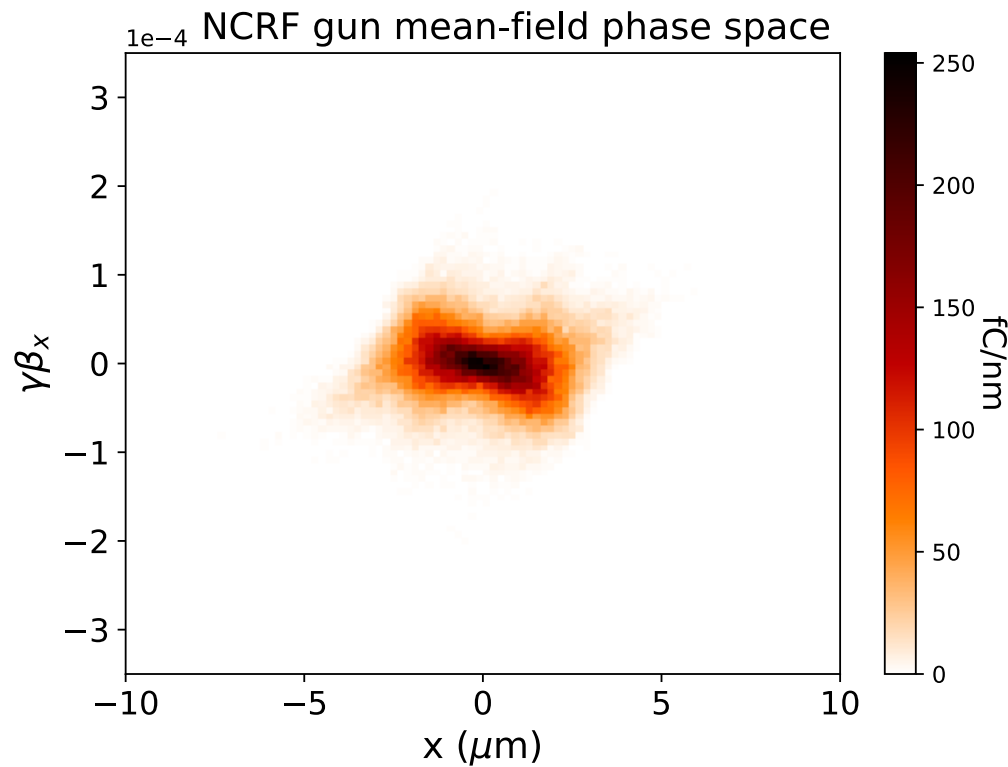
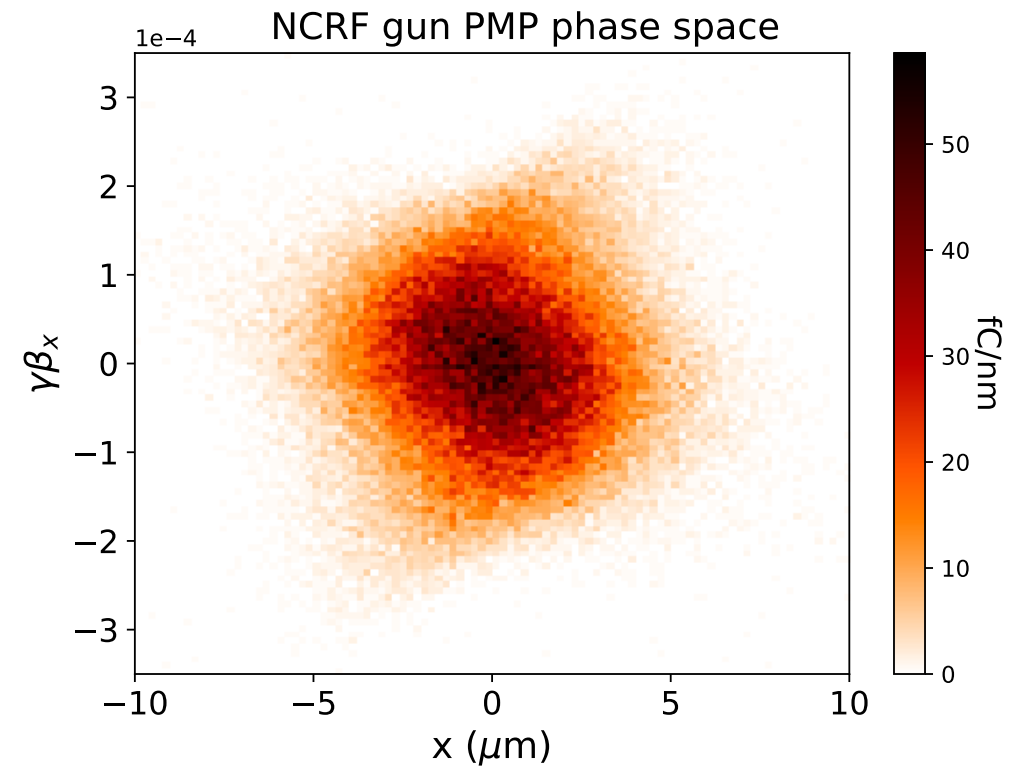


Beyond Space Charge: Simulating Ultracold Photoemission Beamlines

Matt Gordon



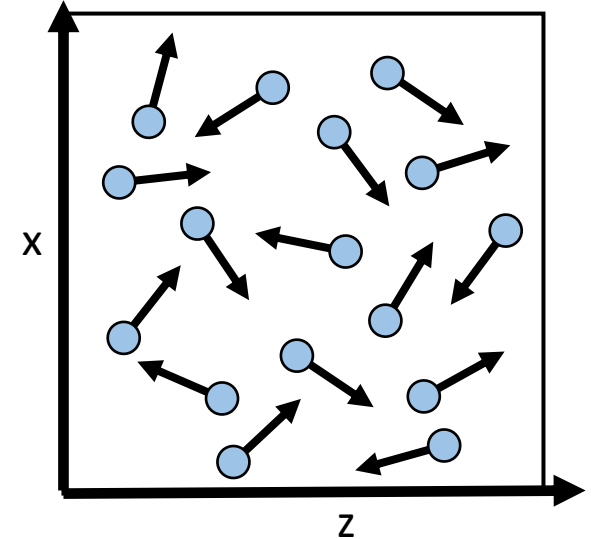
Add Reality



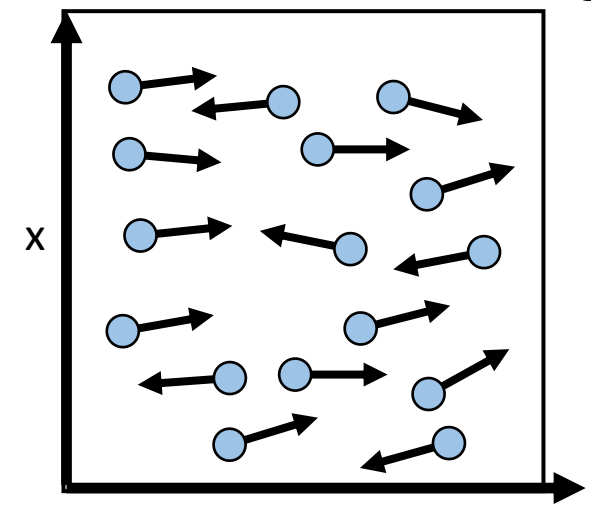
Motivation

- Typical photocathodes have a mean transverse energy of hundreds of meV
- Want to reduce this to tens or even single digit meV
- At low temperatures, typical approximations used for electron interactions in simulation become less valid
- When will reducing MTE no longer improve beam quality?
- **Simulate the transport of cold electron beams**

Large Mean Transverse Energy

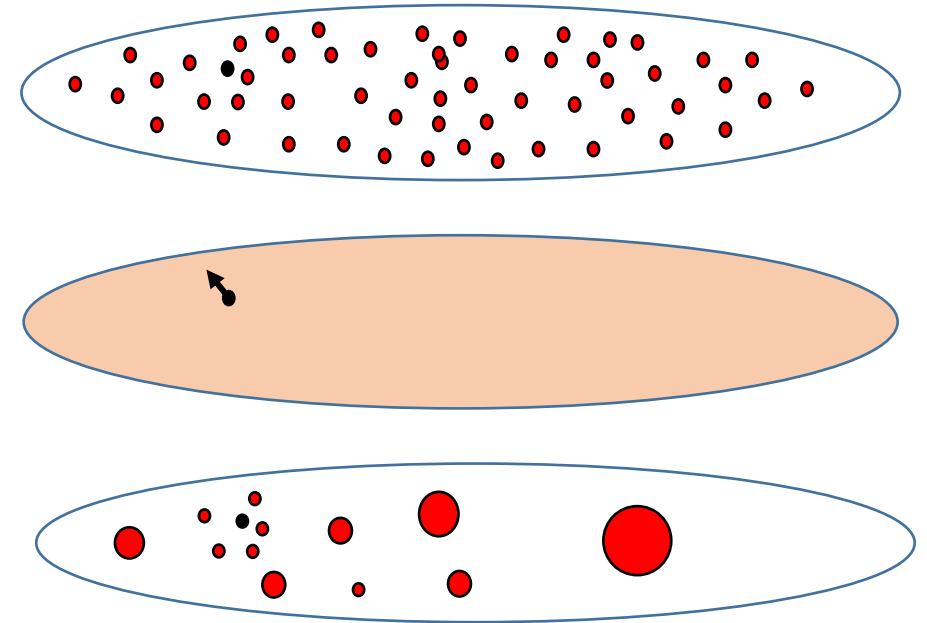


Small Mean Transverse Energy



How are Electron Interactions Simulated?

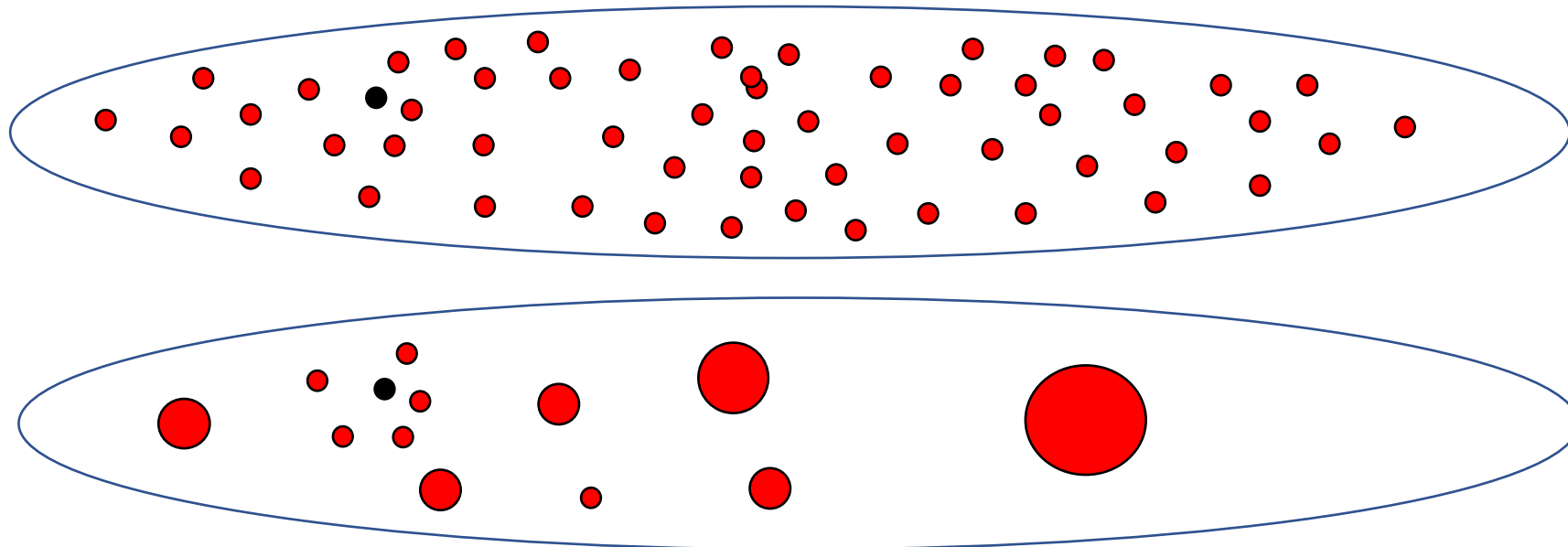
- An electron beam is a collection of point particles
- Exact interaction is computationally expensive
- Charge smoothed in a beam to approximate the interaction (space charge)
- Want to calculate short range interactions precisely and approximate long range interactions
- One method to do this is the Barnes-Hut Algorithm



From top to bottom:
Full Beam Brute Force Calculation
Mean-field Approximation
Barnes-Hut Approximation

Faster Simulation of Non-Mean-field Space Charge

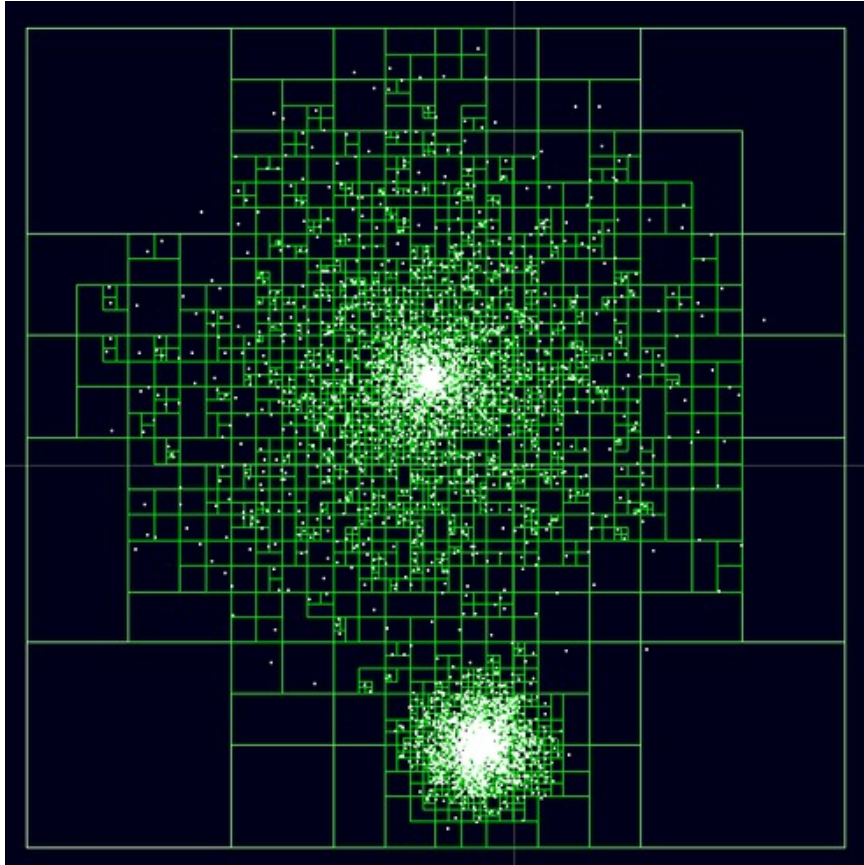
- To make simulating non-mean-field effects possible, rely on the multipole expansion
- E field of collection of charges approximated as series of static multipoles
- The Barnes Hut Tree Algorithm finds which groups of particles are far enough away, such that the monopole term is the only term which contributes to a desired accuracy
- Approximates long range forces, exactly calculates interactions from nearby particles



