



Sextupoles cause nonlinear dynamics, which can be chaotic and unstable.

$$\begin{pmatrix} x_{n+1} \\ x'_{n+1} \end{pmatrix} = \underline{M}_0 \left[ \begin{pmatrix} x_n \\ x'_n \end{pmatrix} - \frac{k_2 l_s}{2} \begin{pmatrix} 0 \\ x_n^2 \end{pmatrix} \right] \quad \begin{pmatrix} x_{n+1} \\ x'_{n+1} \end{pmatrix} = \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \hat{x}_n \\ \hat{x}'_n \end{pmatrix}$$

$$\begin{pmatrix} \hat{x}_{n+1} \\ \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu & \sin \mu \\ -\sin \mu & \cos \mu \end{pmatrix} \left[ \begin{pmatrix} \hat{x}_n \\ \hat{x}'_n \end{pmatrix} - \frac{k_2 l_s}{2} \sqrt{\beta} \begin{pmatrix} 0 \\ \beta \hat{x}_n^2 \end{pmatrix} \right]$$

$$\begin{pmatrix} \hat{x}_f \\ \hat{x}'_f \end{pmatrix} = \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \begin{pmatrix} 1 - \cos \mu & \sin \mu \\ -\sin \mu & 1 - \cos \mu \end{pmatrix}^{-1} \begin{pmatrix} 0 \\ \hat{x}_f^2 \end{pmatrix} = \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \frac{1}{2 \sin \frac{\mu}{2}} \begin{pmatrix} -\cos \frac{\mu}{2} \\ \sin \frac{\mu}{2} \end{pmatrix} \hat{x}_f^2$$

$$\left. \begin{aligned} \hat{x}_f &= -\frac{4}{k_2 l_s} \beta^{-\frac{3}{2}} \tan \frac{\mu}{2} \\ \hat{x}'_f &= \frac{4}{k_2 l_s} \beta^{-\frac{3}{2}} \tan^2 \frac{\mu}{2} \end{aligned} \right\} \hat{x} = \hat{x}_f + \Delta \hat{x} \quad J_f = \frac{1}{2} (\hat{x}_f^2 + \hat{x}'_f^2) = \frac{1}{2 \beta^3} \left( \frac{4}{k_2 l_s} \frac{\tan \frac{\mu}{2}}{\cos \frac{\mu}{2}} \right)^2$$

$$\begin{pmatrix} \Delta \hat{x}_{n+1} \\ \Delta \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu & \sin \mu \\ -\sin \mu & \cos \mu \end{pmatrix} \left[ \begin{pmatrix} \Delta \hat{x}_n \\ \Delta \hat{x}'_n \end{pmatrix} - \begin{pmatrix} 0 \\ \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \Delta \hat{x}_n^2 - 4 \tan \frac{\mu}{2} \Delta \hat{x}_n \end{pmatrix} \right]$$



# The Dynamic Aperture



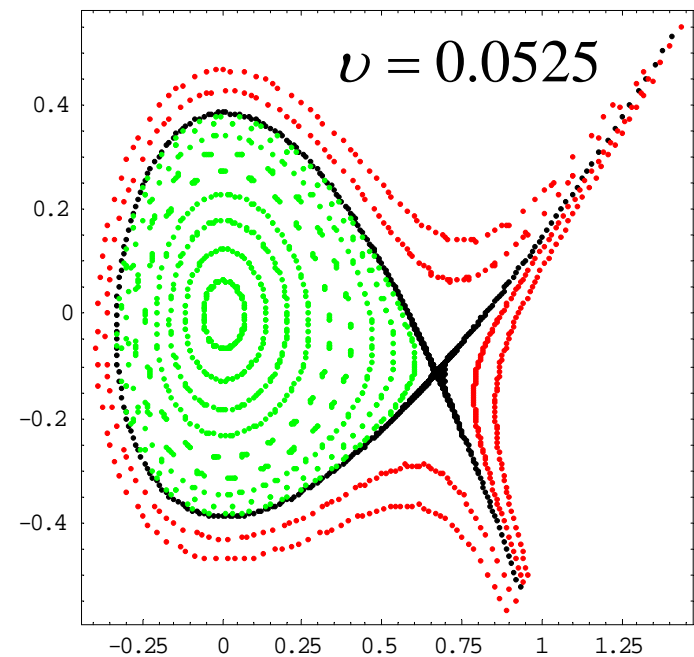
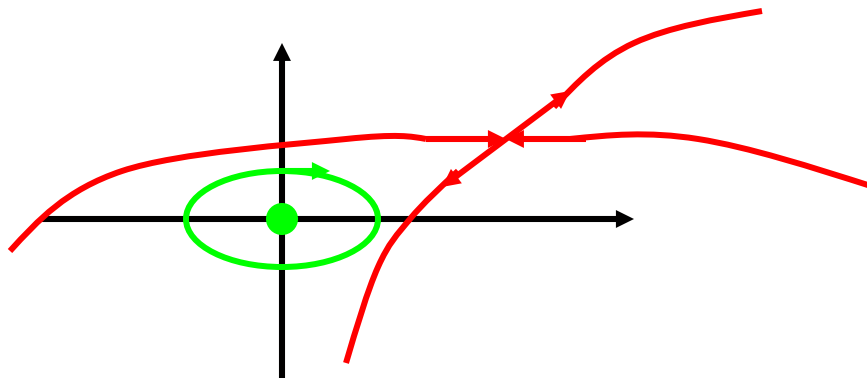
CHESS &amp; LEPP

$$\begin{pmatrix} \Delta \hat{x}_{n+1} \\ \Delta \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu & \sin \mu \\ -\sin \mu & \cos \mu \end{pmatrix} \left[ \begin{pmatrix} \Delta \hat{x}_n \\ \Delta \hat{x}'_n \end{pmatrix} - \begin{pmatrix} 0 \\ \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \Delta \hat{x}_n^2 - 4 \tan \frac{\mu}{2} \Delta \hat{x}_n \end{pmatrix} \right]$$

$$\begin{pmatrix} \Delta \hat{x}_{n+1} \\ \Delta \hat{x}'_{n+1} \end{pmatrix} = \begin{pmatrix} \cos \mu + 4 \sin \mu \tan \frac{\mu}{2} & \sin \mu \\ -\sin \mu + 4 \cos \mu \tan \frac{\mu}{2} & \cos \mu \end{pmatrix} \left[ \begin{pmatrix} \Delta \hat{x}_n \\ \Delta \hat{x}'_n \end{pmatrix} - \begin{pmatrix} 0 \\ \frac{k_2 l_s}{2} \beta^{\frac{3}{2}} \Delta \hat{x}_n^2 \end{pmatrix} \right]$$

$$\text{Tr}[\underline{M}] = 2 \frac{\cos \frac{\mu}{2} (1 + 2 \sin^2 \frac{\mu}{2})}{\cos \frac{\mu}{2}} \geq 2$$

The additional fixed point is unstable !





