## Radiative Decays of the $\Upsilon(1 \mathrm{~S})$ to a Pair of Charged Hadrons*

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(Dated: June 19, 2005)


#### Abstract

Using data obtained with the CLEO III detector, running at the Cornell Electron Storage Ring (CESR), we report on a new study of exclusive radiative $\Upsilon(1 S)$ decays into the final states $\gamma \pi^{+} \pi^{-}$, $\gamma K^{+} K^{-}$, and $\gamma p \bar{p}$. We present branching ratio measurements for the decay modes $\Upsilon(1 \mathrm{~S}) \rightarrow$ $\gamma f_{2}(1270), \Upsilon(1 \mathrm{~S}) \rightarrow \gamma f_{2}^{\prime}(1525)$, and $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma K^{+} K^{-}$; helicity production ratios for $f_{2}(1270)$ and $f_{2}^{\prime}(1525)$; upper limits for the decay $\Upsilon(1 S) \rightarrow \gamma f_{J}(2200)$, with $f_{J}(2220) \rightarrow \pi^{+} \pi^{-}, K^{+} K^{-}, p \bar{p}$; and an upper limit for the decay $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma X(1860)$, with $X(1860) \rightarrow \gamma p \bar{p}$.


[^0]
## I. INTRODUCTION

Radiative decays of heavy-quarkonia, where a photon replaces one of the three gluons from the strong decay of, for example, the $J / \psi$ or $\Upsilon(1 S)$, are useful in studying color-singlet two-gluon systems. The two gluons can, among other things, hadronize into a meson ${ }^{1}$, or directly form a glueball ${ }^{2}$. Further information on radiative decays of heavy-quarkonia can be found in [39].

Light meson production in $J / \psi$ two-body radiative decays has been experimentally well established at the $10^{-3}$ level, based largely on evidence provided by radiative decays to a pair of hadrons ${ }^{3}$. Helicity production ratio measurements have been made for the tensor mesons $f_{2}(1270)$ [41-44] and $f_{2}^{\prime}(1525)[45,46]$ in $J / \psi$ two-body radiative decays and agree with theoretical predictions [47, 48]. In 1996, the BES collaboration claimed to observe the $f_{J}(2220)$ in $J / \psi$ two-body radiative decays, and measured product branching fractions, $\mathcal{B}\left(J / \psi \rightarrow \gamma f_{J}(2220)\right) \times \mathcal{B}\left(f_{J}(2220) \rightarrow h^{+} h^{-}\right.$) (we use the convention $\left.h=\pi, K, p\right)$, of the order of $10^{-5}$ [49]. Much excitement was generated at the time because it is possible to interpret the $f_{J}(2220)$ as a glueball. A candidate similar to $f_{J}(2220)$ was reported in 1986 by the Mark III collaboration in the $K \bar{K}$ mode [50], but was not confirmed by the DM2 collaboration [51]. Recently, BES has claimed to observe the signal candidate $X(1860)$ via $J / \psi \rightarrow \gamma X(1860) \rightarrow p \bar{p}$ [52], a result that is currently being interpreted [53-59].

The experimental observation of radiative $\Upsilon(1 \mathrm{~S})$ decays is challenging because their rate is suppressed by a factor of the order of,

$$
\left(\frac{q_{b}}{q_{c}}\right)^{2}\left(\frac{m_{c}}{m_{b}}\right)^{2} \approx 0.025
$$

with respect to the rate of $J / \psi$ radiative decays. This factor arises because the quarkphoton coupling is proportional to the electric charge, and the quark propagator is roughly proportional to $1 / \mathrm{m}$ for low momentum quarks. Taking into account the total widths [40] of $J / \psi$ and $\Upsilon(1 \mathrm{~S})$, the branching fraction of a particular $\Upsilon(1 \mathrm{~S})$ radiative decay mode is expected to be suppressed by a factor of roughly 0.04. In 1999 CLEO II made the first observation of a radiative $\Upsilon(1 S)$ decay to a pair of hadrons [60], which was consistent with $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma f_{2}(1270)$, where $f_{2}(1270) \rightarrow \pi \pi$. Comparing the measured branching fraction to the $J / \psi \rightarrow \gamma f_{2}(1270)$ branching fraction, a suppression factor of $0.06 \pm 0.03$ was obtained. Recent theoretical works [2, 3], predict a suppression factor between $0.06-0.18$ for this mode, and favor the production of $f_{2}(1270)$ in a helicity 0 state. After the BES result for