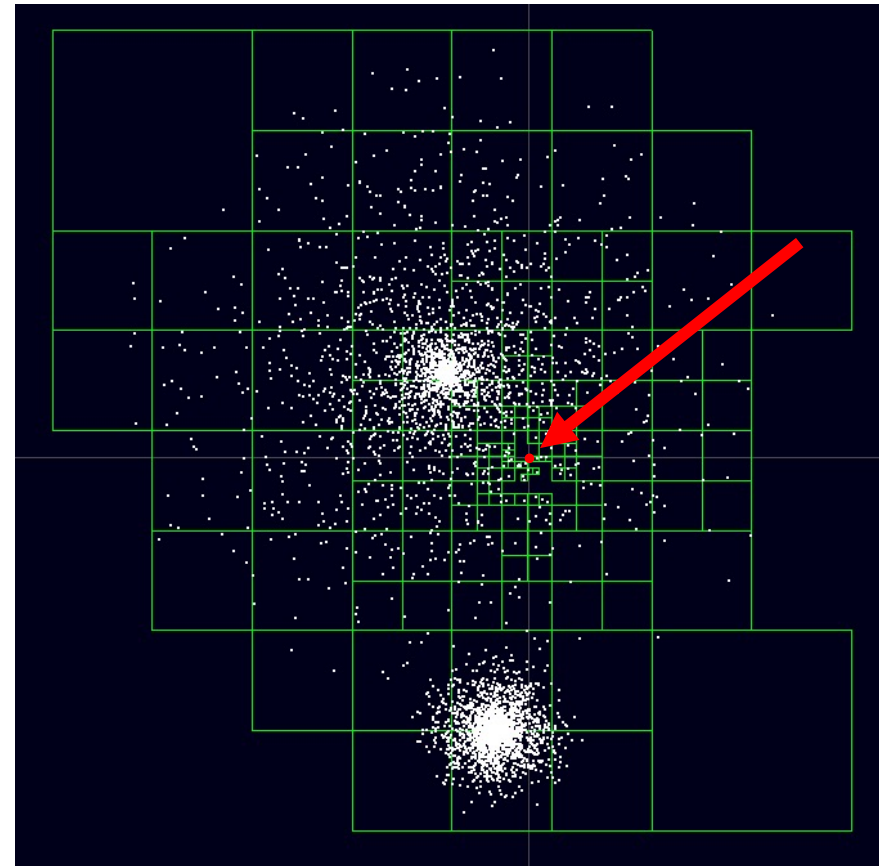
Barnes-Hut Algorithm

- Barnes-Hut tree algorithm for simulation of 2 nearby galaxies

Full Barnes-Hut Tree

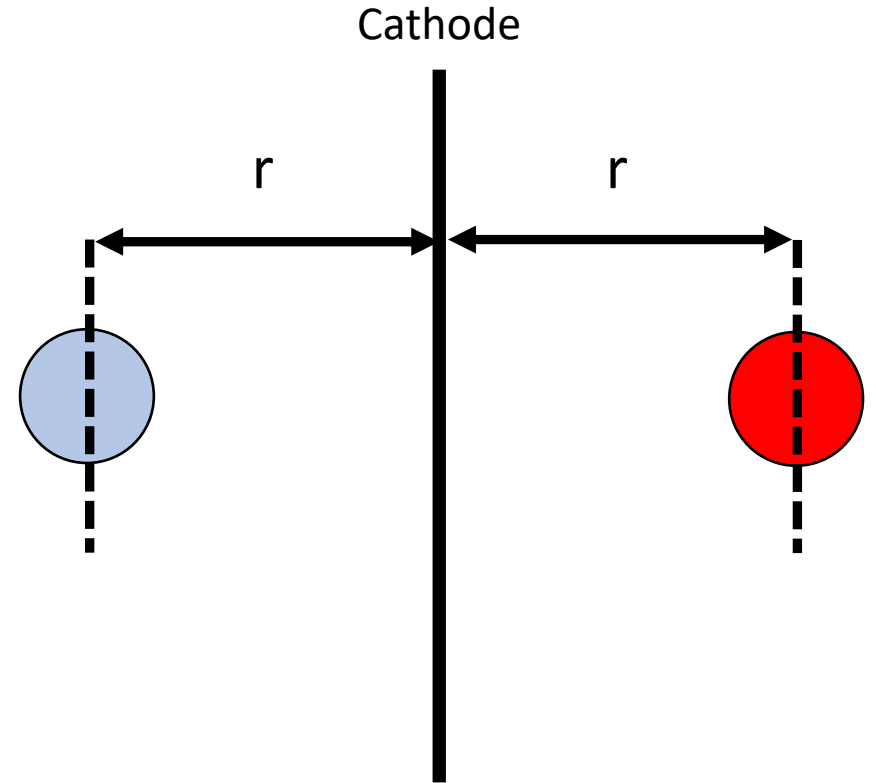


Grouped Nodes for force calculation of particle at origin (red)



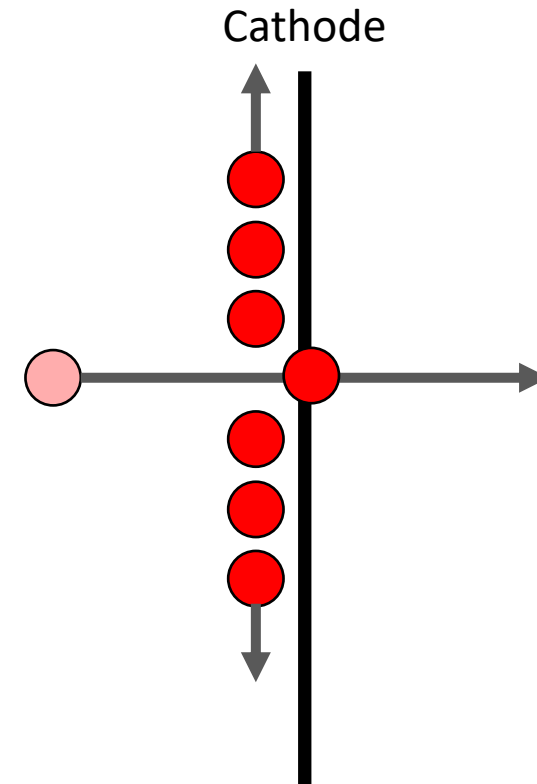
Cathode Divergence

- Model cathode interaction with image charge method
- Image potential diverges as distance from cathode goes to 0
- Not physical, we know electrons can escape
- Need to model the cathode in a different way



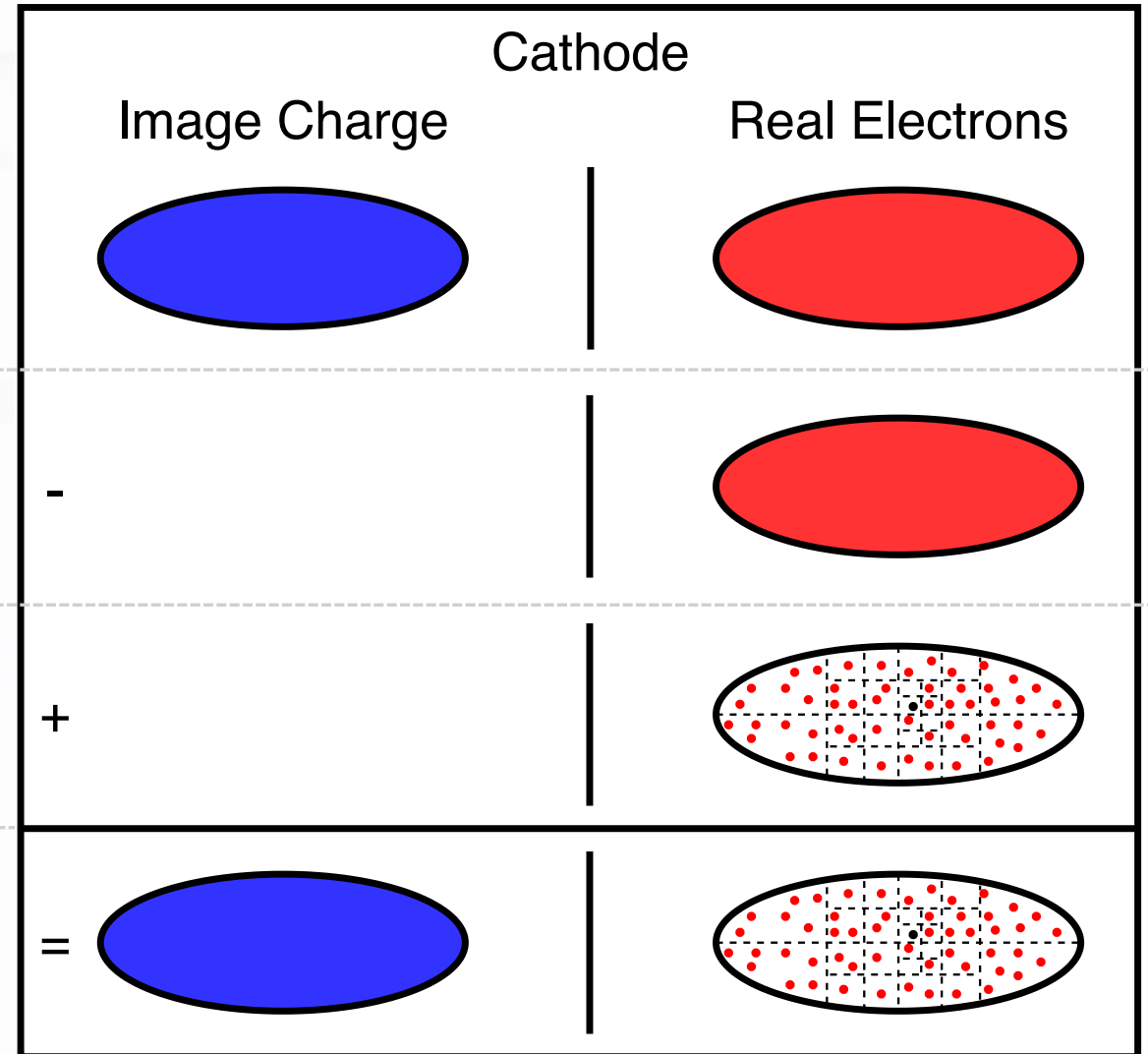
Dynamic Image Charge Method

- Semiclassical approximation of photoemission
- Image charge form on timescale set by the cathode material
- Image potential is velocity dependent and non-divergent
- Self-consistently solve for the image potential for different starting energies

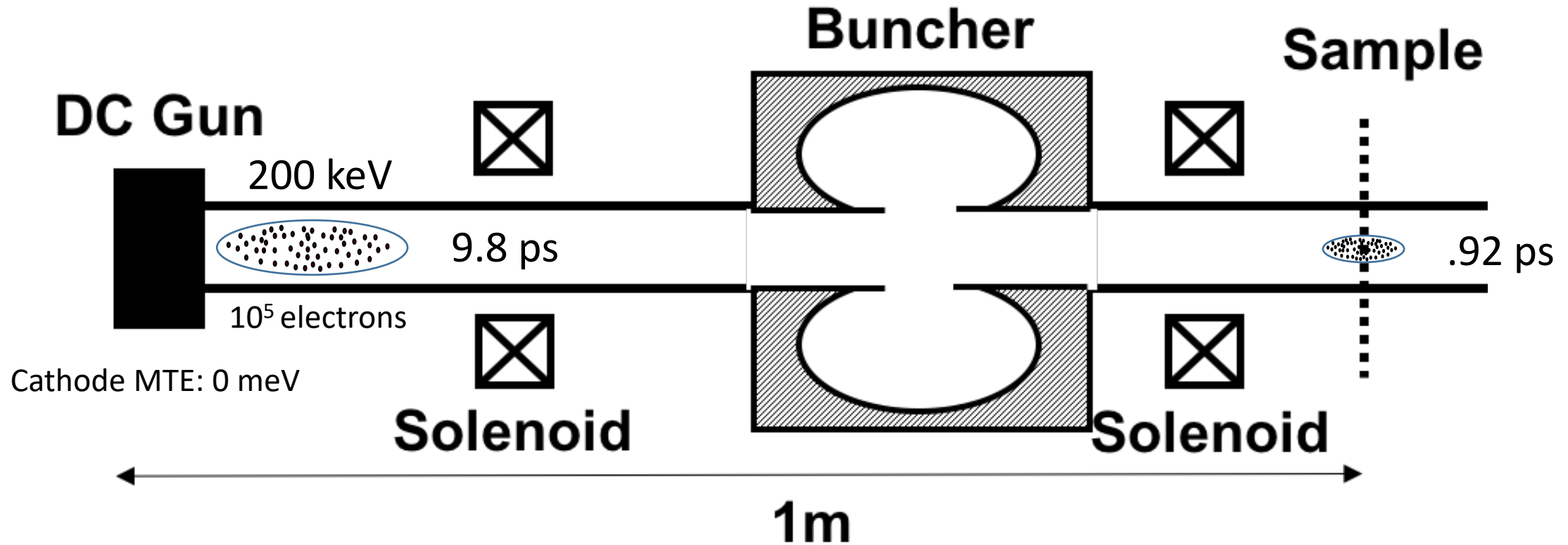


The Plus-Minus-Plus(PMP) Method

- Fields calculated in a 3 step process
- We will calculate the mean-field electric fields including cathode
- Subtract out the mean field calculation without the cathode
- Add in the point-to-point interaction of the real particles
- Final Result