# Sextupole Aperture



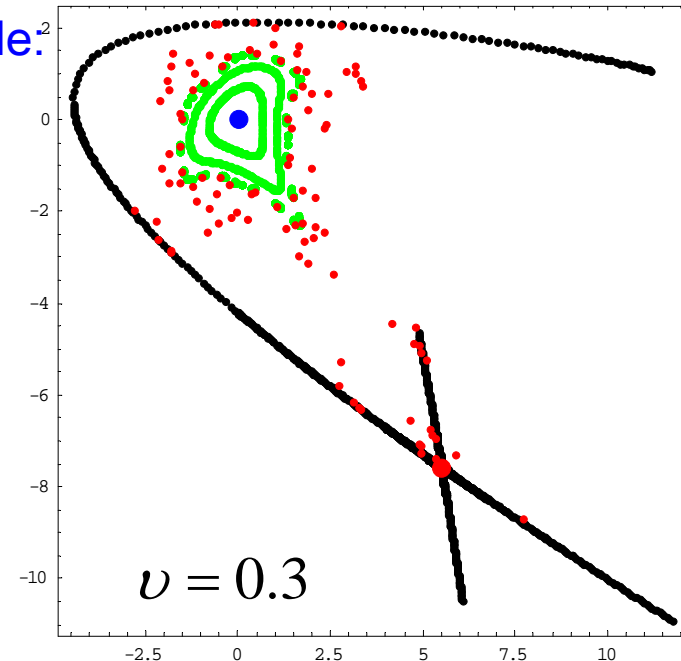
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If the chromaticity is corrected by a single sextupole:

$$\xi_x = \xi_{0x} + \frac{1}{4\pi} \beta_x \eta_x k_2 l \approx 0$$

$$J_f = \frac{1}{2\beta^3} \left( \frac{4}{k_2 l_s} \frac{\tan \frac{\mu}{2}}{\cos \frac{\mu}{2}} \right)^2 \approx \frac{1}{2\beta} \left( \frac{\eta}{\xi_{0x} \pi} \frac{\sin \frac{\mu}{2}}{\cos^2 \frac{\mu}{2}} \right)^2$$

Often the dynamic aperture is much smaller than the fixed point indicates !



When many sextupoles are used:

$$\xi_{0x} + \frac{N}{4\pi} \beta_x \eta_x k_2 l \approx 0$$

The sum of all  $k_2^2$  is then reduced to about  $\sum (k_2 l \beta)^2 \approx N (k_2 l \beta)^2 \approx \frac{1}{N} \left( \frac{4\pi}{\eta} \xi_{0x} \right)^2$

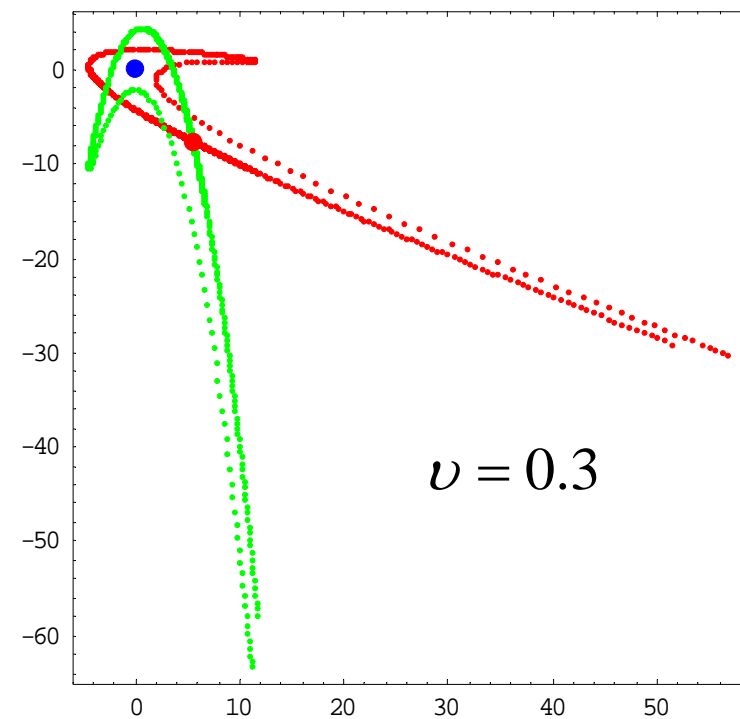
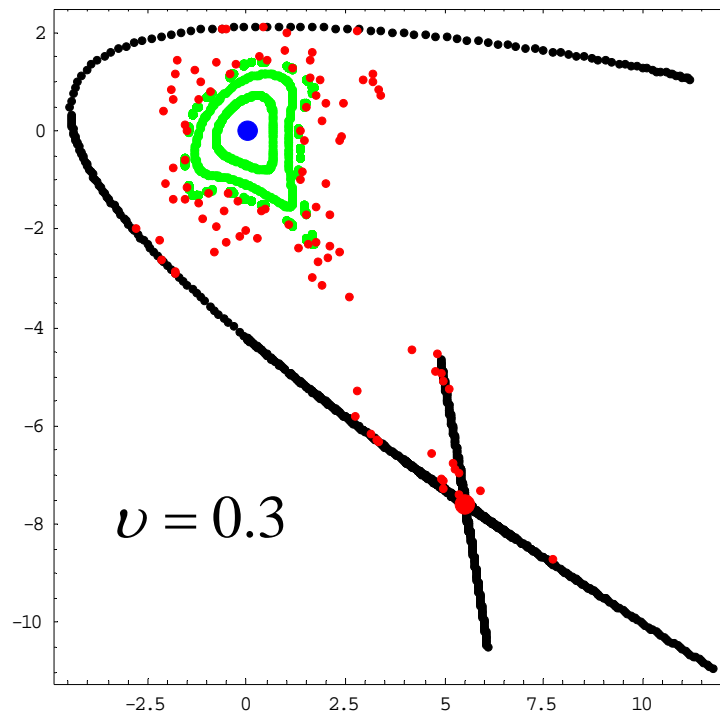
The dynamic aperture is therefore greatly increased when distributed sextupoles are used.



# Sextupole Extraction



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Due to the narrow region of unstable trajectories, sextupoles are used for slow particle extraction at a tune of  $1/3$ .

The intersection of **stable** and **unstable** manifolds is a certain indication of chaos.