[^1]the $f_{J}(2220)$ in radiative $J / \psi$ decays, a corresponding search was performed by CLEO II in the radiative $\Upsilon(1 \mathrm{~S})$ system [61] and limits where put on some of the glueball candidate's product branching ratios.

In this paper, we use a new CLEO III $\Upsilon(1 S)$ data sample, which has fifteen times higher statistics and better particle identification than the CLEO II data sample, to probe the color-singlet two-gluon spectrum by measuring the system's invariant mass using its decays to $\pi^{+} \pi^{-}, K^{+} K^{-}$, and $p \bar{p}$. Further details of this analysis can be found elsewhere [62].

## II. CLEO III DETECTOR, DATA, AND MONTE CARLO SIMULATED SAMPLE

The CLEO III detector is a versatile multi-purpose particle detector described more fully in [63]. It is centered on the interaction region of CESR. From the $e^{+} e^{-}$interaction region radially outward it consists of a silicon strip vertex detector and a wire drift chamber used to measure the position, momenta, and ionization energy losses $(d E / d x)$ of charged tracks based on their fitted path in a 1.5 T solenoidal magnetic field and the amount of charge deposited on the drift chamber wires. The silicon vertex detector and drift chamber tracking system achieves a charged particle momentum resolution of $0.35 \%(1 \%)$ at $1 \mathrm{GeV} / \mathrm{c}(5 \mathrm{GeV} / \mathrm{c})$ and a $d E / d x$ resolution of $6 \%$. Beyond the drift chamber is a Ring Imaging Cherenkov Detector, RICH, which covers $80 \%$ of the solid angle and is used to further identify charged particles by giving for each mass hypothesis the likelihood of a fit to the Cherenkov radiation pattern. After the RICH is a Crystal Calorimeter (CC) that covers $93 \%$ of the solid angle. The CC has a resolution of $2.2 \%(1.5 \%)$ for $1 \mathrm{GeV}(5 \mathrm{GeV})$ photons. After the CC is a superconducting solenoid coil that provides the magnetic field, followed by iron flux return plates with wire chambers interspersed in three layers at 3,5 , and 7 hadronic interaction lengths to provide muon identification.

The data sample has an integrated luminosity of $1.13 \mathrm{fb}^{-1}$ taken at the $\Upsilon(1 \mathrm{~S})$ energy, $\sqrt{s}=9.46 \mathrm{GeV}$, which correspond to $21.2 \pm 0.2$ million $\Upsilon(1 \mathrm{~S})$ decays [64] and $3.49 \mathrm{fb}^{-1}$ taken at the $\Upsilon(4 \mathrm{~S})$ energy, $\sqrt{s}=10.56 \mathrm{GeV}$, used to model the underlying continuum present in the $\Upsilon(1 \mathrm{~S})$ data sample. The continuum background modeling is important because continuum background processes such as $e^{+} e^{-} \rightarrow \gamma \rho$ with $\rho \rightarrow \pi^{+} \pi^{-}, e^{+} e^{-} \rightarrow \gamma \phi$ with $\phi \rightarrow K^{+} K^{-}$, and direct $e^{+} e^{-} \rightarrow \gamma h^{+} h^{-}$, have the same topology as the signal events we are investigating.

Efficiencies are evaluated using a Monte Carlo simulation of the process [65] and a GEANT-based [66] detector response. Monte Carlo samples of $e^{+} e^{-} \rightarrow \gamma X$ with $X \rightarrow h^{+} h^{-}$ are generated [67] at both the $\Upsilon(1 S)$ and $\Upsilon(4 \mathrm{~S})$ energies with uniform angular distributions and flat $h^{+} h^{-}$invariant mass distributions from threshold to $3.5 \mathrm{GeV} / \mathrm{c}^{2}$.

