Physics 443, Solutions to PS 3^1

1. Griffiths 3.23. It is easiest to first write the hamiltonian matrix. By inspection

$$H = \epsilon \begin{pmatrix} 1 & 1 \\ 1 & -1 \end{pmatrix}$$

We find the eigenvalues λ by setting

$$\det(H - \lambda I) = 0$$

Then $\lambda_{\pm} = \pm \epsilon \sqrt{2}$. Let

$$\vec{v} = \begin{pmatrix} a \\ b \end{pmatrix}$$

be an eigenvector. Then

$$\begin{aligned} H \vec{v}_{\pm} &= \lambda_{\pm} \vec{v}_{\pm} \\ &= \epsilon \begin{pmatrix} 1 & 1 \\ 1 & -1 \end{pmatrix} \begin{pmatrix} a \\ b \end{pmatrix} \\ &= \lambda_{\pm} \begin{pmatrix} a \\ b \end{pmatrix} \\ &\Rightarrow \vec{v}_{\pm} = \begin{pmatrix} 1 \\ -1 \pm \sqrt{2} \end{pmatrix} \end{aligned}$$

Or in the $|\rangle$ representation

$$|v_{\pm}\rangle = |1\rangle + (-1 \pm \sqrt{2})|2\rangle$$

2. Griffiths 3.24.

Since the set of orthonormal vectors $|e_n\rangle$ is complete, any state can be written as a linear combination of those vectors. In paricular, the state $|\alpha\rangle$ can be written as

$$|\alpha\rangle = \sum_{m} a_{m} |e_{m}\rangle \tag{1}$$

Then

$$\langle e_n \mid \alpha \rangle = \sum_m a_m \langle e_n \mid e_m \rangle = \sum_m a_m \delta_{mn} = a_n$$
 (2)

where we have used the orthhonormality of the eigenvectors. Finally substitute $\langle e_n | \alpha \rangle = a_n$ from Equation 3 into Equation 2.

$$|\alpha\rangle = \sum_{n} |e_{n}\rangle\langle e_{n} |\alpha\rangle \tag{3}$$

¹Courtesy Shaffique Adam

 So

$$\hat{Q}|\alpha\rangle = \sum_{n} \hat{Q}|e_{n}\rangle\langle e_{n} |\alpha\rangle$$

$$= \sum_{n} q_{n}|e_{n}\rangle\langle e_{n} |\alpha\rangle$$

$$= \left(\sum_{n} q_{n}|e_{n}\rangle\langle e_{n} |\right)|\alpha\rangle$$

3. Griffiths 3.3. We are given that $\langle h | \hat{Q} | h \rangle = \langle \hat{Q}h | h \rangle$ for all states $| h \rangle$. If we define $| h \rangle = | f \rangle + | g \rangle$ then

By hypothesis $\langle f \mid \hat{Q} \mid f \rangle = \langle \hat{Q}f \mid f \rangle$ and similarly for $\mid g \rangle$ so Equation 3 reduces to

$$\left\langle f \mid \hat{Q} \mid g \right\rangle + \left\langle g \mid \hat{Q} \mid f \right\rangle = \left\langle \hat{Q}f \mid g \right\rangle + \left\langle \hat{Q}g \mid f \right\rangle \tag{5}$$

Alternatively, if we let $\mid h \rangle = \mid f \rangle + i \mid g \rangle$ we find that

$$i\left\langle f \mid \hat{Q} \mid g\right\rangle - i\left\langle g \mid \hat{Q} \mid f\right\rangle = i\left\langle \hat{Q}f \mid g\right\rangle - i\left\langle \hat{Q}g \mid f\right\rangle \tag{6}$$

or

$$\left\langle f \mid \hat{Q} \mid g \right\rangle - \left\langle g \mid \hat{Q} \mid f \right\rangle = \left\langle \hat{Q}f \mid g \right\rangle - \left\langle \hat{Q}g \mid f \right\rangle \tag{7}$$

The sum of equations (4) and (6) gives $\langle f \mid \hat{Q} \mid g \rangle = \langle \hat{Q}g \mid f \rangle$ and the difference gives $\langle g \mid \hat{Q} \mid f \rangle = \langle \hat{Q}g \mid f \rangle$