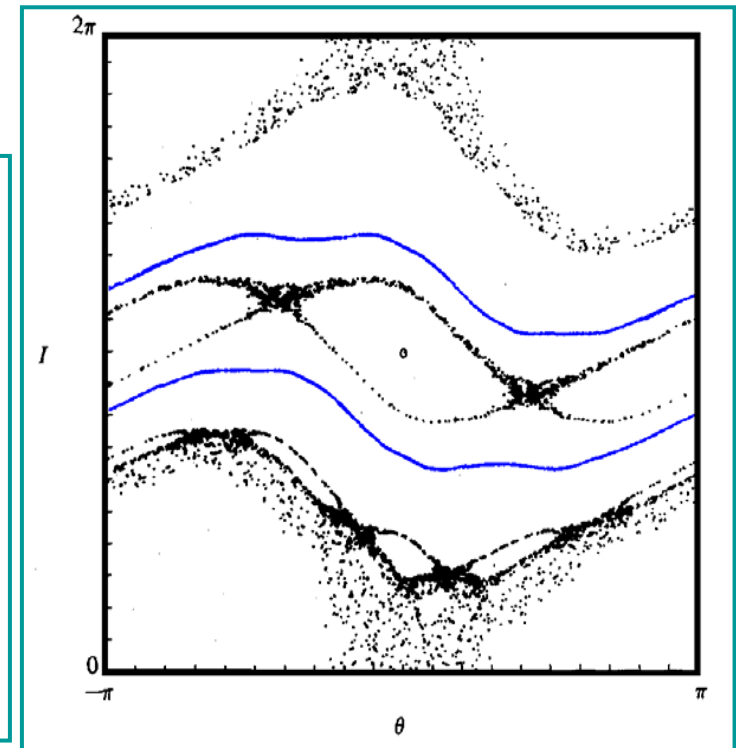
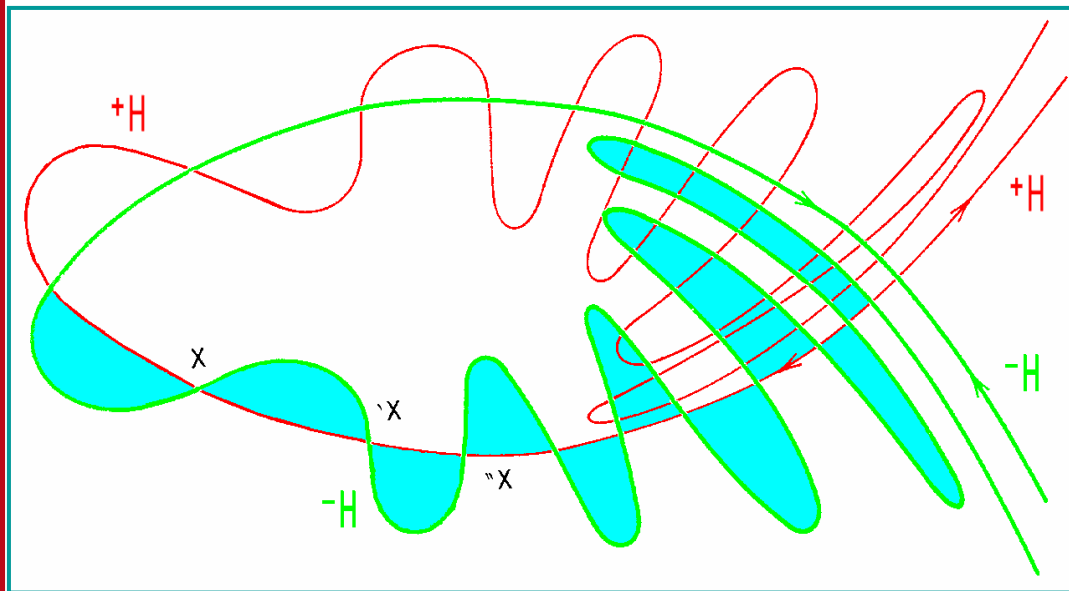


# Homoclinic Points



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- At instable fixed points, there is a **stable** and an **instabile invariant curve**.
- Intersections of these curves (**homoclinic points**) lead to **chaos**.



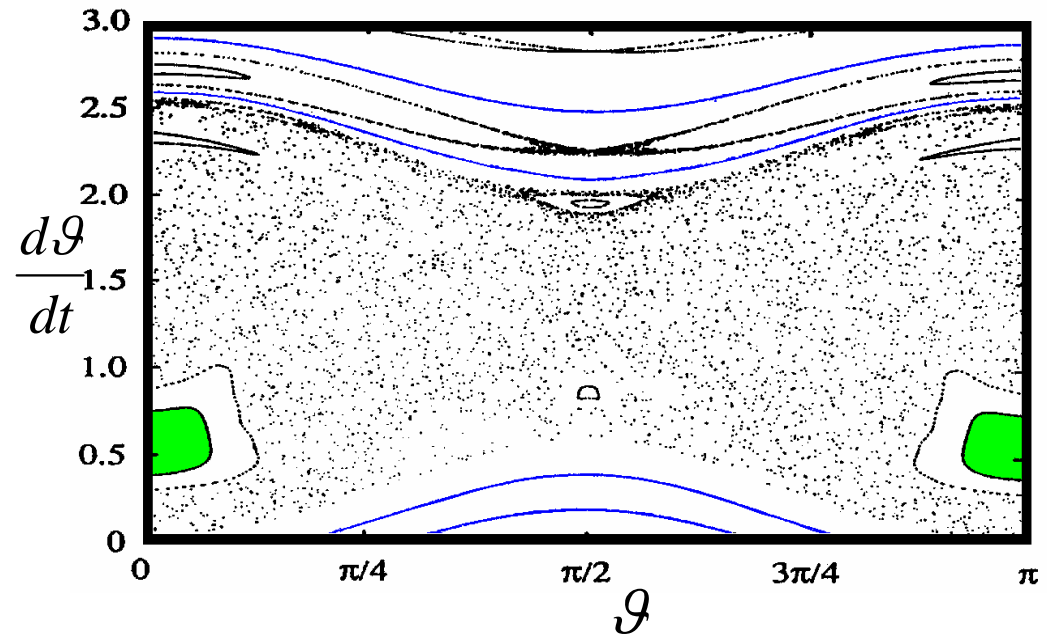
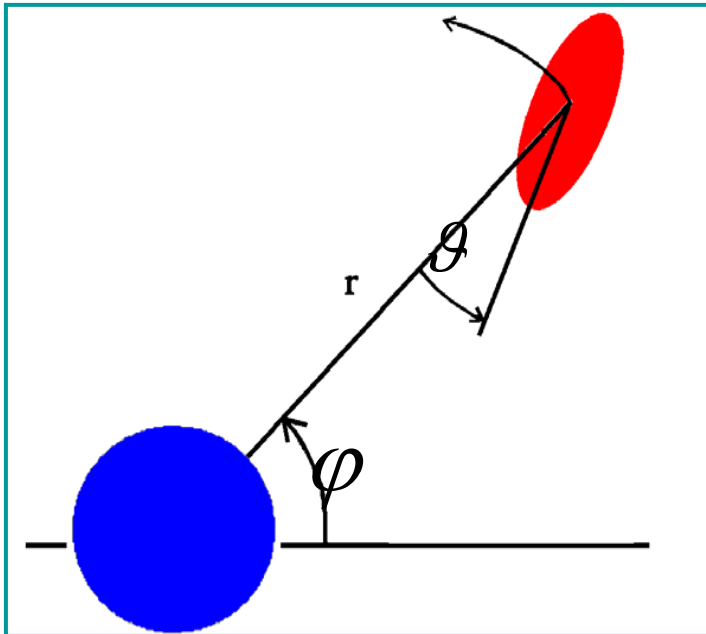