## III. EVENT SELECTION

Events which satisfy the CLEO III trigger [68] are required to meet the following analysis requirements: (a) There are exactly two charged tracks consistent with coming from the beamspot, with $d E / d x$ information, and with good quality track fits. (b) There is exactly one CC shower that is unmatched to any track and whose energy, $E_{\gamma}$, is greater than 4 GeV .

Each event is also required to be consistent with having the 4 -momentum of the initial $e^{+} e^{-}$system by demanding that the chi-squared from a kinematic fit to the following constraint,

$$
\begin{equation*}
\vec{p}_{h^{+} h^{-}}+\left(2 E_{b e a m}-E_{h^{+} h^{-}}\right) \widehat{p}_{\gamma}=\vec{p}_{C M} \tag{1}
\end{equation*}
$$

be less than 100, where $\vec{p}_{h^{+} h^{-}}$is the di-hadron momentum, $E_{h^{+} h^{-}}$is the di-hadron energy, $E_{\text {beam }}$ is the beam energy, $\widehat{p}_{\gamma}$ is the photon's direction, and $\vec{p}_{C M}$ is the momentum of the $e^{+} e^{-}$ system (which has a magnitude of a few $\mathrm{MeV} / \mathrm{c}$ because of the small but finite crossing angle of the $e^{+}$and $e^{-}$beams). Equation 1 is a 3 -constraint subset of the 4 -momentum constraint and has the convenient property of avoiding the use of the measured photon's energy, which has an asymmetric measurement error. We improve the measurement of the di-hadron 4momenta (the di-hadron invariant mass resolution becomes $3.2,2.6$, and $2.0 \mathrm{MeV} / \mathrm{c}^{2}$ for the pion, kaon, and proton modes, respectively) by using the constraint in equation 1 , and then demand that:

$$
0.950<\left(E_{h^{+} h^{-}}+E_{\gamma}\right) / 2 E_{\text {beam }}<1.025
$$

Strong electron and muon vetoes are imposed to suppress the abundant QED processes $e^{+} e^{-} \rightarrow \gamma e^{+} e^{-}$and $e^{+} e^{-} \rightarrow \gamma \mu^{+} \mu^{-}$. To reject $e^{+} e^{-} \rightarrow \gamma e^{+} e^{-}$, we require each track to have a matched CC shower with an energy $E$, together with a measured momentum $p$, such that $|E / p-0.95|>0.1$, and that the combined RICH and $d E / d x$ likelihood for $h$ be higher than the combined likelihood for $e$. To reject $e^{+} e^{-} \rightarrow \gamma \mu^{+} \mu^{-}$, we require that neither track produce a signal in the five hadronic interaction lengths of the muon system. For the $\pi^{+} \pi^{-}$mode, where muon background is a particular problem because of the similar pion and muon masses, we further require that both tracks must be within the barrel part of the muon chambers $(|\cos \theta|<0.7)$, and both have $p>1 \mathrm{GeV} / \mathrm{c}$. To increase the solidangle acceptance of the detector and improve the overall muon suppression efficiency with virtually no increase in muon fakes, we flag an event as "not muonic" and remove the muon suppression requirements if either track deposits more than 600 MeV in the CC .

Events that satisfy all the above requirements are then identified as either $\pi^{+} \pi^{-}, K^{+} K^{-}$, or $p \bar{p}$ using the RICH and $d E / d x$ information. Since the ratios $\pi^{+} \pi^{-} / K^{+} K^{-}$and $K^{+} K^{-} / p \bar{p}$ are much larger than 1 for these types of events, in the 3 cases where we try to reduce the background from a lower-mass hadron (e.g. $\pi$ faking $K$ ), we also use the chi-squared value from the kinematic constraint in Equation 1 to identify the event type. Since the constraint involves the di-hadron energy, the chi-squared value is sensitive to the hadronic masses. After these procedures, the particle identification and fake rates are $90 \%$ ( $0.31 \%$ ), $99 \%$ ( $0.03 \%$ ), $98 \%$ ( $0.10 \%$ ) for kaons (pions), protons (pions), and protons (kaons), respectively.