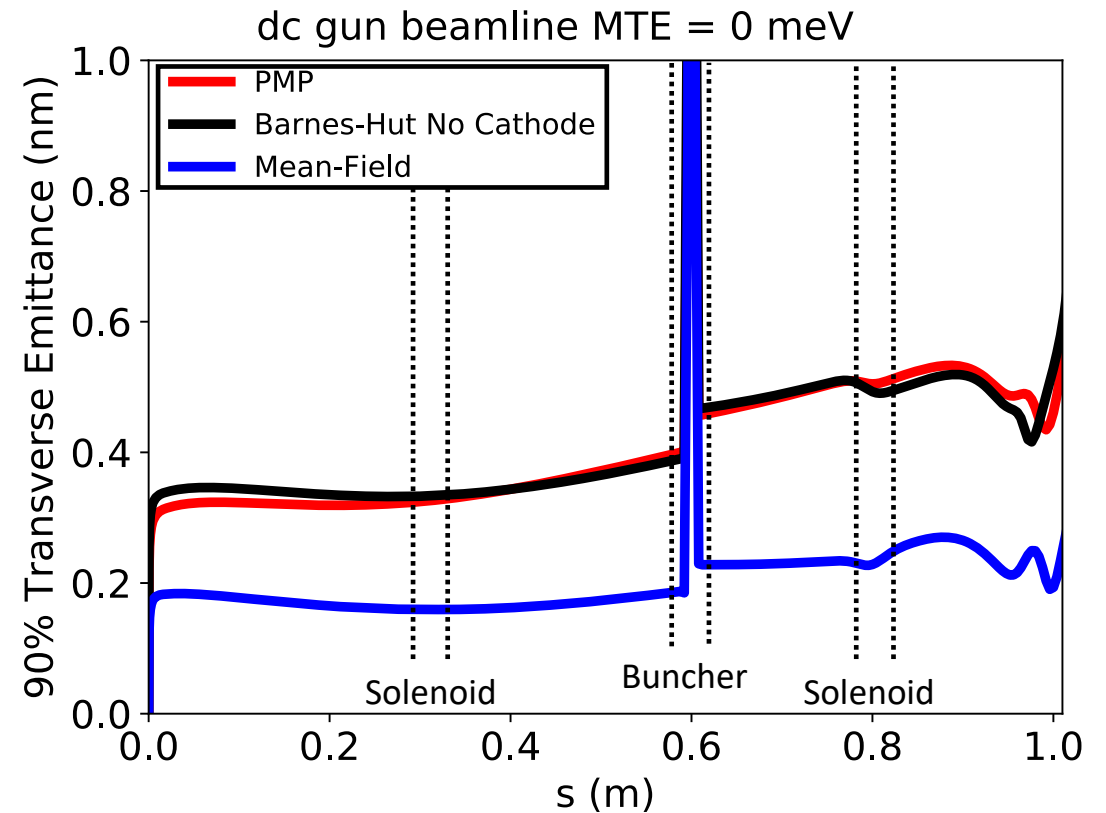
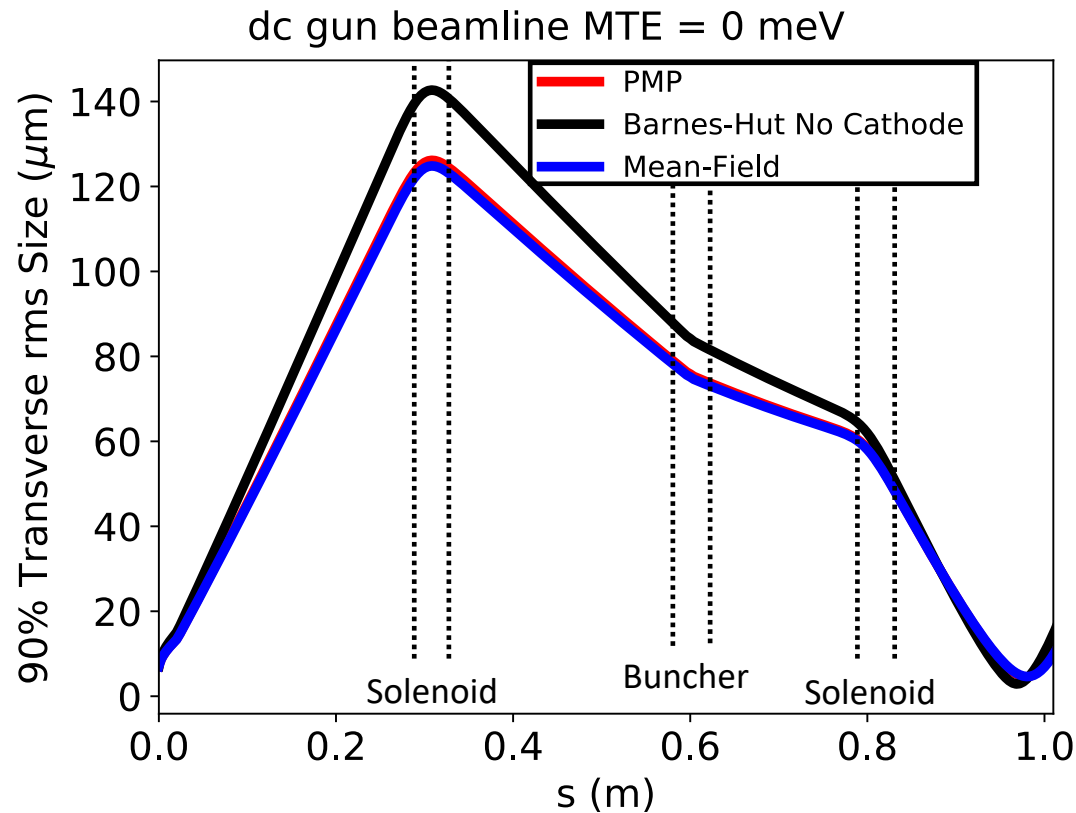


DC Ultrafast Electron Diffraction Beamline Details



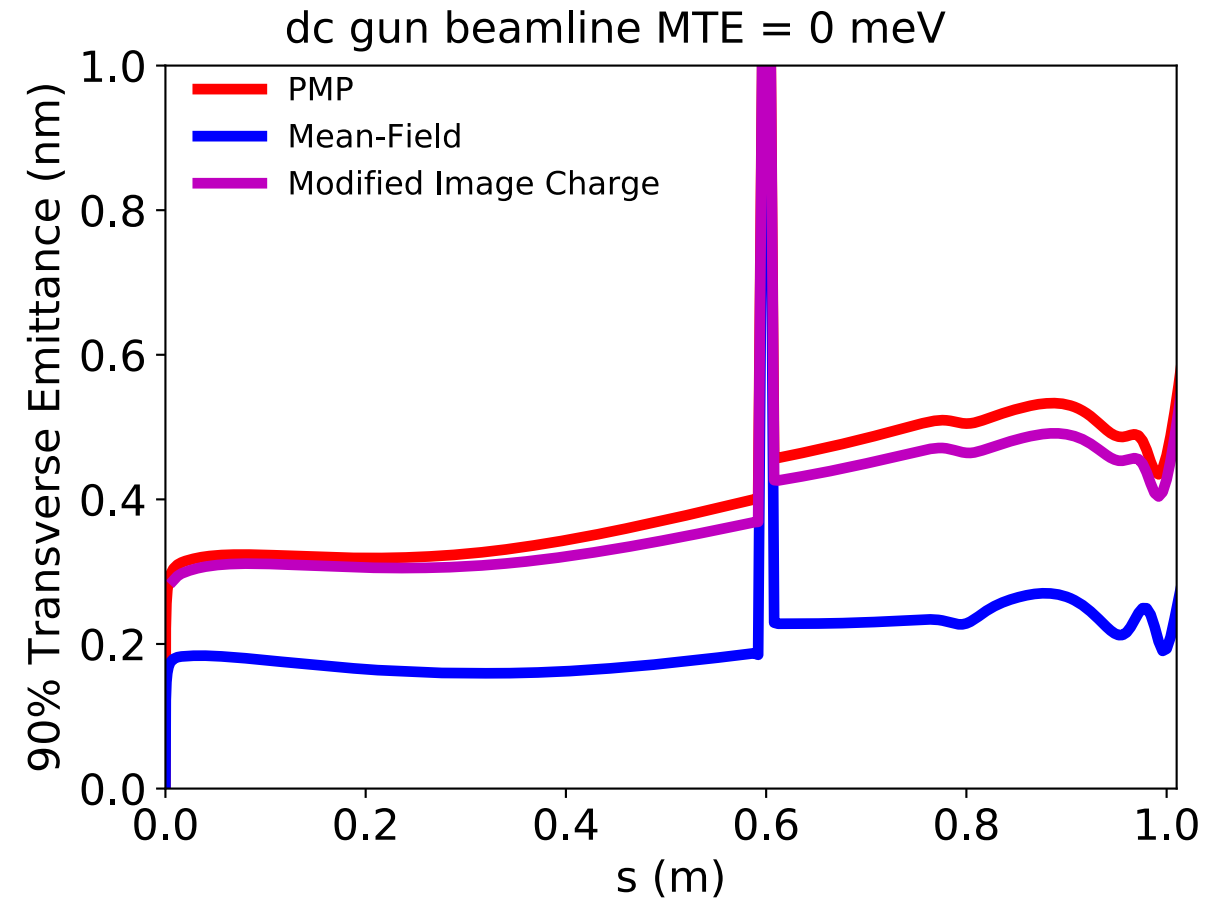
DC Beamline Simulations

- Cathode effects needed to model beam correctly
- Point-to-point interactions increase emittance by a factor of 2

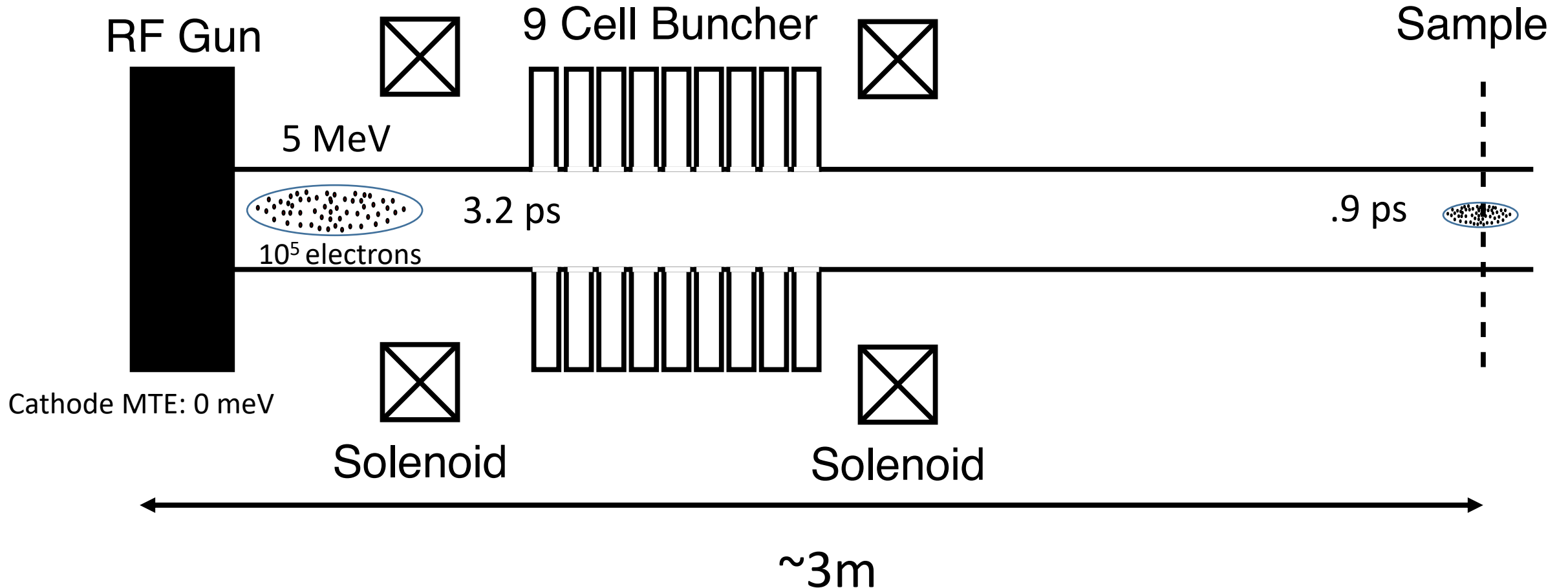


Dynamic Image Charge Comparison

- See how well the more exact calculation compares to the PMP method
- Qualitatively, the graph behave similarly
- Compares to PMP method quite well