# Hyperion: rotation around the vertical



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$$\frac{d^2(\vartheta + \varphi(t))}{dt^2} = -\alpha \left(\frac{a}{r(t)}\right)^3 \sin 2\vartheta$$



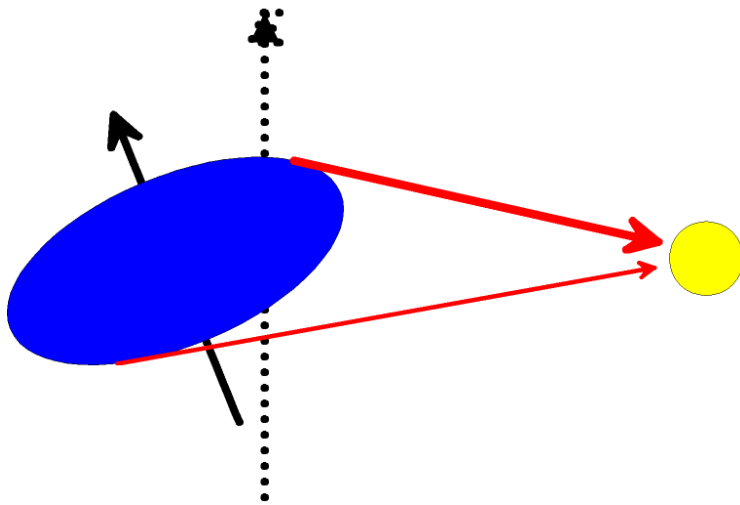
- On the path from Rotation to Libration around the Spin-Orbit-Coupling is a strong chaotic region.



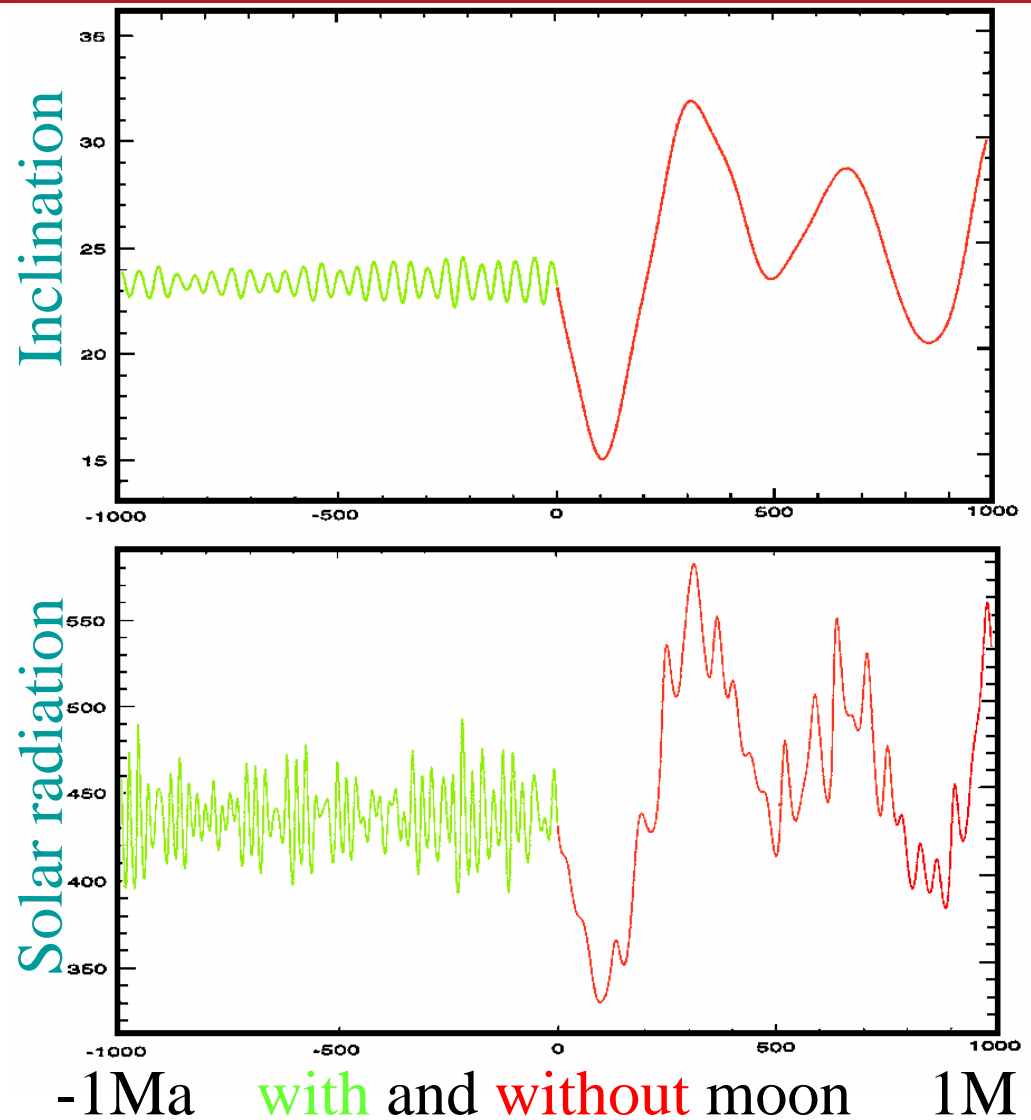
# Tilt of the earth



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- Tidal forces from moon and sun cause a stabilization of the rotation axis.





$$\begin{pmatrix} x' \\ a' \end{pmatrix} = \begin{pmatrix} 0 & 1 \\ -K & 0 \end{pmatrix} \begin{pmatrix} x \\ a \end{pmatrix} + \begin{pmatrix} 0 \\ \Delta f \end{pmatrix}$$

$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix} = \sqrt{2J} \underline{\beta} \vec{S}$$

This would be a solution with constant  $J$  and  $\phi$  when  $\Delta f=0$ .

Variation of constants:

$$\frac{J'}{\sqrt{2J}} \underline{\beta} \vec{S} + \sqrt{2J} \phi_0' \begin{pmatrix} 0 & \sqrt{\beta} \\ -\frac{1}{\sqrt{\beta}} & -\frac{\alpha}{\sqrt{\beta}} \end{pmatrix} \vec{S} = \begin{pmatrix} 0 \\ \Delta f \end{pmatrix}$$

$$\frac{J'}{\sqrt{2J}} \vec{S} + \sqrt{2J} \phi_0' \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \vec{S} = \underline{\beta}^{-1} \begin{pmatrix} 0 \\ \Delta f \end{pmatrix} \quad \text{with} \quad \underline{\beta}^{-1} = \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix}$$

$$\frac{J'}{\sqrt{2J}} = \cos(\psi + \phi_0) \sqrt{\beta} \Delta f \quad , \quad \sqrt{2J} \phi_0' = -\sin(\psi + \phi_0) \sqrt{\beta} \Delta f$$