## IV. DETERMINATION OF SIGNALS AND THEIR SPIN ASSIGNMENTS

The overall reconstruction efficiencies as determined by Monte Carlo simulations, including both event selection and analysis cuts, are $43 \%, 48 \%$, and $56 \%$ for the $\Upsilon(1 S)$ radiative decays to $\pi^{+} \pi^{-}, K^{+} K^{-}$, and $p \bar{p}$, respectively. These efficiencies are only mildly dependent on the di-hadron invariant mass and are very similar for the continuum background events. The continuum subtracted di-hadron invariant mass plots are obtained by efficiency correcting each bin of the di-hadron invariant mass plots for the $\Upsilon(1 \mathrm{~S})$ and $\sqrt{s}=10.56 \mathrm{GeV}$ datasets, scaling the latter plot by a factor of $0.404 \pm 0.002^{4}$, and subtracting it from the $\Upsilon(1 \mathrm{~S})$ dataset invariant mass plot. Possible signals are determined by fitting each spectrum to spin-dependent relativistic Breit-Wigner functions ${ }^{5}$. The spin value for each Breit-Wigner

[^2]

FIG. 1: Invariant mass of $\pi^{+} \pi^{-}$from $e^{+} e^{-} \rightarrow \gamma \pi^{+} \pi^{-}$for the scaled $\sqrt{s}=10.56 \mathrm{GeV}$ dataset (solid line), and the $\Upsilon(1 \mathrm{~S})$ dataset (crosses). The large number of events near $770 \mathrm{MeV} / \mathrm{c}^{2}$ is due to the abundant process $e^{+} e^{-} \rightarrow \gamma \rho$.
is surmised by identifying each possible resonance in the invariant mass plot based on its approximate mass and width. Later, we confirm these spin assignments for the significant resonances by inspecting the angular distributions of the $\Upsilon(1 S)$ decay products.

The $\pi^{+} \pi^{-}$invariant mass plots for the $\Upsilon(1 \mathrm{~S})$ and the scaled $\sqrt{s}=10.56 \mathrm{GeV}$ datasets are shown in Figure 1. The fit to the continuum subtracted $\pi^{+} \pi^{-}$spectrum, shown in Figure 2, has a significant $f_{2}(1270)$ signal of $944 \pm 74$ events. It also has two less significant signal candidates; $340_{-130}^{+140}$ events in the $f_{0}(980)$ region, and $80 \pm 30$ events in the $f_{4}(2050)$ region (see Figure 3) whose significances are $4.3 \sigma$ and $2.6 \sigma$, respectively. Each significance
each dataset,

$$
f=0.404 \approx \frac{1.13 \mathrm{fb}^{-1}}{3.49 \mathrm{fb}^{-1}}\left(\frac{10.56 \mathrm{GeV}}{9.46 \mathrm{GeV}}\right)^{2}
$$

${ }^{5}$ The spin-dependent relativistic Breit-Wigner parameterization used has the following probability distribution for a particular $h^{+} h^{-}$invariant mass $x>x_{0}$,

$$
d P(x) \propto \frac{x x_{m} \Gamma(x)}{\left(x^{2}-x_{m}^{2}\right)^{2}+\left(x_{m} \Gamma(x)\right)^{2}} d x
$$

where:

$$
\Gamma(x)=\Gamma_{0}\left(\frac{x-x_{0}}{x_{m}-x_{0}}\right)^{2 S+1} \frac{2\left(x_{m}-x_{0}\right)^{2}}{\left(x-x_{0}\right)^{2}+\left(x_{m}-x_{0}\right)^{2}} .
$$

In the above expression, $x_{m}$ and $\Gamma_{0}$ represent respectively the most likely mass and the width and are allowed to float during the fit. The values of $x_{0}$ and $S$ are fixed during the fit to the invariant mass threshold for the particular mode and the spin of the resonance, respectively. The number of events for each fitted signal candidate is obtained by integrating this Breit-Wigner parameterization between threshold and $3 \mathrm{GeV} / \mathrm{c}^{2}$.


