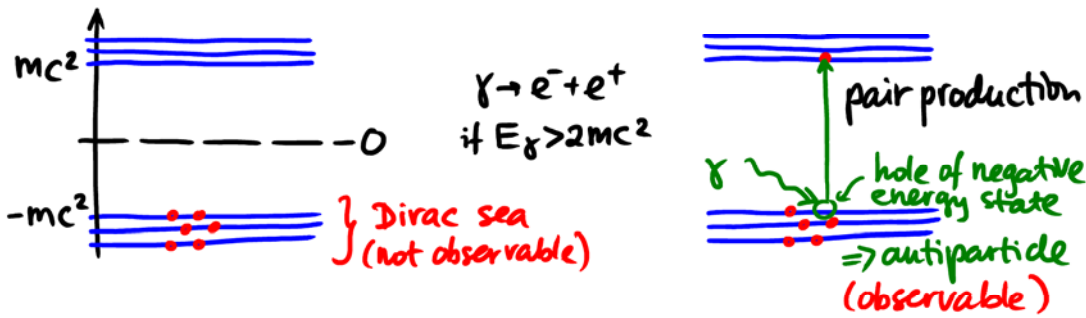
(neg. pressure) dark energy ~ 72%

### ① Antiparticles

consequence of relativistic QM (Dirac eqn. for fermions)

$$E^2 = (pc)^2 + (mc^2)^2, \text{ or } E = \pm \sqrt{p^2c^2 + m^2c^4}$$

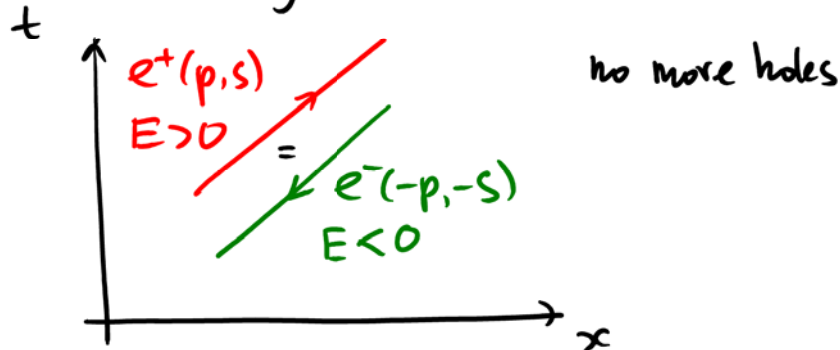
Dirac: the negative energy states completely filled in vacuum



Every fund. fermion: has an antiparticle  
all intrinsic quant. numbers are opposite  
(absence of a particle with negative energy)

Problem with Dirac sea: infinite (negative) energy of vacuum;  $e^+/e^-$  are not treated symmetrically (holes)

A better picture: Stueckelberg-Feynman  $E < 0$  solns of a particle are  $E > 0$  solns of its antiparticle moving backwards in time



## ② Crossing symmetry

if  $A + B \rightarrow C + D$  reaction possible:  
cross any particle and change to antiparticle

$$\left. \begin{array}{l} A \rightarrow \bar{B} + C + D \\ \bar{C} + \bar{D} \rightarrow \bar{A} + \bar{B} \end{array} \right\} \text{also possible}$$

Ex.  $n \rightarrow p + e^- + \bar{\nu}_e$   
 $n + \bar{\nu}_e \rightarrow p + e^-$   
 $n + \bar{p} \rightarrow e^- + \bar{\nu}_e$

neutron  $\beta$ -decay  
inverse  $\beta$ -decay  
neutron/antiproton annihilation

$\swarrow$  used to detect neutrinos  
 $(p + \bar{\nu}_e \rightarrow n + e^+)$



## Elementary Particles

Quarks	$u$ up	$c$ charm	$t$ top	Force Carriers	
	$d$ down	$s$ strange	$b$ bottom		
Leptons	$\nu_e$ electron neutrino	$\nu_\mu$ muon neutrino	$\nu_\tau$ tau neutrino		$Z$ Z boson
	$e$ electron	$\mu$ muon	$\tau$ tau		$W$ W boson
	I	II	III		

Three Families of Matter

- ③ Leptons (elementary)  
 spin  $\frac{1}{2}$ ; 6 types  
 3 families (generations): electron, muon, tau

generation	mass (MeV/c <sup>2</sup> )	charge (e)	flavor			
			e	$\mu$	$\tau$	
1	$e^-$	0.511	-1	1	0	0
	$\nu_e$	$< 2 \times 10^{-6}$	0	1	0	0
2	$\mu^-$	105.7	-1	0	1	0
	$\nu_\mu$	$< 0.17$	0	0	1	0
3	$\tau^-$	1777	-1	0	0	1
	$\nu_\tau$	$< 15.5$	0	0	0	1