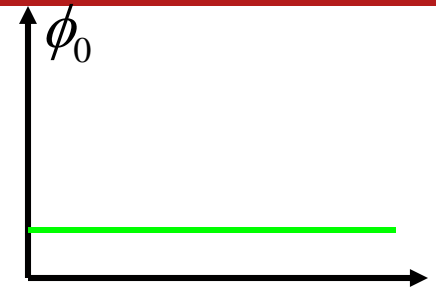


# Simplification of linear motion

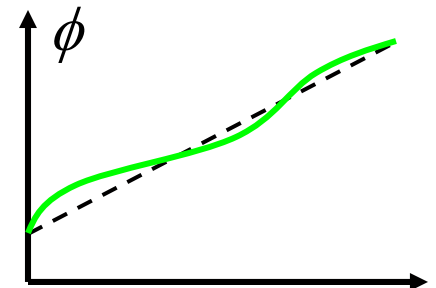


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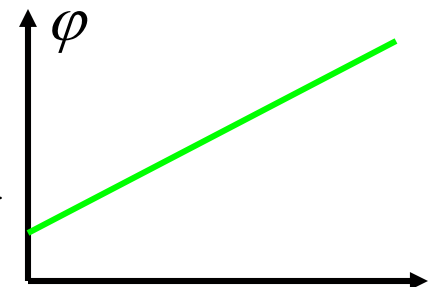
$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi + \phi_0) \\ \cos(\psi + \phi_0) \end{pmatrix} \Rightarrow \begin{aligned} J' &= 0 \\ \phi_0' &= 0 \end{aligned}$$



$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin \phi \\ \cos \phi \end{pmatrix} \Rightarrow \begin{aligned} J' &= 0 \\ \phi' &= \frac{1}{\beta} \end{aligned}$$



$$\begin{pmatrix} x \\ a \end{pmatrix} = \sqrt{2J} \begin{pmatrix} \sqrt{\beta} & 0 \\ -\frac{\alpha}{\sqrt{\beta}} & \frac{1}{\sqrt{\beta}} \end{pmatrix} \begin{pmatrix} \sin(\psi - \mu \frac{s}{L} + \varphi) \\ \cos(\psi - \mu \frac{s}{L} + \varphi) \end{pmatrix} \Rightarrow \begin{aligned} J' &= 0 \\ \varphi' &= \mu \frac{1}{L} \end{aligned}$$



$$\tilde{\psi} = \psi - \mu \frac{s}{L} \Rightarrow \tilde{\psi}(s + L) = \tilde{\psi}(s)$$

Corresponds to Floquet's Theorem



## Quasi-periodic Perturbation



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$$J' = \cos(\psi + \phi) \sqrt{2J\beta} \Delta f \quad , \quad \phi' = -\sin(\psi + \phi) \sqrt{\frac{\beta}{2J}} \Delta f$$

$$J' = \cos(\tilde{\psi} + \phi) \sqrt{2J\beta} \Delta f \quad , \quad \phi' = \mu \frac{1}{L} - \sin(\tilde{\psi} + \phi) \sqrt{\frac{\beta}{2J}} \Delta f$$

New independent variable  $\mathcal{G} = 2\pi \frac{s}{L}$

$$\frac{d}{d\mathcal{G}} J = \cos(\tilde{\psi} + \phi) \sqrt{2J\beta} \Delta f \frac{L}{2\pi} \quad , \quad \frac{d}{d\mathcal{G}} \phi = \nu - \sin(\tilde{\psi} + \phi) \sqrt{\frac{\beta}{2J}} \Delta f \frac{L}{2\pi}$$

$$\Delta f(x) = \Delta f(\sqrt{2J\beta} \sin(\tilde{\psi} + \phi))$$

The perturbations are  $2\pi$  periodic in  $\mathcal{G}$  and in  $\phi$

$\phi$  is approximately  $\phi \approx \nu \cdot \mathcal{G}$

For irrational  $\nu$ , the perturbations are **quasi-periodic**.



# Tune Shift with Amplitude



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$$\frac{d}{d\mathcal{G}} J = \cos(\tilde{\psi} + \varphi) \sqrt{2J\beta} \Delta f \frac{L}{2\pi} \quad , \quad \frac{d}{d\mathcal{G}} \varphi = \nu - \sin(\tilde{\psi} + \varphi) \sqrt{\frac{\beta}{2J}} \Delta f \frac{L}{2\pi}$$

$$\frac{d}{d\mathcal{G}} \varphi = \partial_J H \quad , \quad \frac{d}{d\mathcal{G}} J = -\partial_\varphi H \quad , \quad H(\varphi, J, \mathcal{G}) = \nu \cdot J - \frac{L}{2\pi} \int_0^x \Delta f(\hat{x}, s) d\hat{x}$$

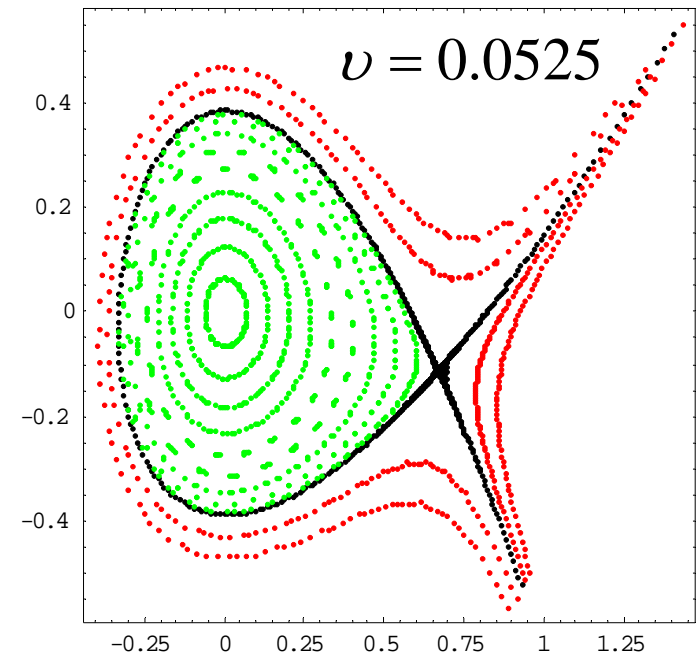
The motion remains Hamiltonian in the perturbed coordinates !