FIG. 2: Invariant mass of $\pi^{+} \pi^{-}$from $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma \pi^{+} \pi^{-}$and the fit to the three spin-dependent relativistic Breit-Wigner functions described in the text.
is obtained by doing multiple chi-squared fits to the invariant mass plot fixing the signal area to different values, assigning each of these multiple fits a probability proportional to $e^{-\chi^{2} / 2}$, normalizing the resulting probability distribution, and calculating the probability for negative or 0 signal. The $K^{+} K^{-}$invariant mass plots for the $\Upsilon(1 S)$ and the scaled $\sqrt{s}=10.56 \mathrm{GeV}$ datasets are shown in Figure 4. The fit to the continuum subtracted $K^{+} K^{-}$spectrum, shown in Figure 5, has a significant signal of $312_{-61}^{+69}$ events identified as the $f_{2}^{\prime}(1525)$, and two non-significant signal candidates indicating possible $f_{2}(1270)$ and $f_{0}(1710)$ production with $109 \pm 36$ and $73 \pm 29$ events whose significances are $3.2 \sigma$ and $3.3 \sigma$, respectively. The excess in the $f_{2}(1270)$ region is consistent with that expected using the $\gamma \pi^{+} \pi^{-}$data and the known branching ratios for the $f_{2}(1270)$. We also note that there is a significant excess of $220 \pm 20$ events above 1.9 GeV in the $K^{+} K^{-}$invariant mass distribution which is not associated with any resonant structure. The $p \bar{p}$ invariant mass plots for the $\Upsilon(1 \mathrm{~S})$ and the scaled $\sqrt{s}=10.56 \mathrm{GeV}$ datasets are shown in Figure 6. No recognizable structure is seen in the continuum subtracted $p \bar{p}$ spectrum, which is shown in Figure 7, and has an excess of $85 \pm 18$ events in the $2-3 \mathrm{GeV} / \mathrm{c}^{2}$ invariant mass region. In particular, we do not note a significant enhancement near threshold.

To confirm the spins of our $f_{2}(1270) \rightarrow \pi^{+} \pi^{-}$and $f_{2}^{\prime}(1525) \rightarrow K^{+} K^{-}$signals, we examine the absolute value of the cosine of the polar angle of the photon with respect to the beam axis, $\left|\cos \theta_{\gamma}\right|$, and the absolute value of the cosine of the angle formed by the 3 -momentum vector of one of the hadrons measured in the di-hadron rest frame with the photon's direction, $\left|\cos \theta_{h}\right|$. The event selection efficiency is slightly dependent on both angles, so to minimize systematic effects, the $\left|\cos \theta_{\gamma}\right|$ and $\left|\cos \theta_{h}\right|$ efficiency-corrected distributions are obtained by projecting the 2-dimensional bin-by-bin efficiency-corrected $\left|\cos \theta_{\gamma}\right|-\left|\cos \theta_{h}\right|$ distribution. We also subtract the background contributions from the tails of nearby resonances. The resulting angular distributions (shown in Figures 8 and 9) are simultaneously fit to the helicity formalism prediction $[48,62,69]$ for different resonance spin hypotheses up to $J=4$. For the $f_{2}(1270)$ the different fit confidence levels are $8 \times 10^{-19}, 2 \times 10^{-19}, 0.05,8 \times 10^{-12}$,


FIG. 3: Invariant mass of $\pi^{+} \pi^{-}$from $\Upsilon(1 S) \rightarrow \gamma \pi^{+} \pi^{-}$and the fit to the $f_{4}(2050)$ candidate in the region $1.5-3.0 \mathrm{GeV} / \mathrm{c}^{2}$.