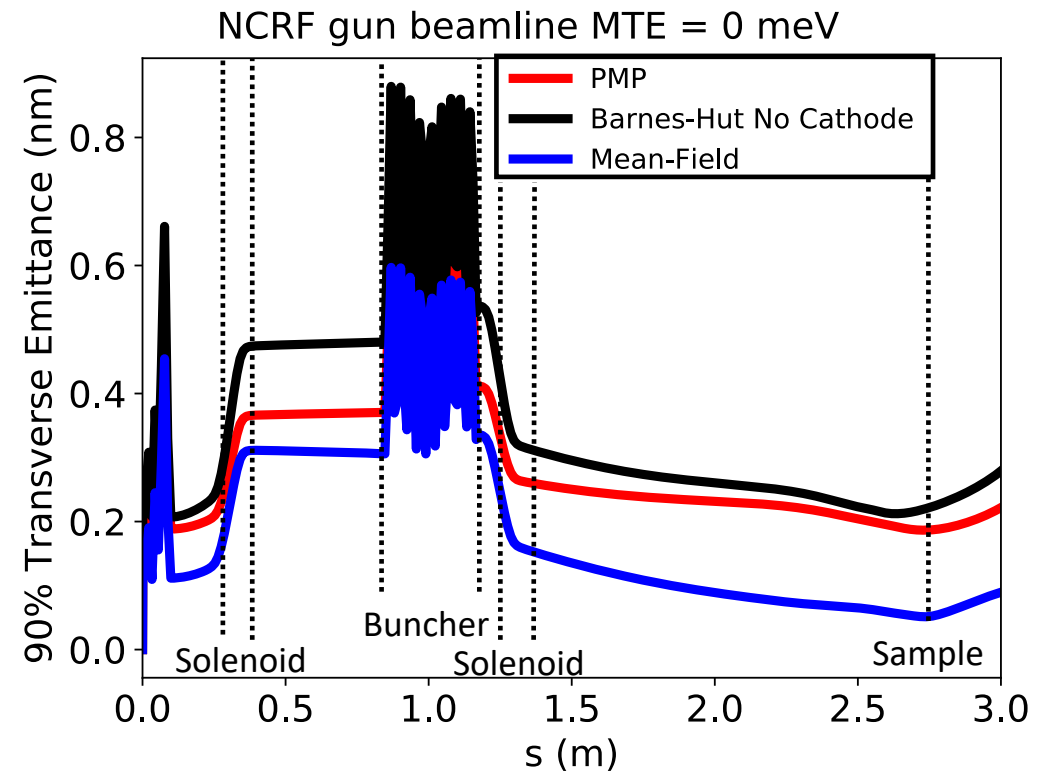
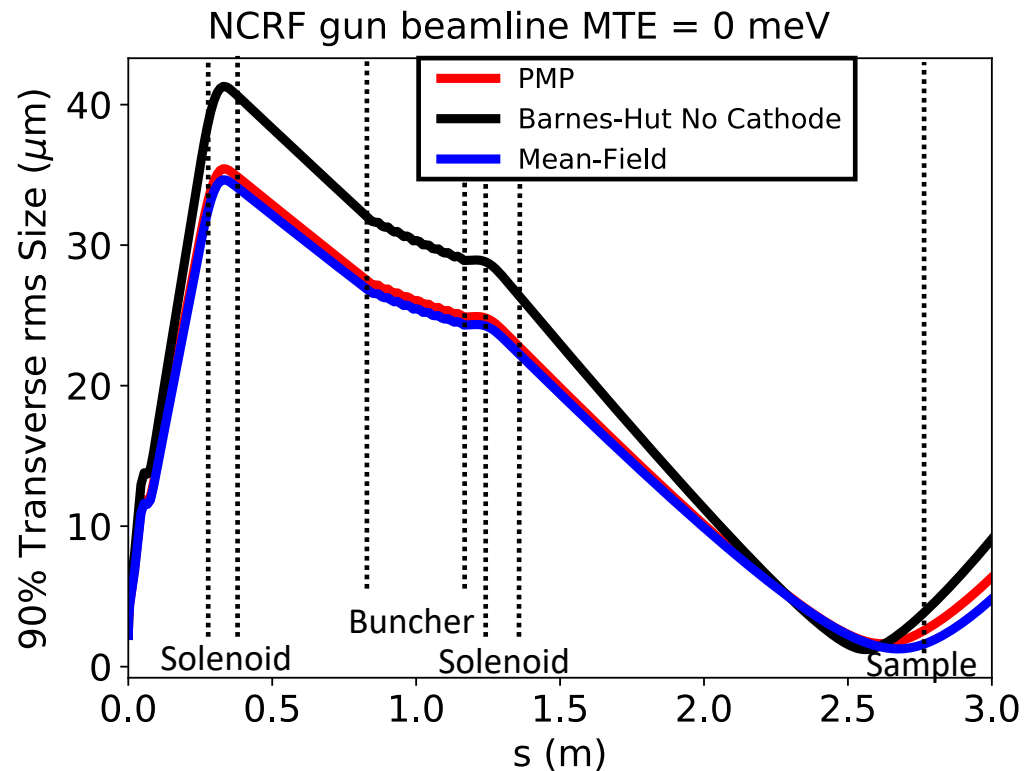


RF Ultrafast Electron Diffraction Beamline Details



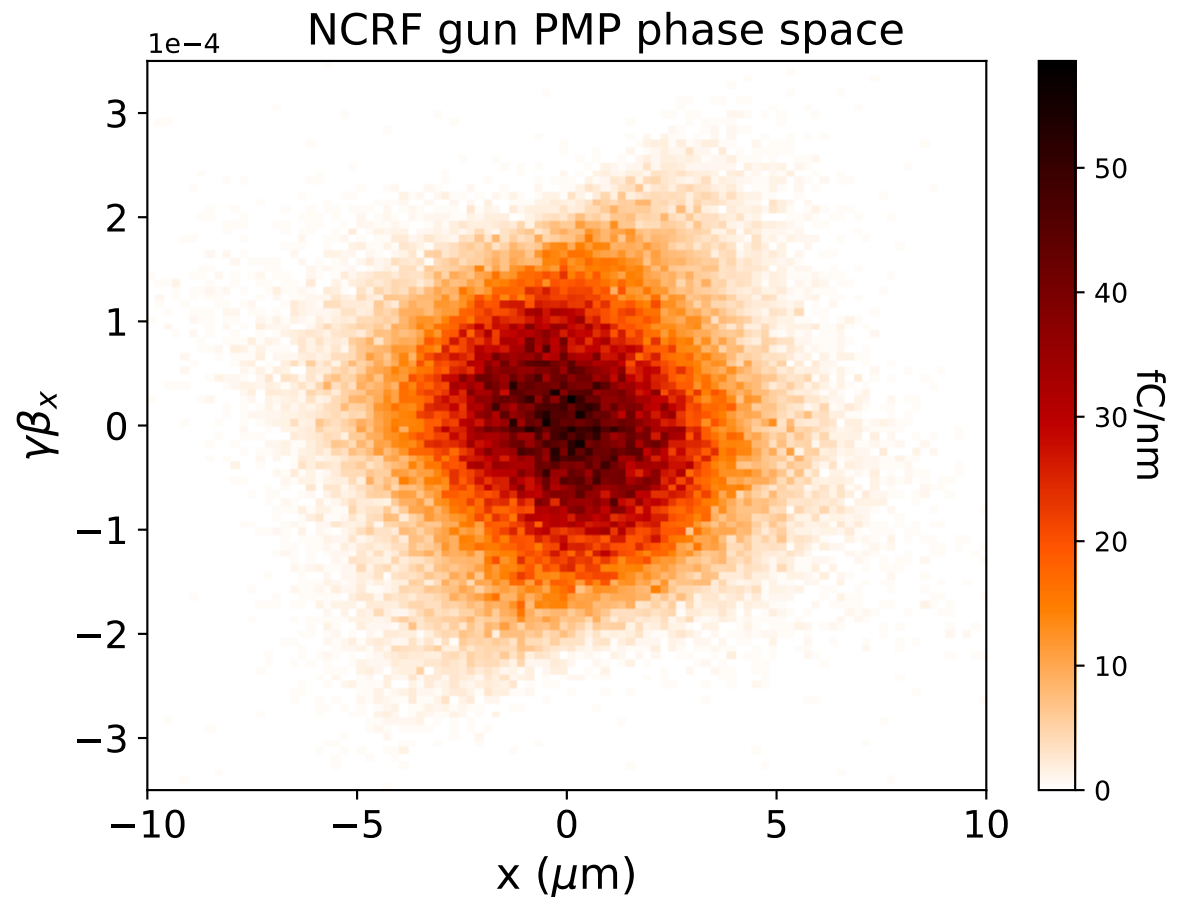
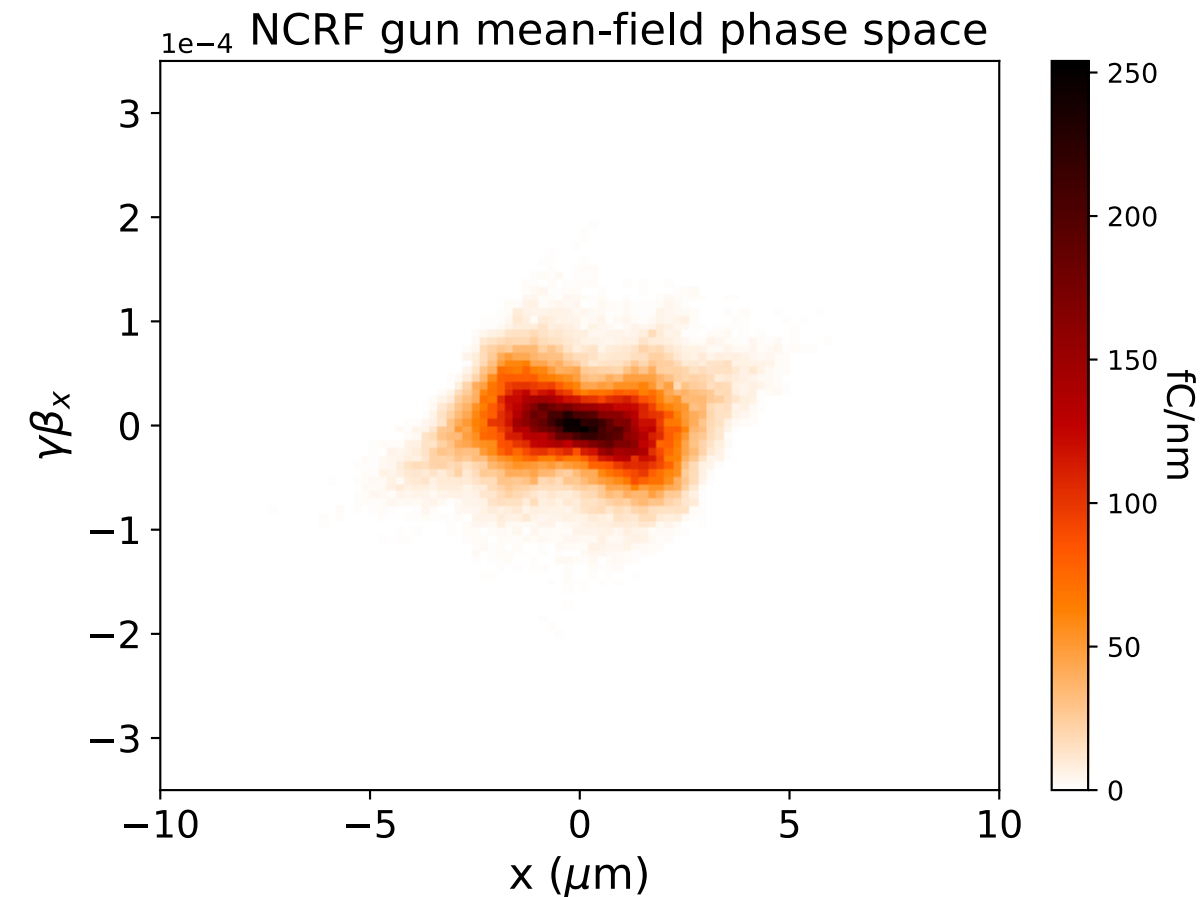
RF Beamline Simulations

- At higher beam densities P2P effects matter more
- Point-to-point interactions increase emittance by a factor of 3.7



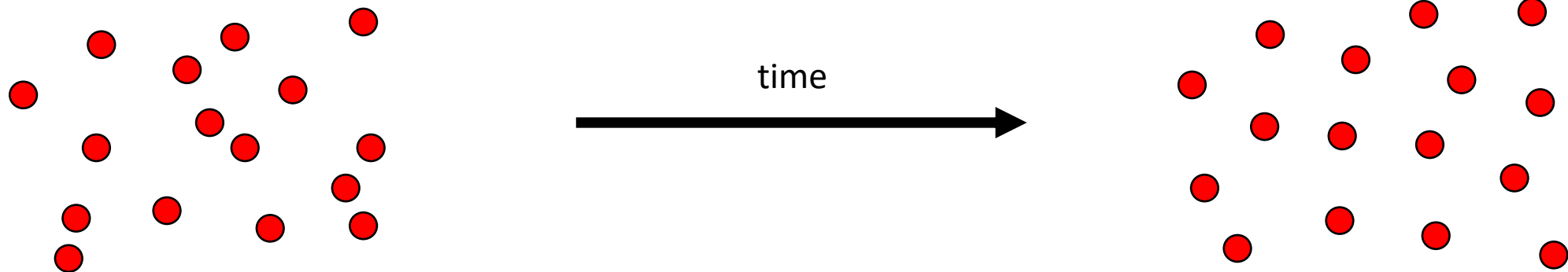
Phase Space Portraits

- Core density drops by a factor of 4
- Not just an effect in the tails
- What caused this?