If there is a part in  $\partial_J H$  that does not depend on  $\varphi, s \Rightarrow$  **Tune shift**

The effect of other terms tends to average out.

$$\varphi(\mathcal{G}) - \varphi_0 \approx \mathcal{G} \cdot \partial_J \langle H \rangle_{\varphi, \mathcal{G}}(J)$$

$$\nu(J) = \nu + \partial_J \langle \Delta H \rangle_{\varphi, \mathcal{G}}(J)$$





# Tune Shift Examples



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$$H(\varphi, J) = \nu \cdot J - \frac{L}{2\pi} \int_0^x \Delta f(\hat{x}, s) d\hat{x} \quad , \quad \Delta \nu(J) = \partial_J \langle \Delta H \rangle_{\varphi, \vartheta}$$

Quadrupole:  $\Delta f = -\Delta k x$

$$\Delta H = \frac{L}{2\pi} \Delta k \frac{1}{2} x^2 = \frac{L}{2\pi} \Delta k J \beta \sin^2(\tilde{\psi} + \varphi)$$

$$\langle \Delta H \rangle_{\varphi, \vartheta} = \frac{1}{2\pi} \int_0^{2\pi} \Delta k \beta d\vartheta L \frac{J}{4\pi} = \int_0^L \Delta k \beta ds \frac{J}{4\pi} \Rightarrow \Delta \nu = \frac{1}{4\pi} \oint \Delta k \beta ds$$

Sextupole:  $\Delta f = -k_2 \frac{1}{2} x^2$

$$\Delta H = \frac{L}{2\pi} k_2 \frac{1}{3!} x^3 = \frac{L}{2\pi} k_2 \frac{1}{3!} \sqrt{2J\beta}^3 \sin^3(\tilde{\psi} + \varphi)$$

$$\langle \Delta H \rangle_{\varphi, \vartheta} = 0 \Rightarrow \Delta \nu = 0$$

Octupole:  $\Delta f = -k_3 \frac{1}{3!} x^3$

$$\Delta H = \frac{L}{2\pi} k_3 \frac{1}{4!} x^4 = \frac{L}{2\pi} k_3 \frac{1}{3!} (J\beta)^2 \sin^4(\tilde{\psi} + \varphi)$$

$$\langle \Delta H \rangle_{\varphi, \vartheta} = \frac{J^2}{3!2\pi} \oint k_3 \beta^2 ds \left\langle \frac{1}{2^4} (e^{i\varphi} - e^{-i\varphi})^4 \right\rangle_{\varphi} \Rightarrow \Delta \nu = J \frac{1}{16\pi} \oint k_3 \beta^2 ds$$



$$\frac{d}{d\vartheta} J = \cos(\tilde{\psi} + \varphi) \sqrt{2J\beta} \Delta f \frac{L}{2\pi} \quad , \quad \frac{d}{d\vartheta} \varphi = \nu - \sin(\tilde{\psi} + \varphi) \sqrt{\frac{\beta}{2J}} \Delta f \frac{L}{2\pi}$$

$$\frac{d}{d\vartheta} \varphi = \partial_J H \quad , \quad \frac{d}{d\vartheta} J = -\partial_\varphi H \quad , \quad H(\varphi, J, \vartheta) = \nu \cdot J - \frac{L}{2\pi} \int_0^x \Delta f(\hat{x}, s) d\hat{x}$$

The effect of the perturbation is especially strong when

$$\cos(\tilde{\psi} + \varphi) \sqrt{\beta} \Delta f \quad \text{or} \quad \sin(\tilde{\psi} + \varphi) \sqrt{\beta} \Delta f$$

has contributions that hardly change, i.e. the change of

$$\sqrt{\beta(\vartheta)} \Delta f(x(\vartheta), \vartheta) \quad \text{is in resonance with the rotation angle} \quad \varphi(\vartheta) \quad .$$

Periodicity allows Fourier expansion:

$$H(\varphi, J, \vartheta) = \sum_{n,m=-\infty}^{\infty} \hat{H}_{nm}(J) e^{i[n\vartheta+m\varphi]} = \sum_{n,m=-\infty}^{\infty} H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))$$

$$H_{00}(J) = \langle H(\varphi, J, s) \rangle_{\varphi, s} \Rightarrow \text{Tune shift}$$



# The Single Resonance Model



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$$\frac{d}{d\mathcal{G}} J = \sum_{n,m=-\infty}^{\infty} m H_{nm}(J) \sin(n\mathcal{G} + m\varphi + \Psi_{nm}(J))$$

$$\frac{d}{d\mathcal{G}} \varphi = \nu + \partial_J \sum_{n,m=-\infty}^{\infty} H_{nm}(J) \cos(n\mathcal{G} + m\varphi + \Psi_{nm}(J))$$

Strong deviation from:  $J = J_0$ ,  $\varphi = \nu \mathcal{G} + \varphi_0$

Occur when there is coherence between the

perturbation and the phase space rotation:  $n + m \frac{d}{ds} \varphi \approx 0$

Resonance condition: tune is rational  $n + m \nu = 0$

On resonance the integral would increase indefinitely !

Neglecting all but the most important term

$$H(\varphi, J, \mathcal{G}) \approx \nu J + H_{00}(J) + H_{nm}(J) \cos(n\mathcal{G} + m\varphi + \Psi_{nm}(J))$$



## Fixed points



$$\frac{d}{d\vartheta} J = mH_{nm}(J) \sin(n\vartheta + m\varphi + \Psi_{nm}(J))$$

$$\frac{d}{d\vartheta} \varphi = \nu + \Delta\nu(J) + \partial_J [H_{nm}(J) \cos(n\vartheta + m\varphi + \Psi_{nm}(J))]$$

$$\Phi = \frac{1}{m} [n\vartheta + m\varphi + \Psi_{nm}(J)] , \quad \delta = \nu + \frac{n}{m}$$

$$\frac{d}{d\vartheta} J = mH_{nm}(J) \sin(m\Phi) , \quad \frac{d}{d\vartheta} \Phi = \delta + \Delta\nu(J) + H'_{nm}(J) \cos(m\Phi)$$

$$H(\varphi, J, \vartheta) \approx \delta J + H_{00}(J) + H_{nm}(J) \cos(m\Phi)$$

Fixed points:  $\frac{d}{d\vartheta} J = mH_{nm}(J_f) \sin(m\Phi_f) = 0 \Rightarrow \Phi_f = \frac{k}{m} \pi$

If  $\delta + \Delta\nu(J_f) \pm H'_{nm}(J_f) = 0$  has a solution.

$$\frac{d}{d\vartheta} \Delta J = \pm m^2 H_{nm}(J_f) \Delta\Phi , \quad \frac{d}{d\vartheta} \Delta\Phi = [\Delta\nu'(J_f) \pm H''_{nm}(J_f)] \Delta J$$

Stable fixed point for:  $H_{nm}(J_f) [H''_{nm}(J_f) \pm \Delta\nu'(J_f)] < 0$



# Third Integer Resonances



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Sextupole:  $\Delta f = -k_2 \frac{1}{2} x^2$