4. Griffiths 3.31. We have:

$$\frac{d}{dt} \langle xp \rangle = \frac{i}{\hbar} \langle [\frac{p^2}{2m} + V(x), xp] \rangle,$$

$$= \frac{i}{\hbar} \langle [\frac{p^2}{2m}, xp] + [V(x), xp] + \rangle,$$

$$= \frac{i}{\hbar} \langle [\frac{pp}{2m}, x]p + x[V(x), p] + \rangle,$$

$$= \frac{i}{\hbar} \langle -2i\hbar \frac{p^2}{2m} + x(\hbar i \frac{\partial V}{\partial x}) \rangle,$$

$$= 2 \langle T \rangle - \langle x \frac{\partial V}{\partial x} \rangle.$$
(8)

For a stationary state, we see that $2\langle T \rangle = \langle x \partial_x V \rangle$, which is the **Virial Theorem**. For the Harmonic Oscillator $V(x) = m\omega^2 x^2/2$, using the Virial Theorem, we see that $\langle T \rangle = \langle m\omega^2 x^2/2 \rangle = \langle V(x) \rangle$.

5. Griffiths 3.33. We begin by using that

$$\begin{aligned} a_{\pm} &= \frac{1}{\sqrt{2m}} (P \pm im\omega x), \\ P &= \sqrt{\frac{m}{2}} (a_{+} + a_{-}), \\ x &= \frac{-i}{\sqrt{2m\omega}} (a_{+} - a_{-}), \\ a_{+}|n-1\rangle &= i\sqrt{n\hbar\omega}|n\rangle, \\ a_{-}|n\rangle &= -i\sqrt{n\hbar\omega}|n-1\rangle, \\ \langle n|x|n'\rangle &= \frac{-i}{\sqrt{2m\omega}} (\langle n|a_{+} - a_{-}|n'\rangle), \\ &= \frac{-i}{\sqrt{2m\omega}} (\delta_{n',n-1}i\sqrt{n\hbar\omega} - \delta_{n,n'-1}(-i\sqrt{n'\hbar\omega})), \\ &= \sqrt{\frac{\hbar}{2m\omega}} (\delta_{n',n-1}\sqrt{n} + \delta_{n,n'-1}\sqrt{n'}), \\ \langle n|p|n'\rangle &= \sqrt{\frac{m}{2}} (\langle n|a_{+} + a_{-}|n'\rangle), \\ &= i\sqrt{\frac{\hbar m\omega}{2}} (\delta_{n',n-1}\sqrt{n} - \delta_{n,n'-1}\sqrt{n'}). \end{aligned}$$

We can then write these out in matrix notation as

$$x = \sqrt{\frac{\hbar}{2m\omega}} \begin{pmatrix} 0 & \sqrt{2} & 0 & 0 & \dots \\ \sqrt{2} & 0 & \sqrt{3} & 0 & \dots \\ 0 & \sqrt{3} & 0 & \sqrt{4} & \dots \\ 0 & 0 & \sqrt{4} & 0 & \dots \\ \vdots & \vdots & \vdots & \vdots & \end{pmatrix},$$

$$p = i\sqrt{\frac{\hbar m\omega}{2}} \begin{pmatrix} 0 & -\sqrt{2} & 0 & 0 & \dots \\ \sqrt{2} & 0 & -\sqrt{3} & 0 & \dots \\ 0 & \sqrt{3} & 0 & -\sqrt{4} & \dots \\ 0 & 0 & \sqrt{4} & 0 & \dots \\ \vdots & \vdots & \vdots & \vdots & \end{pmatrix}.$$
(9)

And you can verify that $p^2/2m + (m\omega^2/2)x^2$ is diagonal with the matrix element given by $\hbar\omega(n+1/2)$.

- 6. Griffiths 3.38.
 - (a) By inspection the eignenvalues of **H** are $E_1 = \hbar\omega$, $E_2 = E_3 = 2\hbar\omega$ and the eigenvectors are

$$e_1 = \begin{pmatrix} 1\\0\\0 \end{pmatrix}, \quad e_2 = \begin{pmatrix} 0\\1\\0 \end{pmatrix}, \quad e_3 = \begin{pmatrix} 0\\0\\1 \end{pmatrix}$$

To find the eigenvalues, γ , we set

$$det(\mathbf{A} - \gamma \mathbf{I}) = \begin{vmatrix} \gamma & \lambda & 0 \\ \lambda & \gamma & 0 \\ 0 & 0 & 2\lambda - \gamma \end{vmatrix} = 0$$
$$\Rightarrow \gamma^2 (2\lambda - \gamma) - \lambda^2 (2\lambda - \gamma) = 0$$