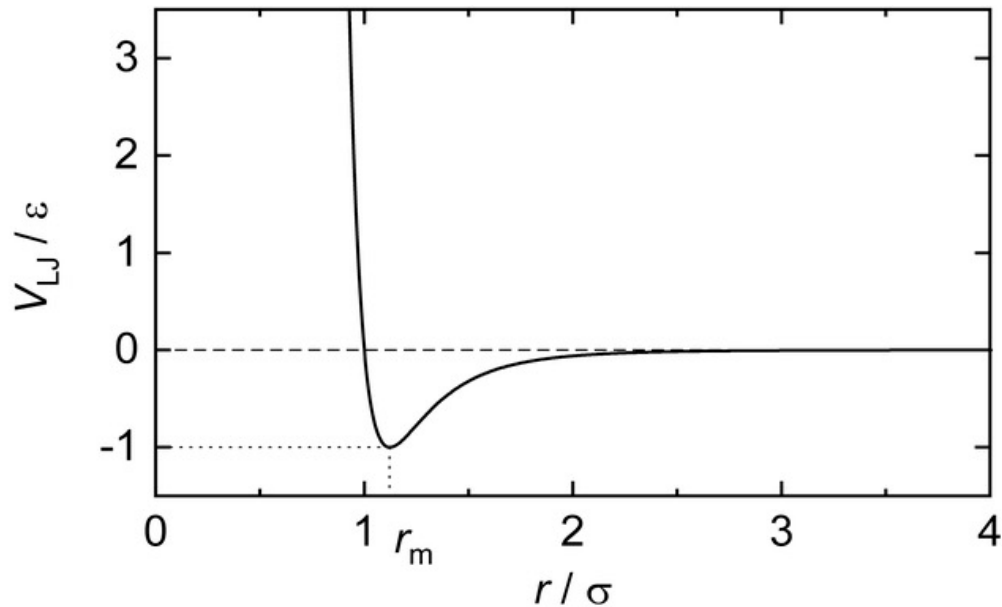
Disorder Induced Heating (The Coulomb Hole)

- The probability of finding an electron in a small region near another electron is near 0 due to Coulomb repulsion
- Charges with positions randomly chosen from a uniform distribution have a higher potential energy than if the charges were ordered
- If the kinetic energy of the particles is small enough, the charges will interact such that the charges become more evenly spaced and thus will warm up

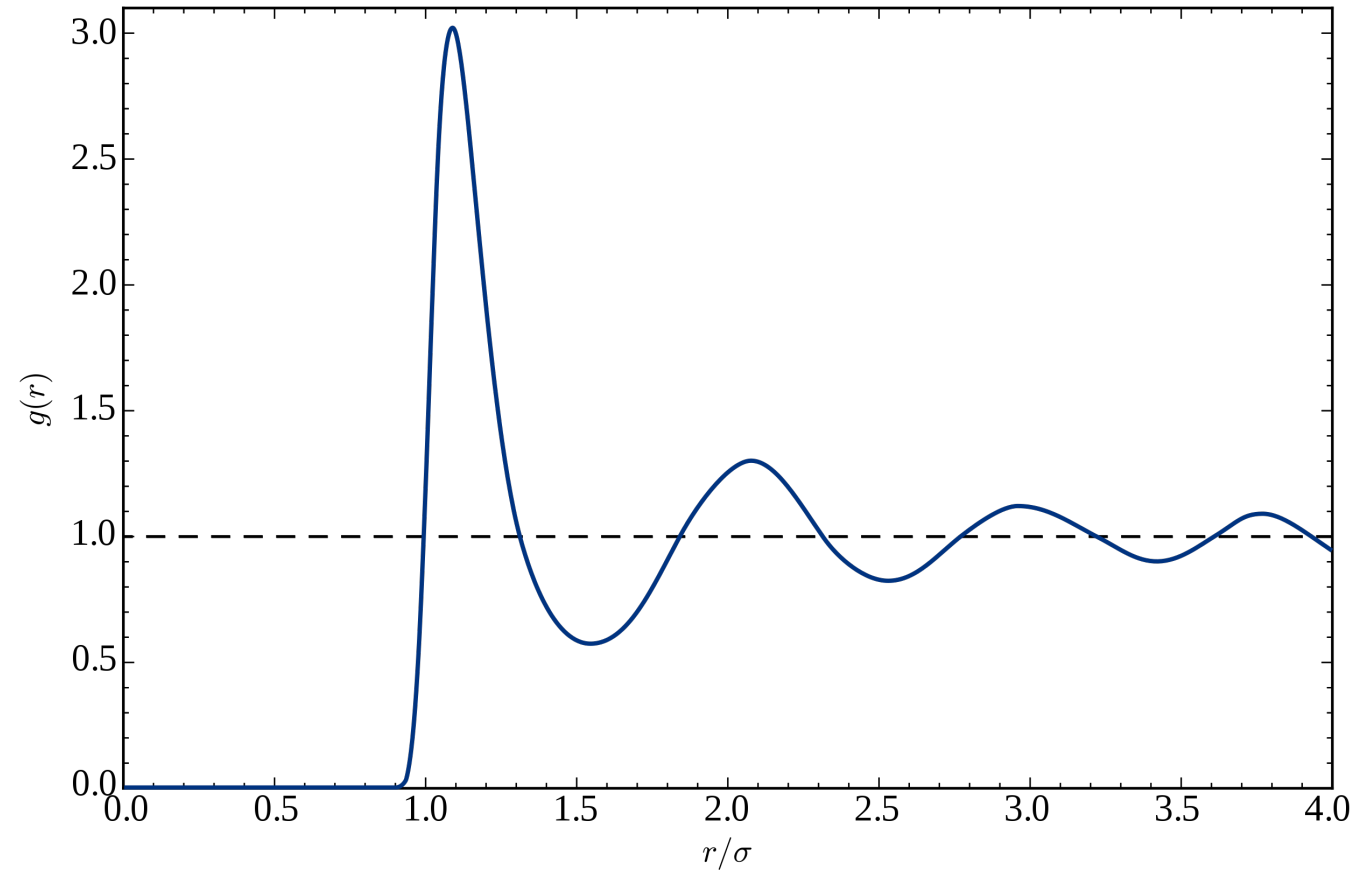


Radial Distribution Function $g(r)$

- How the density of particles varies as a function of distance from a reference particle
- $\rho(r) = \rho g(r)$



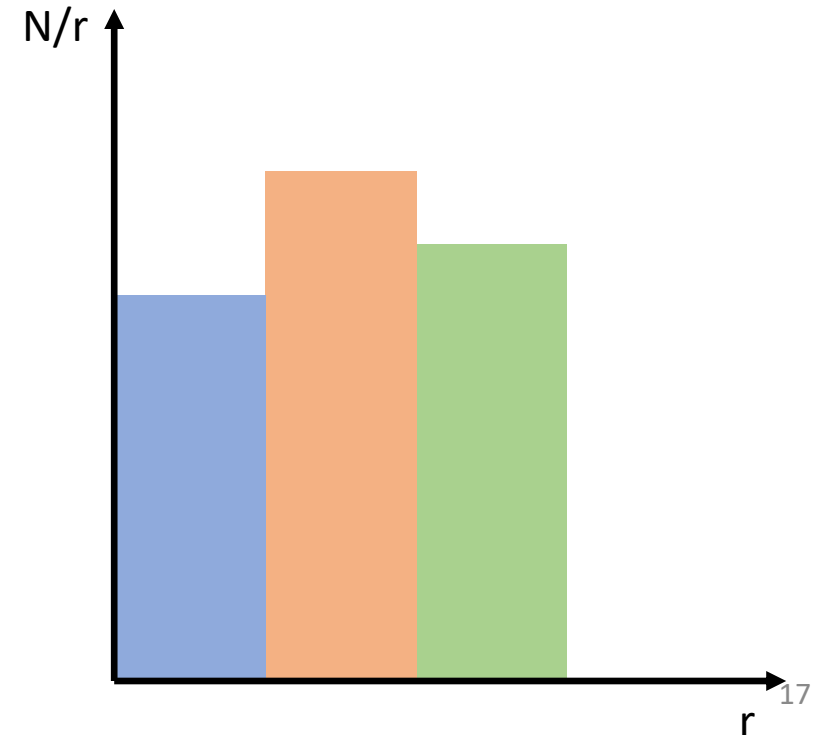
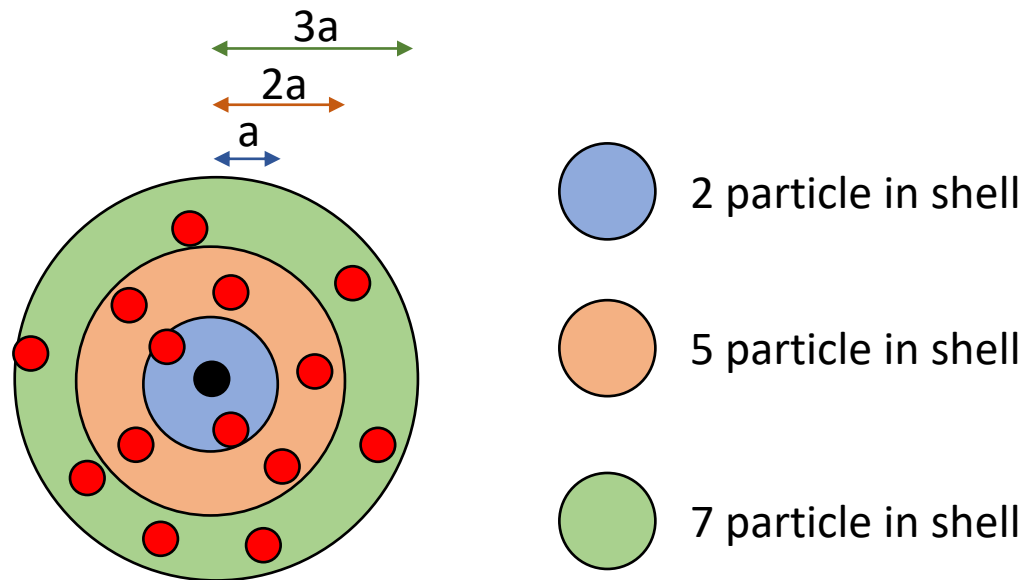
Lennard-Jones potential



Radial distribution function of the Lennard-Jones potential

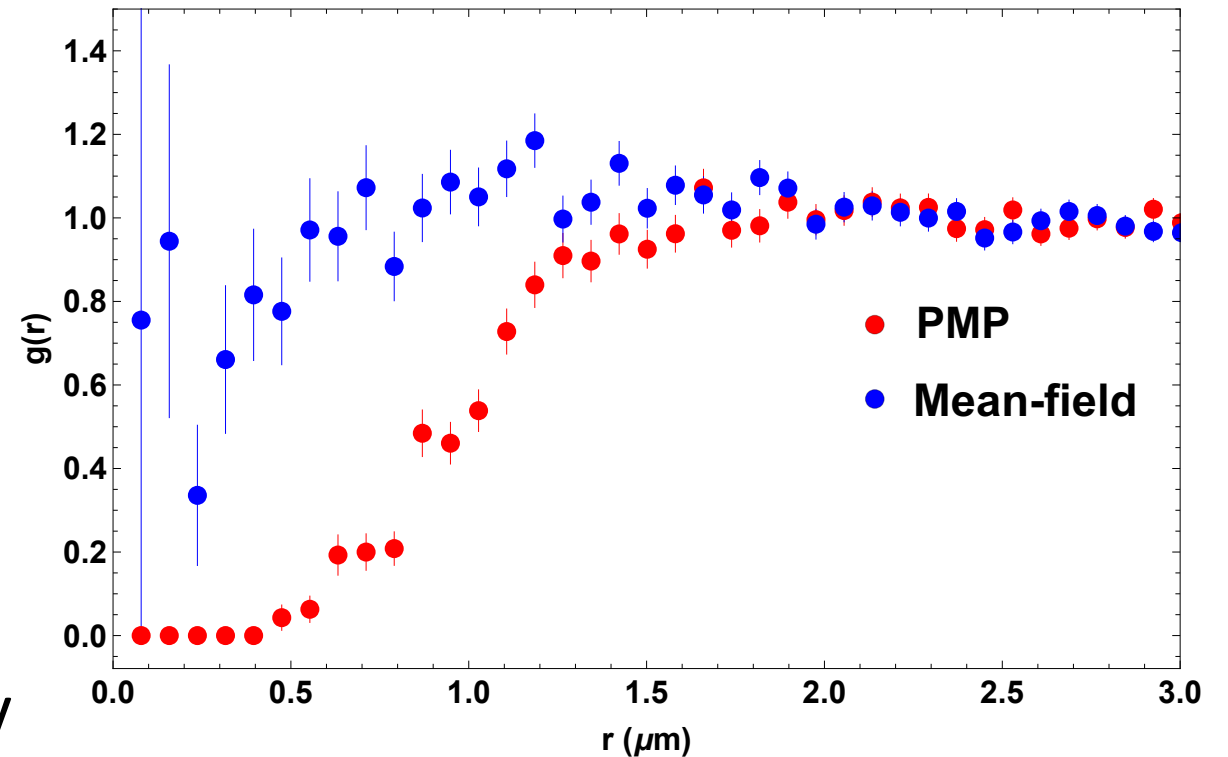
Calculating $g(r)$

- Find distance between 1 particle and every other particle
- Bin the results and normalize the number of particles in the shell by the volume of the shell (for 3D $4\pi r^2 \Delta r$)
- Repeat with all other particles and make an average
- This histogram plots $\rho^* g(r)$ vs r



The Coulomb Hole

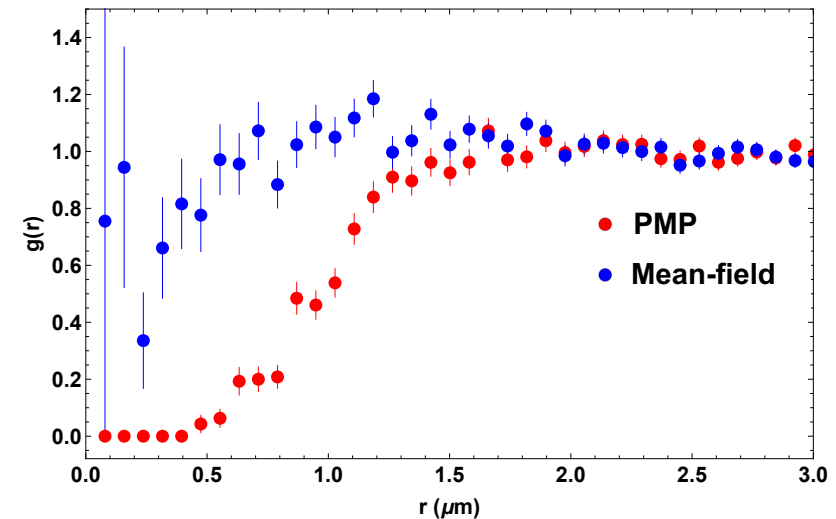
- $g(r)$ 3 mm away from the cathode
- A perfectly uniform distribution would have a constant $g(r)$
- Both distributions started flat up to statistical noise
- PMP simulation $g(r)$ becomes visibly non-flat for small r