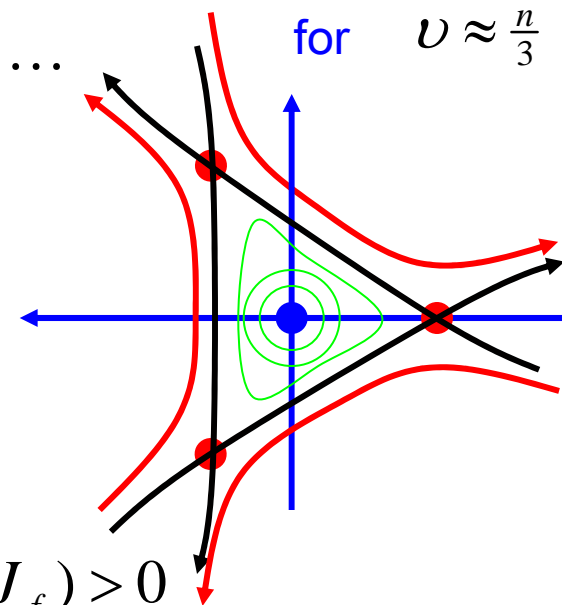
$$\begin{aligned} \Delta H &= \frac{L}{2\pi} k_2 \frac{1}{3!} x^3 = \frac{L}{2\pi} k_2 \frac{1}{3!} \sqrt{2J\beta}^3 \sin^3(\tilde{\psi} + \varphi) \\ &= \frac{L}{2\pi} k_2 \frac{1}{3!4} \sqrt{2J\beta}^3 [\sin(3[\tilde{\psi} + \varphi]) + 3\sin(\tilde{\psi} + \varphi)] \end{aligned}$$

Simplification: one sextupole  $k_2(\mathcal{G}) = k_2 \delta(\mathcal{G}) = k_2 \frac{1}{2\pi} \sum_{n=-\infty}^{\infty} \cos(n\mathcal{G})$

$$\Delta H = \frac{L}{2\pi} k_2 \frac{1}{3!4} \sqrt{2J\beta}^3 \frac{1}{2\pi} \cos(-n\mathcal{G} + 3\varphi + \tilde{\psi} - \frac{\pi}{2}) + \dots$$

$$\Delta H \approx A_2 \sqrt{J}^3 \cos(3\Phi)$$

$$\left. \begin{aligned} \Phi_f = 0, \frac{1}{3}\pi, \frac{2}{3}\pi, \dots \\ \delta \pm A_2 \frac{3}{2} \sqrt{J} = 0 \end{aligned} \right\} \begin{aligned} \Phi_f = \frac{1}{3}\pi, \pi, \frac{5}{3}\pi \\ \text{for } \delta > 0 \end{aligned}$$



All these fixed points are instable since  $H_{nm}(J_f) H''_{nm}(J_f) > 0$





# Fourth Integer Resonances



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Octupole:  $\Delta f = -k_3 \frac{1}{3!} x^3$  ,  $\Delta H = \frac{L}{2\pi} k_3 \frac{1}{4!} x^4 = \frac{L}{2\pi} k_3 \frac{1}{3!} J^2 \beta^2 \sin^4(\tilde{\psi} + \varphi)$

$$= \frac{L}{2\pi} k_3 \frac{1}{3!} J^2 \beta^2 [\cos(4[\tilde{\psi} + \varphi]) - 4 \cos(\tilde{\psi} + \varphi) + 3]$$

Simplification: one octupole  $k_3(\mathcal{G}) = k_3 \delta(\mathcal{G}) = k_3 \frac{1}{2\pi} \sum_{n=-\infty}^{\infty} \cos(n\mathcal{G})$

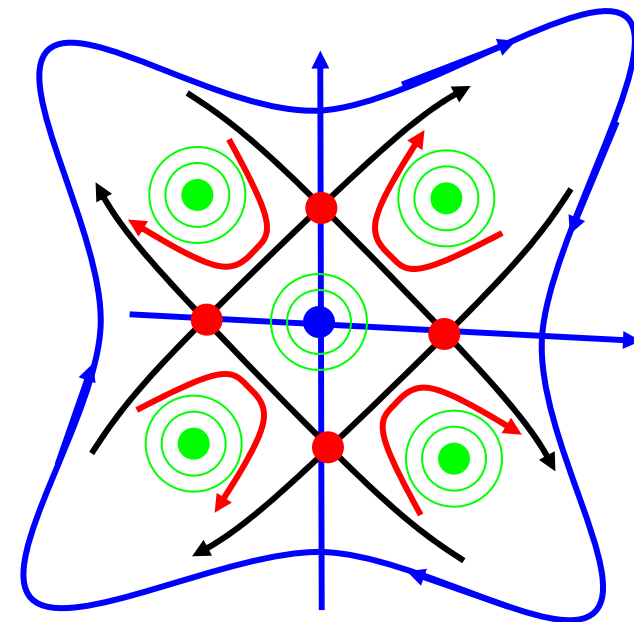
$$\Delta H \approx A_3 J^2 [3 + \cos(4\Phi)] \quad \text{for } \nu \approx \frac{n}{4}$$

$$\Phi_f = 0, \frac{1}{4}\pi, \frac{2}{4}\pi, \dots \quad \text{Either 8 fixed points: } \delta < 0$$

$$\delta + A_3 2J (3 \pm 1) = 0 \quad \text{or none for: } \delta > 0$$

$$H_{nm}(J_f) [H_{nm}''(J_f) \pm \Delta \nu'(J_f)] < 0$$

Stability for  $(2A_3 J)^2 [1 \pm 3] < 0$ ,  
i.e. for the 4 outer fixed points.





## Resonance Width (Strength)

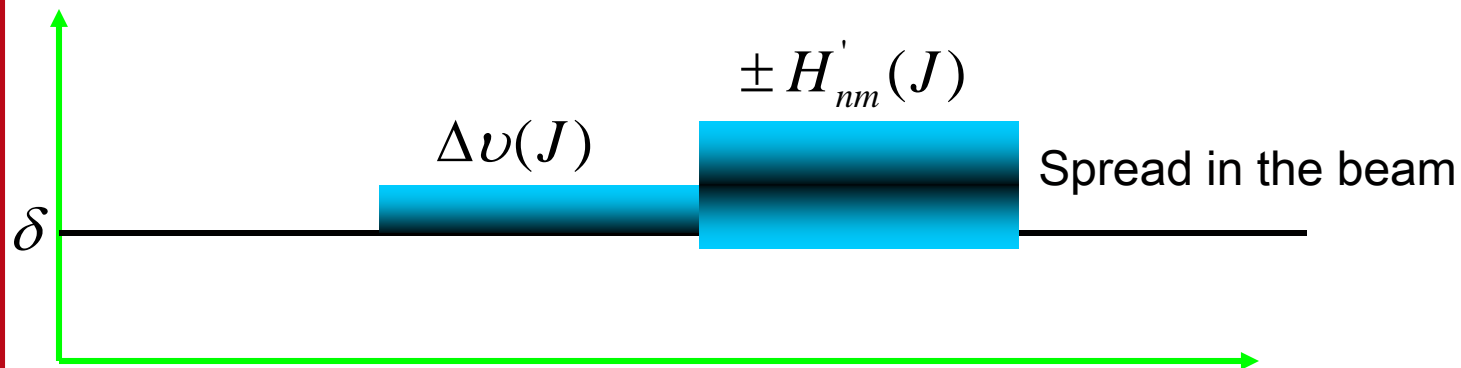


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Fixed points:  $\frac{d}{d\theta} J = mH_{nm}(J_f)\sin(m\Phi_f) = 0 \Rightarrow \Phi_f = \frac{k}{m}\pi$

If  $\delta + \Delta\nu(J_f) \pm H'_{nm}(J_f) = 0$  has a solution.