Then the normalized eigenvalues of **A** are $A_1 = \lambda$, $A_2 = -\lambda$, $A_3 = 2\lambda$ and the eigenvectors are

$$a_1 = \frac{1}{\sqrt{2}} \begin{pmatrix} 1\\1\\0 \end{pmatrix}, \quad a_2 = \frac{1}{\sqrt{2}} \begin{pmatrix} 1\\-1\\0 \end{pmatrix}, \quad a_3 = \begin{pmatrix} 0\\0\\1 \end{pmatrix}$$

The same strategy gives the eigenvalues **B** are $B_1 = 2\mu$, $B_2 = \mu$, $B_3 = -\mu$ and the eigenvectors are

$$b_3 = \begin{pmatrix} 1\\0\\0 \end{pmatrix}, \quad b_2 = \frac{1}{\sqrt{2}} \begin{pmatrix} 0\\1\\1 \end{pmatrix}, \quad b_3 = \frac{1}{\sqrt{2}} \begin{pmatrix} 0\\1\\-1 \end{pmatrix}$$

(b)

$$\begin{array}{lll} \langle H \rangle &=& \langle S \mid \mathbf{H} \mid S \rangle \\ &=& \hbar \omega \left(c_1 & c_2 & c_3 \right) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 2 & 0 \\ 0 & 0 & 2 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} \\ &=& \hbar \omega (|c_1|^2 + 2(|c_2|^2 + |c_3|^2)) \\ &=& \hbar \omega (2 - |c_1|^2) \end{array}$$

$$\begin{array}{lll} \langle A \rangle & = & \lambda \left(\, c_1 & c_2 & c_3 \, \right) \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 2 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} \\ & = & \lambda (c_1^* c_2 + c_2^* c_1 + 2 |c_3|^2) \end{array}$$

$$\langle B \rangle = \mu \left(c_1 \quad c_2 \quad c_3 \right) \begin{pmatrix} 2 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix}$$
$$= \mu \left(2|c_1|^2 + c_2^* c_3 + c_3^* c_2 \right)$$

(c)

$$S(t)\rangle = c_1 e^{-i\omega t} + c_2 e^{-2i\omega t} + c_3 e^{-2i\omega t}$$

The probability of measuring energies E_1, E_2 and E_3 is $|c_1|^2, |c_2|^2$, and $|c_3|^2$ respectively independent of time. The probability of measuring A_i is $|\langle a_i | S(t) \rangle|^2$

$$\begin{aligned} |\langle a_1 | S(t) \rangle|^2 &= |\frac{1}{\sqrt{2}} \begin{pmatrix} 1 & 1 & 0 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} |^2 &= \frac{1}{2} |c_1 e^{-i\omega t} + c_2 e^{-2i\omega t}|^2 \\ |\langle a_2 | S(t) \rangle|^2 &= |\frac{1}{\sqrt{2}} \begin{pmatrix} 1 & -1 & 0 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} |^2 &= \frac{1}{2} |c_1 e^{-i\omega t} - c_2 e^{-2i\omega t}|^2 \\ |\langle a_3 | S(t) \rangle|^2 &= |(0 & 0 & 1) \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} |^2 &= |c_3|^2 \end{aligned}$$

The probability of measuring B_i is $|\langle b_i \mid S(t) \rangle|^2$

$$|\langle b_1 | S(t) \rangle|^2 = |(1 \ 0 \ 0) \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix}|^2 = \frac{1}{2} |c_1|^2$$

$$\begin{aligned} |\langle b_2 | S(t) \rangle|^2 &= \left| \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 1 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} \right|^2 = \frac{1}{2} |c_2 e^{-2i\omega t} + c_3 e^{-2i\omega t}|^2 = \frac{1}{2} |c_2 + c_3|^2 \\ |\langle b_3 | S(t) \rangle|^2 &= \left| \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & -1 \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \\ c_3 \end{pmatrix} \right|^2 = \frac{1}{2} |c_2 e^{-2i\omega t} - c_3 e^{-2i\omega t}|^2 = \frac{1}{2} |c_2 - c_3|^2 \end{aligned}$$