Potential Energy From $g(r)$

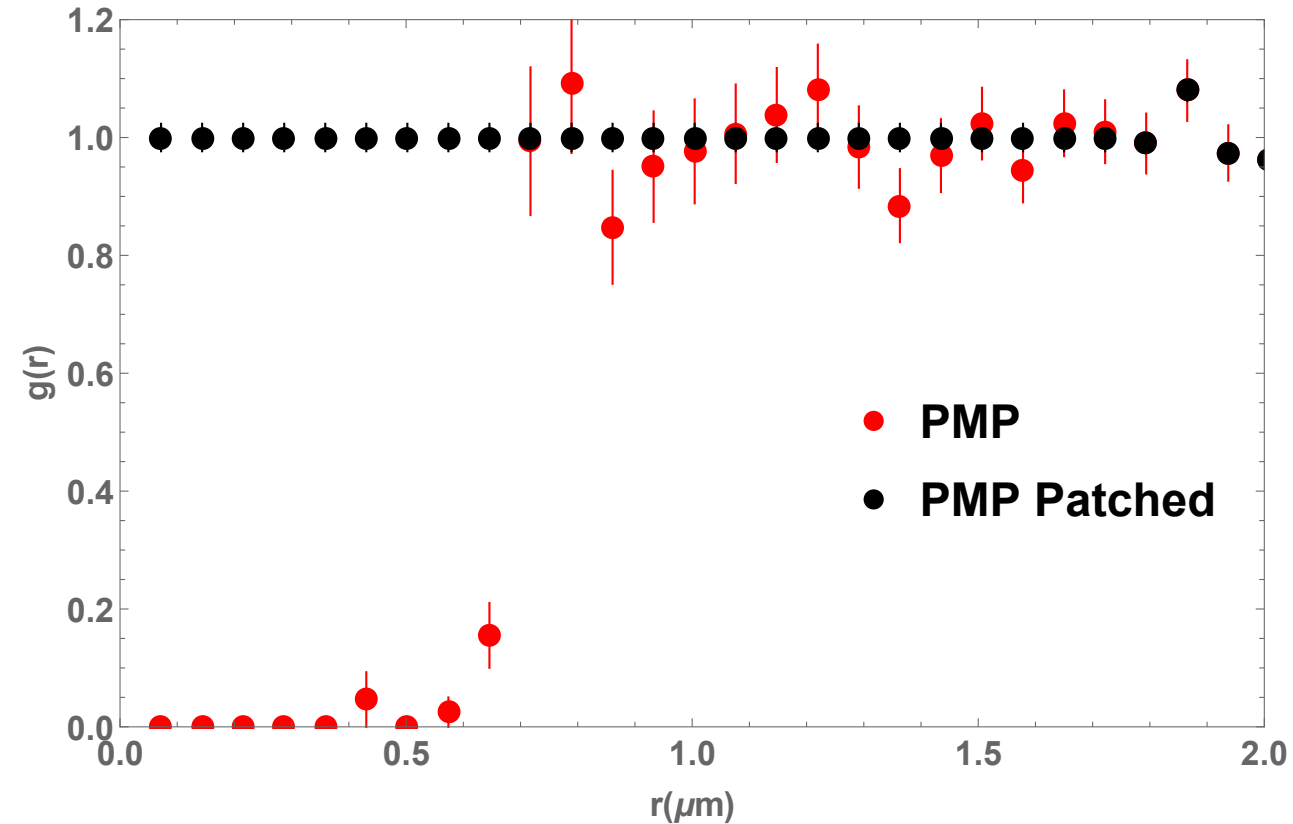
- To find the heating from this, we will find the change in potential energy
- Let $u(r)$ be the interaction potential between particles in the system
- The potential energy of a single particle due to this interaction is:

$$E_{pot}(t) = \int_0^{\infty} dr 4\pi r^2 \rho g(r, t) u(r)$$



Energy of the Coulomb Hole

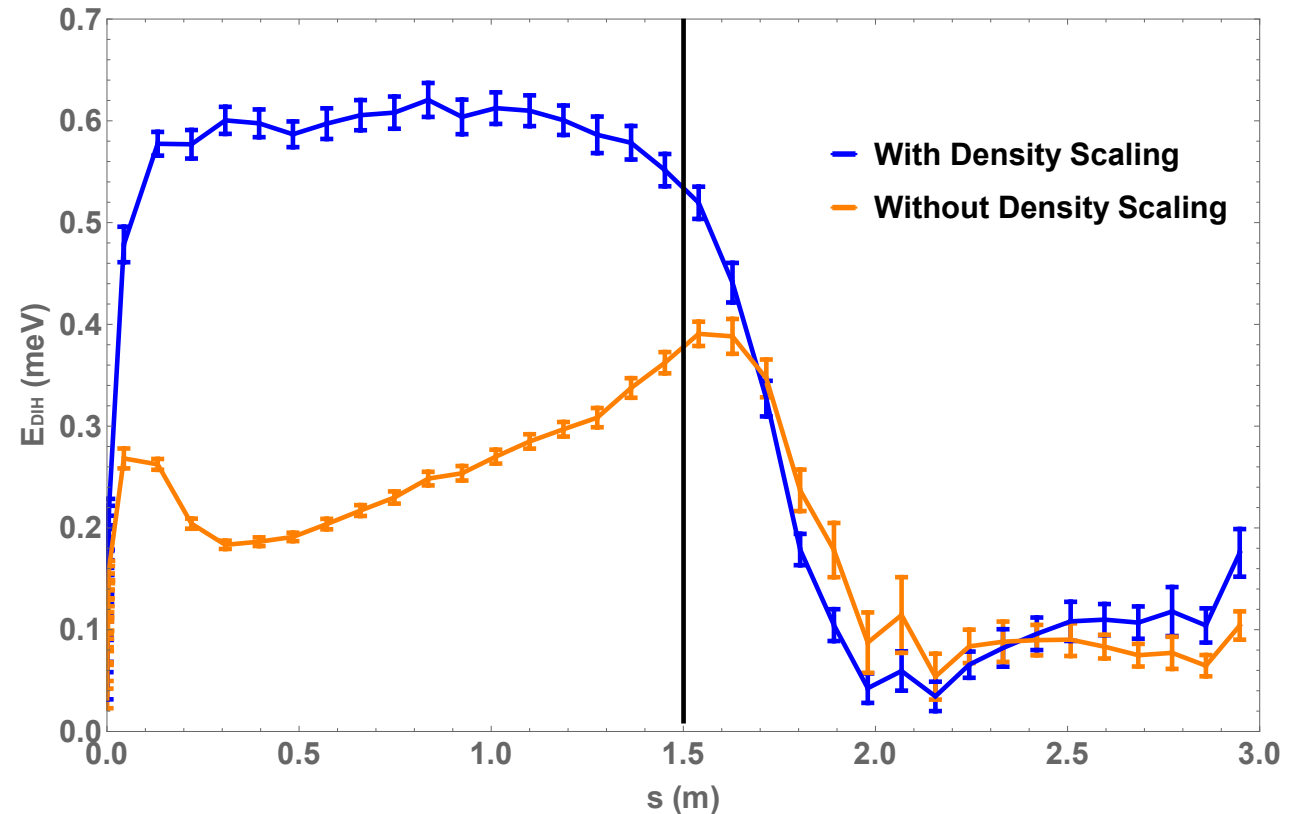
- “Patch up” the Coulomb hole
- Potential energy for these 2 different $g(r)$ can be calculated
- Subtract to roughly calculate the energy from the Coulomb hole



*This plot comes from a simulation at a lower particle density than the plots shown before

Energy of the Coulomb Hole

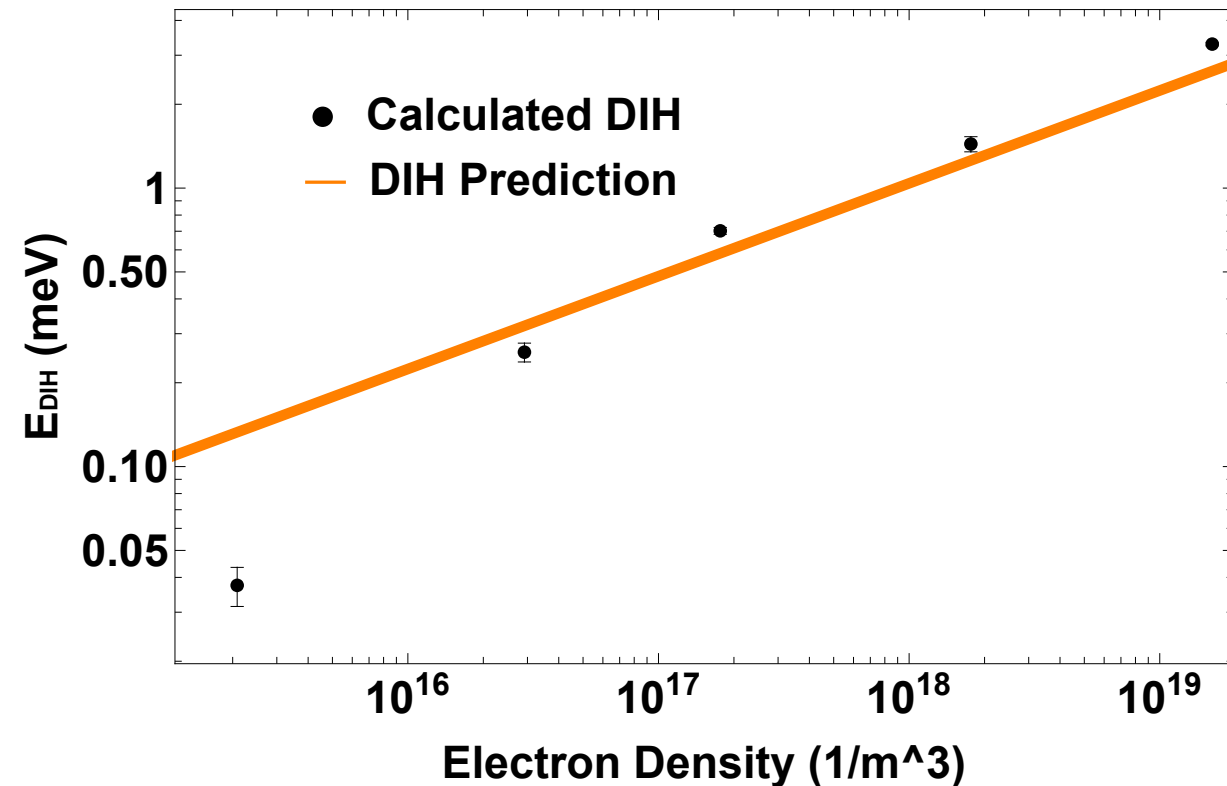
- Subtracting these energies gives us the energy of disorder induced heating, E_{DIH}
- When the beam travels further down the beamline, you can extract out E_{DIH} as long as the shape of $g(r)$ doesn't change
- One common way to lose the shape is to go through a focus



*This plot comes from a simulation at a lower particle density than the plots shown before

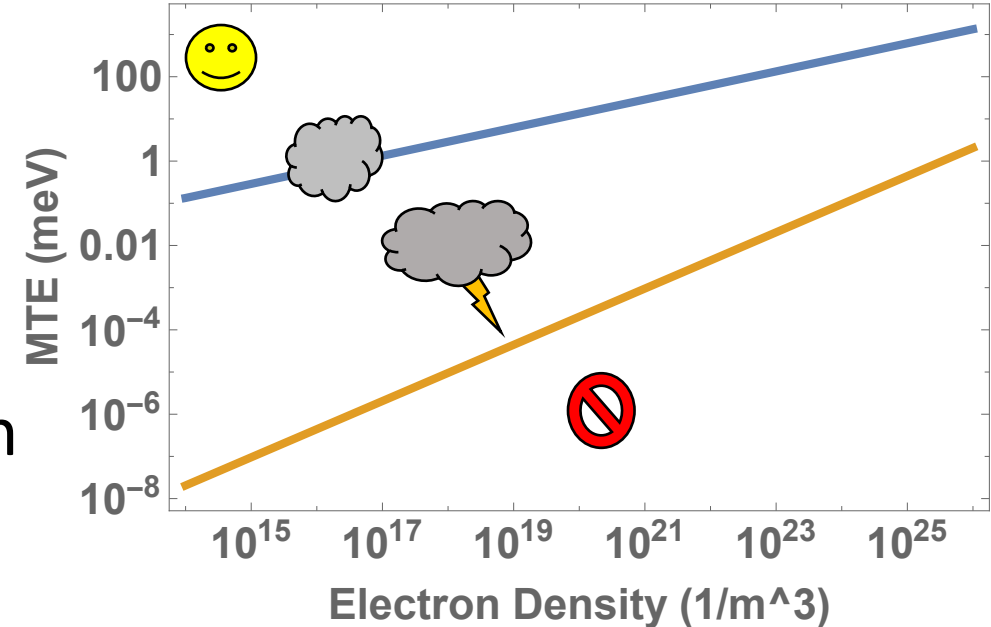
Disorder Induced Heating

- Found heating for several densities
- Disorder induced heating scales with density to the $1/3$ power
- For large densities, heating scales with density with power of $.39 \pm .03$
- At low densities, you can “outrun” the heating effect



“What can I do with this?”