$\delta$  has to avoid the region  $\delta + \Delta\nu(J) \pm H'_{nm}(J) = 0$  for all particles.



Assuming that the tune shift and perturbation are monotonous in J:

This tune region has the width  $\Delta_{nm} = 2 |H'_{nm}(J_{\max})|$  for strong resonances.

$\Delta_{nm}$  Is called **Resonance Width**, Resonance Strength, or Stop-Band Width



$$\frac{d}{d\vartheta} J_x = \cos(\tilde{\psi}_x + \varphi_x) \sqrt{2J_x \beta_x} \Delta f_x \frac{L}{2\pi} \quad , \quad \frac{d}{d\vartheta} \varphi_x = \nu_x - \sin(\tilde{\psi}_x + \varphi_x) \sqrt{\frac{\beta_x}{2J_x}} \Delta f_x \frac{L}{2\pi}$$

$$\frac{d}{d\vartheta} J_y = \cos(\tilde{\psi}_y + \varphi_y) \sqrt{2J_y \beta_y} \Delta f_y \frac{L}{2\pi} \quad , \quad \frac{d}{d\vartheta} \varphi_y = \nu_y - \sin(\tilde{\psi}_y + \varphi_y) \sqrt{\frac{\beta_y}{2J_y}} \Delta f_y \frac{L}{2\pi}$$

$$\frac{d}{d\vartheta} \vec{\varphi} = \vec{\partial}_J H \quad , \quad \frac{d}{d\vartheta} \vec{J} = -\vec{\partial}_\varphi H \quad , \quad H(\vec{\varphi}, \vec{J}, \vartheta) = \vec{\nu} \cdot \vec{J} - \frac{L}{2\pi} \int_0^{\vec{x}} \Delta \vec{f}(\hat{x}, s) d\hat{x}$$

The integral form can be chosen since it is path independent. This is due to the Hamiltonian nature of the force:

$$\Delta f_{x,y}(x, y, s) = -\partial_{x,y} \Delta H(x, y, s)$$

Single Resonance model for two dimensions means retaining only the amplitude dependent tune shift and one term in the two dimensional Fourier expansion:

$$H(\vec{\varphi}, \vec{J}, \vartheta) = \vec{\nu} \cdot \vec{J} + H_{00}(\vec{J}) + H_{n\vec{m}}(\vec{J}) \cos(n\vartheta + m_x \varphi_x + m_y \varphi_y + \Psi_{n\vec{m}}(\vec{J}))$$

For  $n + m_x \nu_x + m_y \nu_y \approx 0$

$$m_x \varphi_x + m_y \varphi_y = \vec{m} \cdot \vec{\varphi}$$



$n + m_x \nu_x + m_y \nu_y \approx 0$  means that oscillations in y can drive oscillations in x in

$$x'' = -K x + \Delta f_x(x, y, s)$$

The resonance term in the Hamiltonian then changes only slowly:

$$H(\vec{\varphi}, \vec{J}, \vartheta) = \vec{\nu} \cdot \vec{J} + H_{00}(\vec{J}) + H_{n\vec{m}}(\vec{J}) \cos(n\vartheta + m_x \varphi_x + m_y \varphi_y + \Psi_{n\vec{m}}(\vec{J}))$$

$$\frac{d}{d\vartheta} \vec{\varphi} = \vec{\partial}_{\vec{J}} H \quad , \quad \frac{d}{d\vartheta} \vec{J} = -\vec{\partial}_{\varphi} H$$

$$J = \vec{m} \cdot \vec{J}$$

$$J_{\perp} = m_x J_x - m_y J_y = \vec{m} \times \vec{J} \quad \Rightarrow \quad \frac{d}{d\vartheta} J_{\perp} = 0$$

Difference resonances lead to stable motion since:

$$n + |m_x| \nu_x - |m_y| \nu_y \approx 0 \Rightarrow |m_x| J_x + |m_y| J_y = \text{const.}$$

Sum resonances lead to unstable motion since:

$$n + |m_x| \nu_x + |m_y| \nu_y \approx 0 \Rightarrow |m_x| J_x - |m_y| J_y = \text{const.}$$



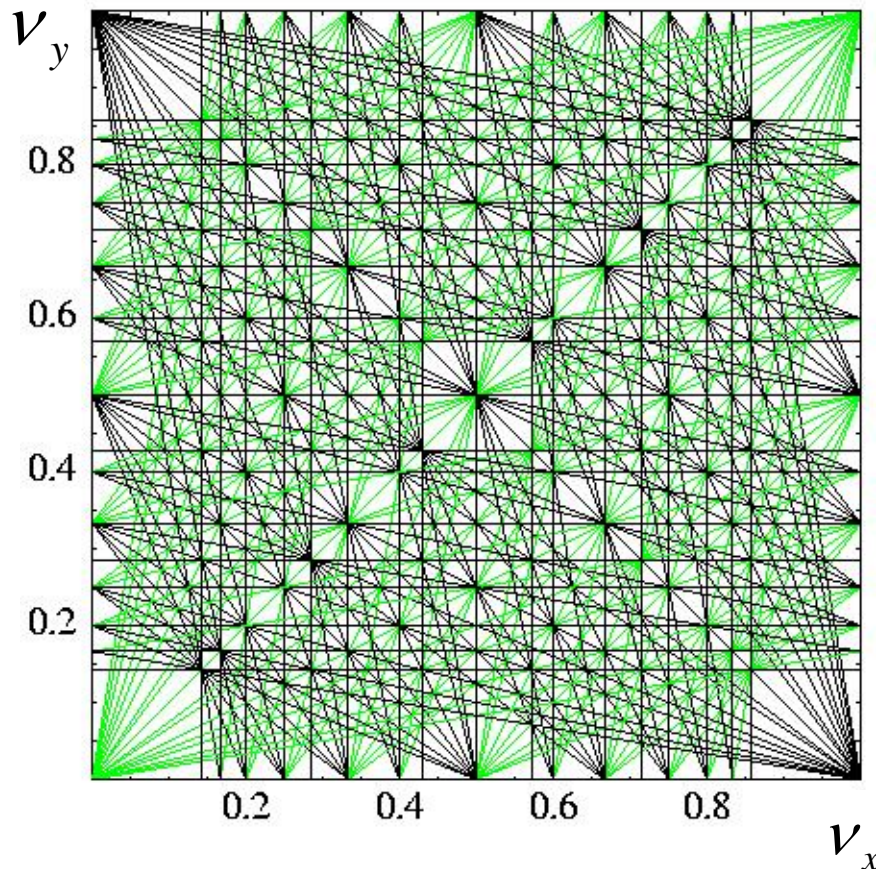
# Resonances Diagram



CHESS & LEPP

$n + m_x \nu_x + m_y \nu_y \approx 0$  means that oscillations in  $y$  can drive oscillations in  $x$  in

$$x'' = -K x + \Delta f_x(x, y, s)$$



All these resonances have to be avoided by their respective resonance width.

The position of an accelerator in the tune plane is called its Working Point.