7. Charmonium. The Schrödinger equation for charmonium is

$$-\frac{\hbar^2}{2m}\nabla^2\psi + \alpha r\psi = E\psi$$

Define $u(r) = r\psi$ and for spherically symmetric wave functions, the Schrödinger equation reduces to

$$-\frac{\hbar^2}{2m}\frac{d^2u}{dr^2} + \alpha ru = Eu$$

Let $r = l_0 z$ and $E = \epsilon E_0$ where z and ϵ are dimensionless and the Schrödinger equation becomes

$$-\frac{\hbar^2}{2ml_0^2}\frac{d^2u}{dz^2} + \alpha l_0 zu = \epsilon E_0 u$$

or

$$-\frac{d^{2}u}{dz^{2}} + \alpha \frac{2ml_{0}^{2}}{\hbar^{2}}l_{0}zu = \epsilon \frac{2ml_{0}^{2}E_{0}}{\hbar^{2}}u$$

Set $l_0 = \left(\frac{\hbar^2}{2m\alpha}\right)^{1/3}$ and $E_0 = \frac{\hbar^2}{2ml_0^2} = \left(\frac{\hbar^2\alpha^2}{2m}\right)^{1/3}$ and our differential equation looks like

$$-\frac{d^2u}{dz^2} + zu = \epsilon u$$

Now let $y = z - \epsilon$ and we have Airy's equation

$$-\frac{d^2u}{dy^2} + yu = 0$$

Since $u(r) = \psi(r)/r$, then it must be that u(0) = 0 so that $\psi(0)$ is finite. Therefore $u(z - \epsilon) = u(-\epsilon) = 0$. The energy eigenvalues, ϵ are the zeros of the Airy function a_i . The first two zeros are 2.3 and 4.1 so $E_1 = 2.3E_0$ and $E_2 = 4.1E_0$.

We have that

$$m_{1s}c^2 = 2m_cc^2 + E_1$$

$$m_{2s}c^2 = 2m_cc^2 + E_2$$
(10)

where m_c is the charmed quark mass and E_1 and E_2 are the binding energies. The difference of the two equations yields $E_0 = (m_{2s} -$ $m_{1s})c^2/1.8 = 0.325 GeV$. We find $m_c c^2 = 1.176 GeV$. Finally the reduced mass $m = m_c/2$. Meanwhile,

$$E_0 = \left(\frac{\hbar^2 \alpha^2}{2m}\right)^{1/3}$$

$$\Rightarrow \alpha = \frac{2mE_0^3}{\hbar^2} = \frac{m_c c^2 E_0^3}{\hbar^2 c^2} = \frac{(1.176 GeV)(0.325 GeV)^3}{(.197 GeV - fm)^2} = 1.02 GeV/fm$$

And

$$l_0 = \left(\frac{\hbar^2 c^2}{m_c c^2 \alpha}\right)^{1/3} = \left(\frac{(0.197 GeV - fm)^2}{(1.176 GeV)(1.02 GeV/fm)}\right)^{1/3} = 0.318 fm.$$

8. Rotations. We define $x = iL_y\theta/\hbar$. Notice that

$$\begin{aligned} x &= \frac{\theta}{2} \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}, \\ x^2 &= -\left(\frac{\theta}{2}\right)^2 \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \\ x^3 &= -\left(\frac{\theta}{2}\right)^3 \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix}, \\ x^2 &= \left(\frac{\theta}{2}\right)^4 \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}, \end{aligned}$$

In particular

$$\begin{aligned} R(\theta) &= e^x &= 1 + x + \frac{x^2}{2} + \frac{x^3}{3!} + \dots, \\ &= \begin{pmatrix} 1 & 0\\ 0 & 1 \end{pmatrix} \left(1 - \frac{1}{2!} \begin{pmatrix} \theta\\ 2 \end{pmatrix}^2 + \dots \right) + \begin{pmatrix} 0 & 1\\ -1 & 0 \end{pmatrix} \left(\begin{pmatrix} \theta\\ 2 \end{pmatrix} - \frac{1}{3!} \begin{pmatrix} \theta\\ 2 \end{pmatrix}^3 + \dots \right) + \begin{pmatrix} 0 & 1\\ -1 & 0 \end{pmatrix} \right) \\ &= \begin{pmatrix} 1 & 0\\ 0 & 1 \end{pmatrix} \cos(\frac{\theta}{2}) + \begin{pmatrix} 0 & 1\\ -1 & 0 \end{pmatrix} \sin(\frac{\theta}{2}), \\ &= \begin{pmatrix} \cos(\frac{\theta}{2}) & \sin(\frac{\theta}{2})\\ -\sin(\frac{\theta}{2}) & \cos(\frac{\theta}{2}) \end{pmatrix}. \end{aligned}$$

You can see that $L_y^{\dagger} = L_y$ and $R(\theta)^T R(\theta) = 1$, making L_y Hermitian and $R(\theta)$ unitary.