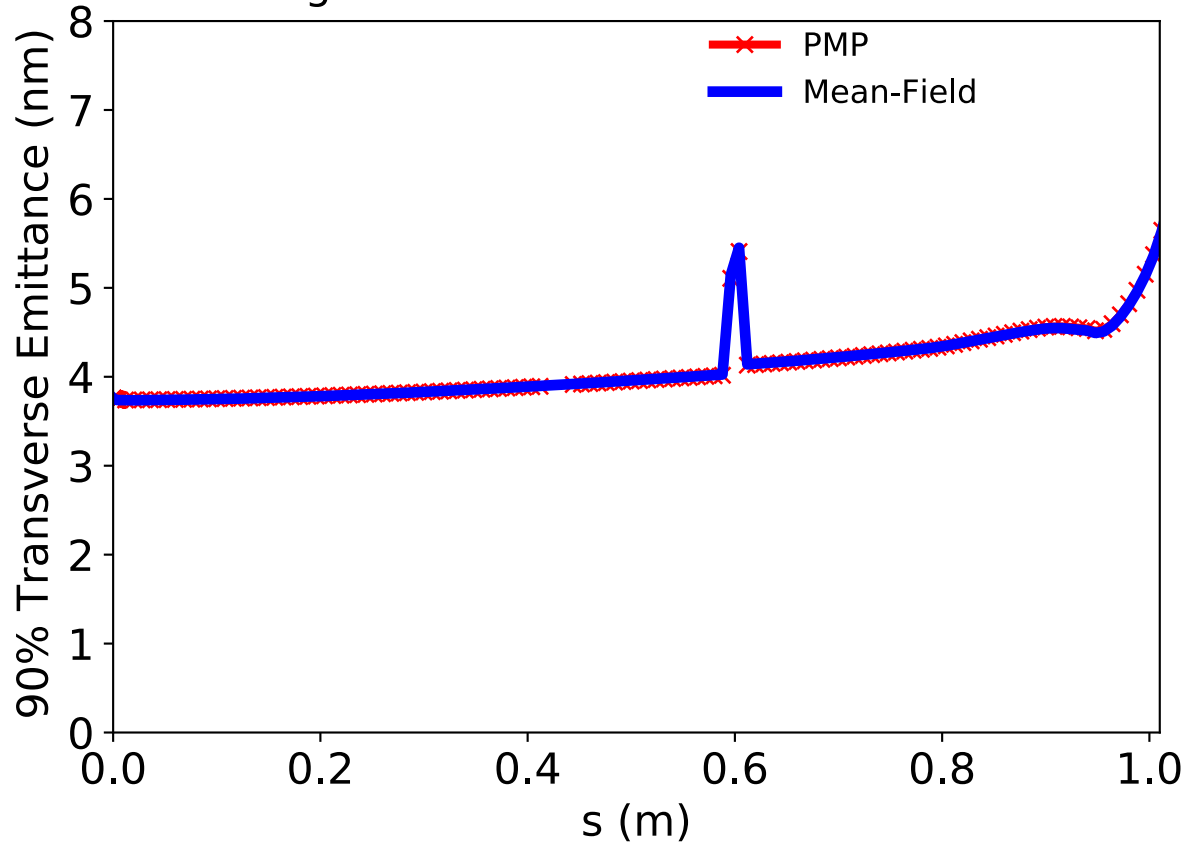
- Simulating these interactions is computational intensive and should be avoided when possible
- Calculate order of magnitude of effect
- If near the order of magnitude of MTE, consider instantaneous heating approximation
- Macroparticle extrapolation method (only for the brave)



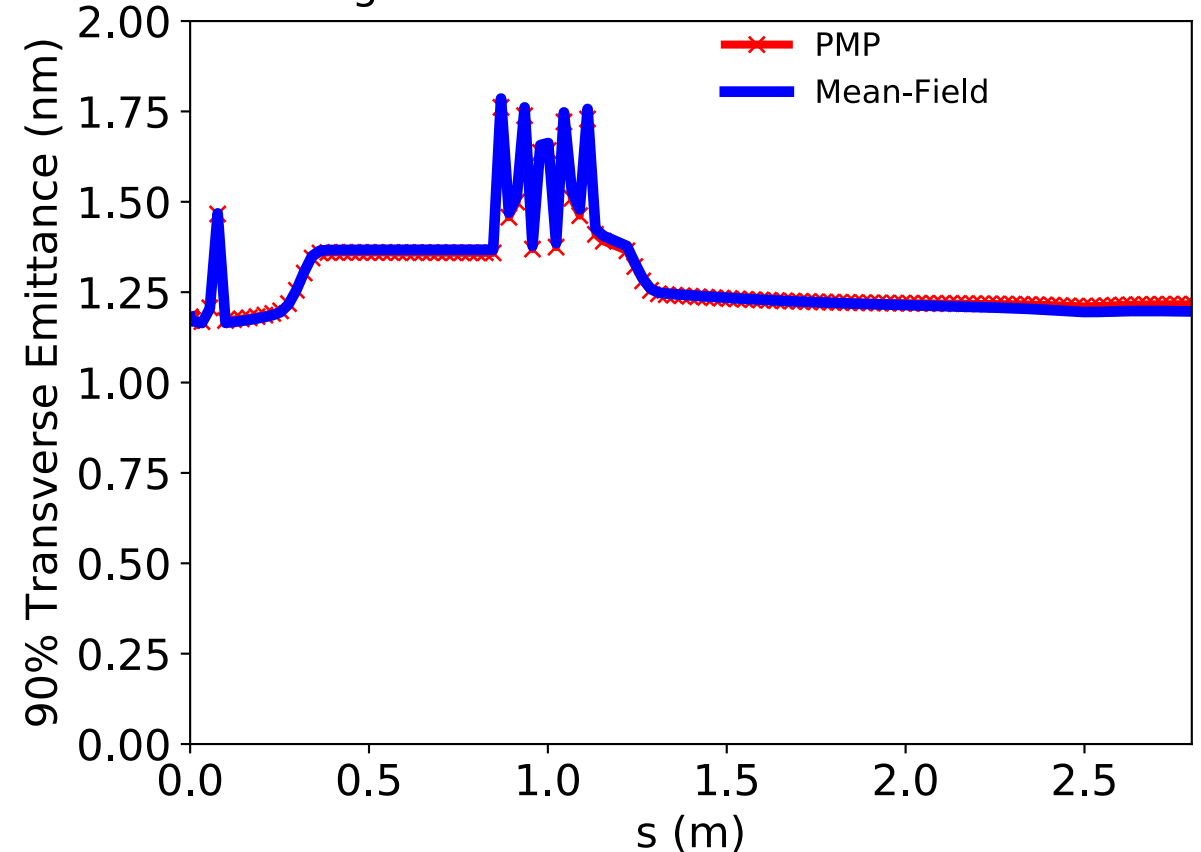
Large MTE Comparison

- With a 150 meV MTE the simulations are identical as expected
- Noticeable changes occur only below ~ 30 meV for these densities 10^{17} - 10^{18}

dc gun beamline MTE = 150 meV

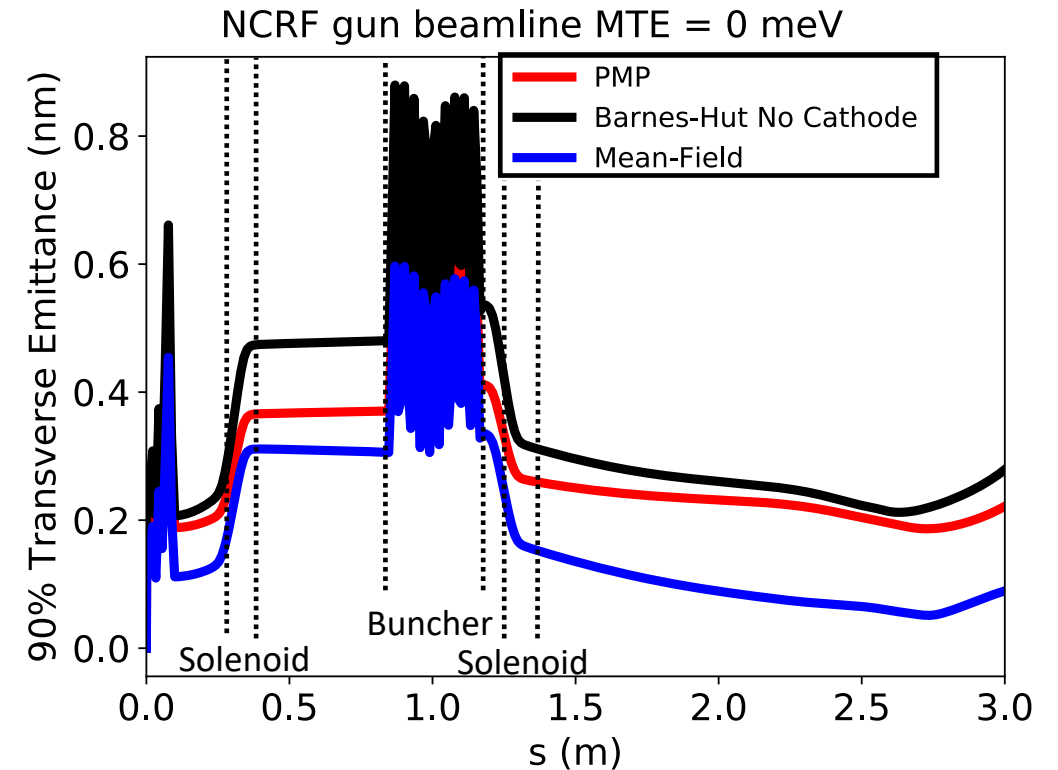


NCRF gun beamline MTE = 150 meV



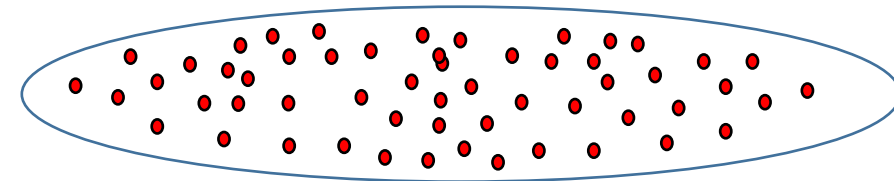
Instantaneous Heating Approximation

- Ignore finite size effects
- The heating is isotropic
- If the heating is “quick”, $\frac{2}{3}$ of E_{DIH} can be added as an effective MTE
- Approximations aren't great, but the results are good
- ~80% of the RMS emittance growth can be explained in this way

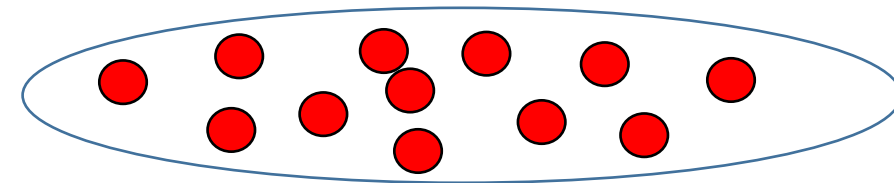


Simulating Systems with a lot more Particles

- What if you wanted to do the full P2P simulation anyway, but you have far too many particles to track them all
- People often use macroparticles to speed up simulations
- Point-to-point effects are number density dependent
- If you use macroparticles, you will **overestimate** DIH
- But it still can be useful



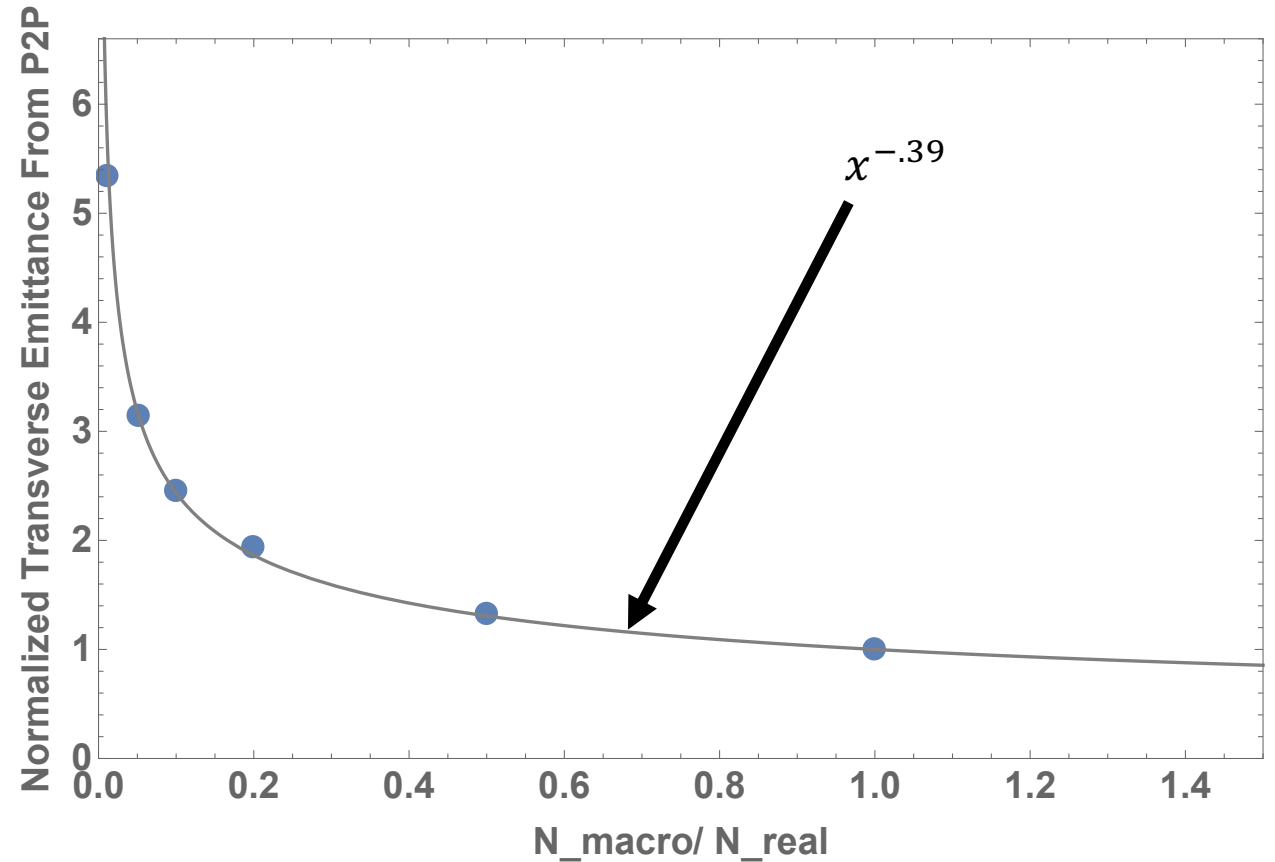
Full Simulation



Simulation With Macroparticles

Point-to-Point Macroparticle Extrapolation

- The effect of macroparticles is the inverse of increasing the actual charge density
- Run a few simulations with different macroparticle numbers
- Extract out the DIH density dependence and determine impact on emittance for full simulation



Summary

- Interactions between large numbers of particles cannot be computed exactly
- First results on simulating cold photoelectron beamlines with a non-mean-field electron interaction
- Beam quality decreases significantly, up to and including the core
- Heating effect consistent with disorder induced heating
- Ways to include point-to-point effects without full simulation

Questions?