+ corresponding antiparticles

- \* leptons interact via electroweak interaction; no color (= no strong interaction)
- \* lepton number conservation in all reactions (antilepton counts as -1)
- \* lepton flavor conservation: almost always explains why  $e^- + \nu_e$  in  $\beta^-$ -decay (electron flavor is 0)  
 But "neutrino oscillations": only  $\frac{1}{3}$  to  $\frac{1}{2}$  of  $\nu_e$  flux from the Sun is observed.  $\nu_e$  oscillates into  $\nu_\mu$  and  $\nu_\tau$ .
- \* leptons decay into lighter ones:

$$\tau^- \rightarrow e^- + \bar{\nu}_\tau + \bar{\nu}_e \quad \left. \begin{array}{l} \tau^- \rightarrow \mu^- + \bar{\nu}_\tau + \bar{\nu}_\mu \end{array} \right\} \tau^- \text{ half-life } 3.3 \times 10^{-13} \text{ s}$$

$$\mu^- \rightarrow e^- + \bar{\nu}_\mu + \bar{\nu}_e \quad 2 \times 10^{-6} \text{ s half-life}$$

$e^-$  is the lightest of leptons  $\Rightarrow$  stable

\* 4<sup>th</sup> generation?? at least 100.8 GeV/c<sup>2</sup> and 45 GeV/c<sup>2</sup>  
 Not observed (yet). lepton neutrino

#### ④ Quarks (elementary)

spin 1/2 ; 6 flavors ; 3 families

\* quarks interact via strong and electroweak forces

\* quark confinement: not found free in nature, only bound as hadrons

\* fractional charge

	mass (MeV/c <sup>2</sup> )		charge (e)	flavor				
	bare	bound		Isospin	S	C	B	T
up (u)	1-5	310	2/3	1/2				
down (d)	3-9	310	-1/3	-1/2				
strange (s)	75-170	483	-1/3		-1			
charm (c)	1150-1350	1500	2/3			1		
bottom (b)	4000-4400	4700	-1/3				-1	
top (t)	174.3 × 10 <sup>3</sup>	175 × 10 <sup>3</sup>	2/3					1

\* quark number conservation in all reactions

\* quark flavor conservation in strong and electromagn. but not in weak

#### ⑤ Hadrons - composite systems of quarks

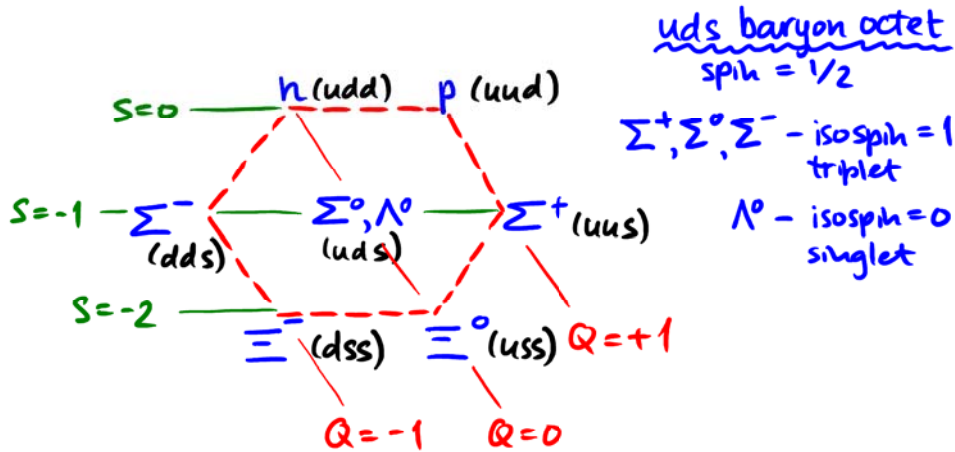
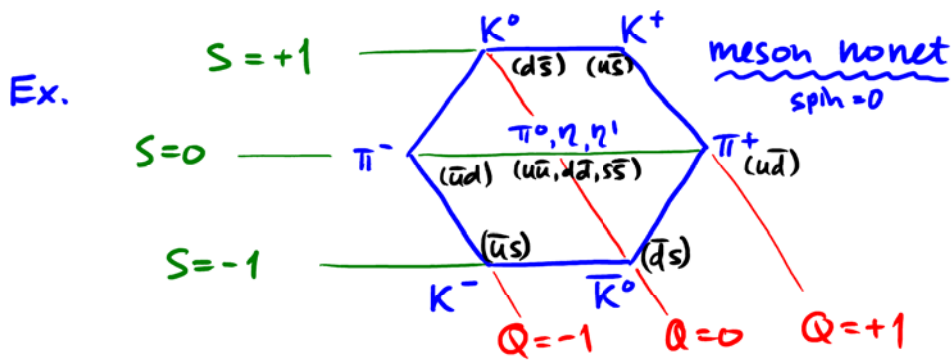
mesons - quark and antiquark, integer spin

baryon - 3 quarks, half-integer spin

measured first, before discovery of quarks  
 (remember: quark confinement)

tetraquark - made of 4 quarks (possibly seen in Fermilab)  
pentaquark - " " " " " " ( " " JLAB)





## Color

- \* Additional quant. number required to make wavefn antisymmetric for baryon decuplet (spin = 3/2). red, green, blue
- \* like electric charge for strong force
- \* each quark must come in 3 colors
- \* naturally occurring baryons are colorless (RGB)  
mesons — " — (color-anticolor)
- \* leptons have no color