Low Mean Transverse Energy(MTE)

- MTE is the transverse momentum spread of a particle bunch
- When the momentum spread is large enough, the electron beam acts as a liquid, and the mobility of the charges screen the effect of local density fluctuations in the beam

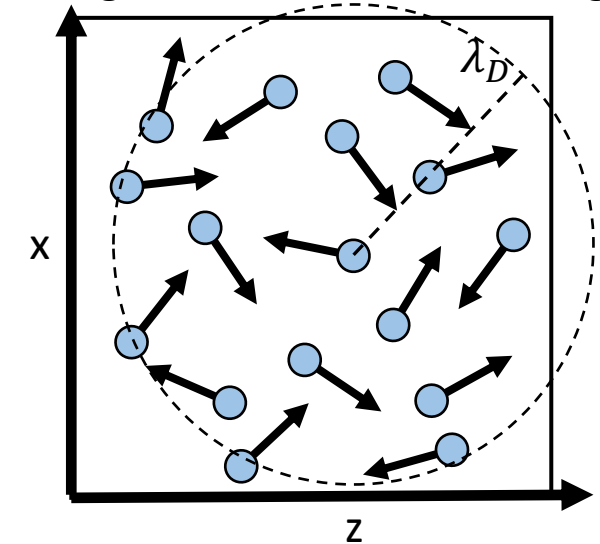
- Screened electric potential: $\varphi(r) = \frac{q}{4\pi\epsilon_0 r} e^{-\frac{r}{\lambda_D}}$

- A low MTE leads to a small Debye screening length:

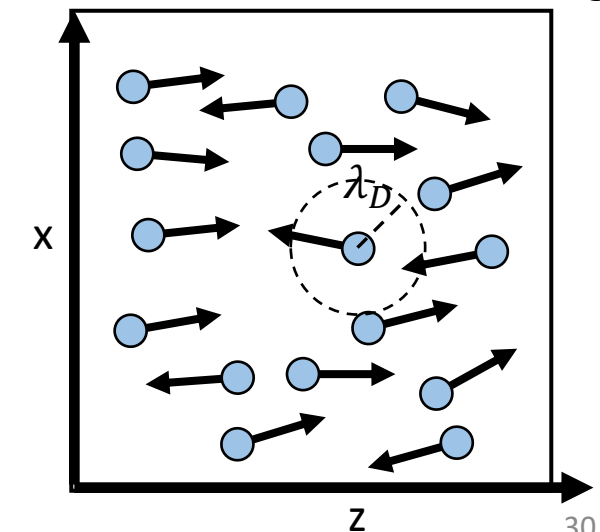
$$\lambda_D = \sqrt{\frac{\epsilon_0 k_b T}{\rho e^2}}$$

- We are at a point where λ_D is less than the average inter particle spacing(IPS) ($\lambda_D \sim .5 \mu\text{m}$, IPS $\sim 1\mu\text{m}$)
- Thus we have reached the breaking point of this approximation

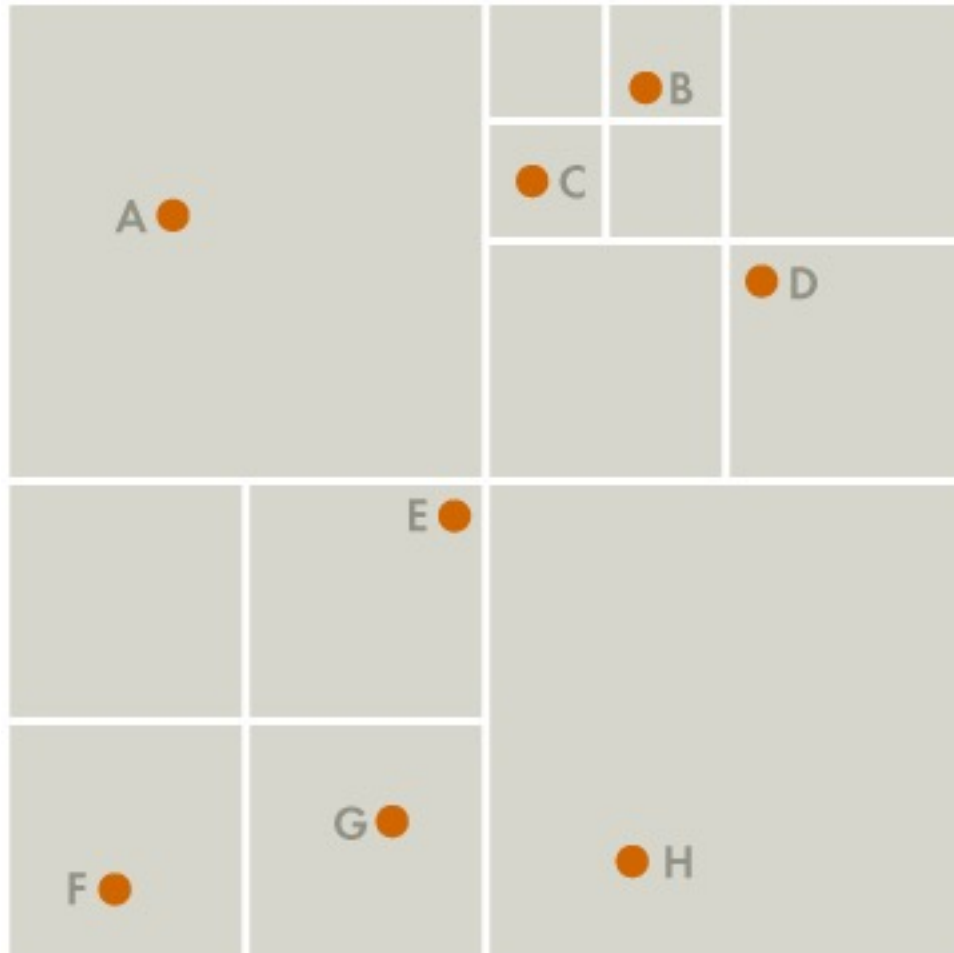
Large Mean Transverse Energy



Small Mean Transverse Energy



Barnes Hut Algorithm



Making a Barnes Hut Tree

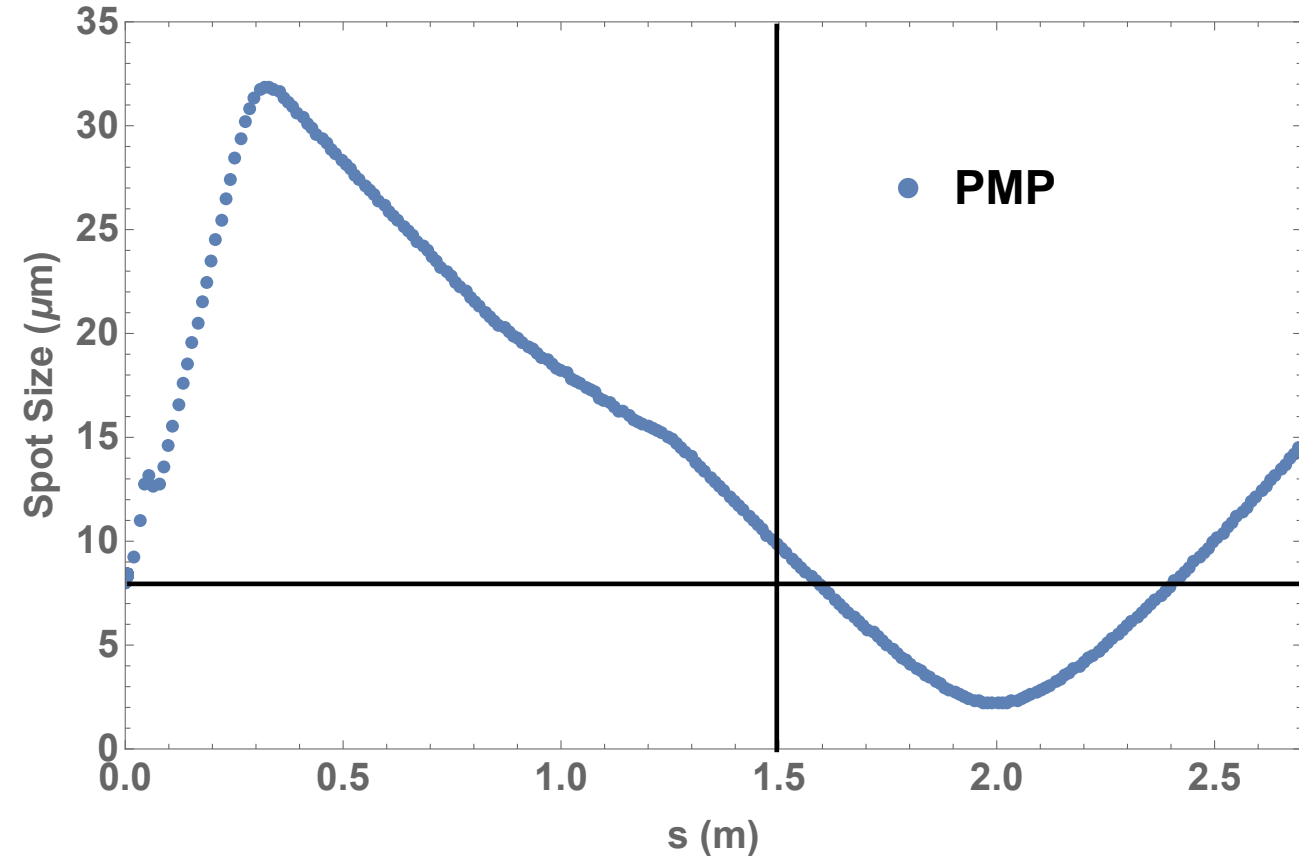
- 1) Divide 3D Space into Octants
- 2) For Each Octant
 - Store center of mass charge
 - If (Octant contains < 2 particles) Stop
 - Else Bring Octant to Step 1

Calculating Forces

- For Each Particle
- 1) Take ratio of distance from particle to center of mass charge to size of whole space
 - 2) If larger than user specified value calculate force
Else Repeat for 8 octants

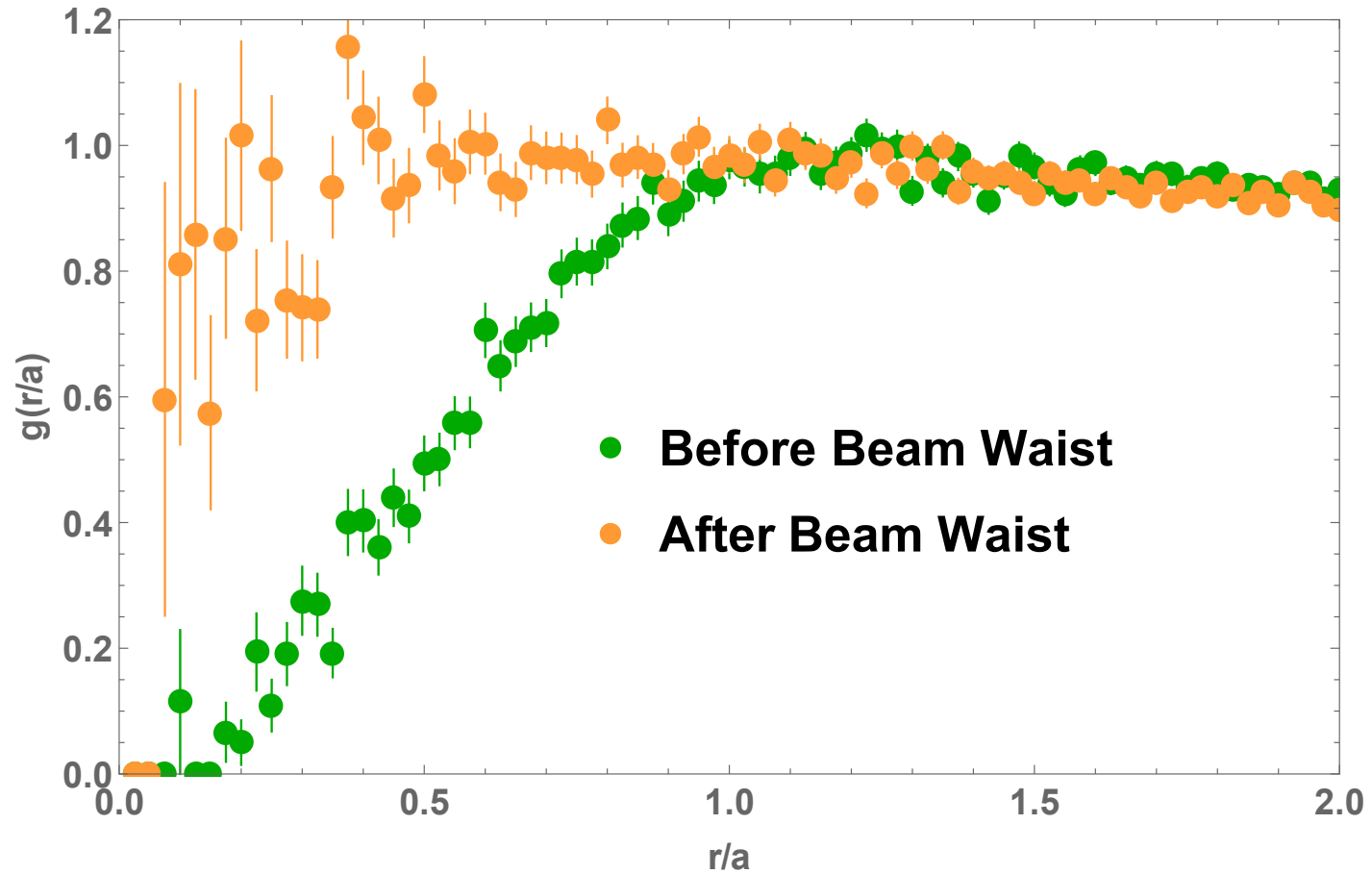
Going Through a Focus

- E_{DIH} decreases when the bunch becomes smaller than its original size
- The warm beam can fill in the original Coulomb hole
- $g(r)$ can no longer be used this way to calculate E_{DIH}



*This plot comes from a simulation at a lower particle density than the plots shown before

$g(r)$ through a Focus



DIH Scaling

- Plasma frequency: $\omega_p = \sqrt{\frac{n_e e^2}{m \epsilon_0}}$
- Heating: $E_{DIH} [meV] = 1.04 * 10^{-6} (n_0 [m^{-3}])^{1/3}$

Theoretical Scaling with Macroparticles

- DIH Kinetic energy per particle approx. potential
- $k_b T \propto \frac{e^2}{r}$
- For a fixed number of real particles N in a volume V and a varying number of macroparticles N_m :

$$e \propto N_m^{-1}, m \propto N_m^{-1}, r \propto N_m^{-\frac{1}{3}}$$

- $k_b T \propto N_m^{-\frac{5}{3}}$

- $\epsilon \propto \sqrt{\frac{k_b T}{mc^2}}$

- $\epsilon \propto N_m^{-1/3}$