FIG. 4: Invariant mass of $K^{+} K^{-}$from $e^{+} e^{-} \rightarrow \gamma K^{+} K^{-}$for the scaled $\sqrt{s}=10.56 \mathrm{GeV}$ dataset (solid line), and the $\Upsilon(1 \mathrm{~S})$ dataset (crosses). The large number of events near $1.050 \mathrm{GeV} / \mathrm{c}^{2}$ is due to the abundant process $e^{+} e^{-} \rightarrow \gamma \phi$.
and $1 \times 10^{-12}$ for the hypotheses $J=0,1,2,3,4$, respectively. For the $f_{2}^{\prime}(1525)$ the different fit confidence levels are $2 \times 10^{-4}, 2 \times 10^{-4}, 0.23,8 \times 10^{-3}$, and $2 \times 10^{-3}$ for the hypotheses $J=0,1,2,3,4$, respectively. These results confirm our identification of the resonances, as


FIG. 5: Invariant mass of $K^{+} K^{-}$from $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma K^{+} K^{-}$and the fit to the three spin-dependent relativistic Breit-Wigner functions described in the text.


FIG. 6: Invariant mass of $p \bar{p}$ from $e^{+} e^{-} \rightarrow \gamma p \bar{p}$ for the scaled $\sqrt{s}=10.56 \mathrm{GeV}$ dataset (solid line), and the $\Upsilon(1 \mathrm{~S})$ dataset (crosses). The events near $3.1 \mathrm{GeV} / \mathrm{c}^{2}$ are due to the process $e^{+} e^{-} \rightarrow \gamma J / \psi$.


FIG. 7: Invariant mass of $p \bar{p}$ from $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma p \bar{p}$.
in both cases the angular distributions of the data strongly favor the $J=2$ hypothesis $^{6}$,

$$
\begin{align*}
\frac{d P_{\theta_{h}, \theta_{\gamma}}}{d \cos \theta_{h} d \cos \theta_{\gamma}}= & \left|a_{0}\right|^{2} \times \frac{5}{8}\left(3 \cos ^{2} \theta_{h}-1\right)^{2} \times \frac{3}{8}\left(1+\cos ^{2} \theta_{\gamma}\right)+ \\
& \left|a_{1}\right|^{2} \times \frac{15}{16} \sin ^{2} 2 \theta_{h} \times \frac{3}{4} \sin ^{2} \theta_{\gamma}+  \tag{2}\\
& \left|a_{2}\right|^{2} \times \frac{15}{16} \sin ^{4} \theta_{h} \times \frac{3}{8}\left(1+\cos ^{2} \theta_{\gamma}\right)
\end{align*}
$$

where $a_{\lambda}, \lambda=0,1,2$, are the normalized helicity amplitudes, $\int d P_{\theta_{h}, \theta_{\gamma}}=\left|a_{0}\right|^{2}+\left|a_{1}\right|^{2}+$ $\left|a_{2}\right|^{2}=1$. In other words, $\left|a_{\lambda}\right|^{2}$ is the probability of $X$ in $\Upsilon(1 S) \rightarrow \gamma X$ to have helicity $\pm \lambda$. Because of the normalization condition, the $\theta_{h}-\theta_{\gamma}$ probability distribution can be described by two free parameters, traditionally chosen to be the helicity production ratios,

$$
\mathrm{x}^{2}=\frac{\left|a_{1}\right|^{2}}{\left|a_{0}\right|^{2}}, \text { and } \mathrm{y}^{2}=\frac{\left|a_{2}\right|^{2}}{\left|a_{0}\right|^{2}} .
$$

To measure $\mathrm{x}^{2}$ and $\mathrm{y}^{2}$, we simultaneously fit the data to the individual $\theta_{h}$ and $\theta_{\gamma}$ distribu-

[^3]

FIG. 8: Distributions of $\left|\cos \theta_{\pi}\right|$ (top) and $\left|\cos \theta_{\gamma}\right|$ (bottom) for the signal events in the $f_{2}(1270)$ invariant mass region. The solid lines correspond to a simultaneous fit to the $J=2$ helicity formalism prediction (Equation 3).
tions ${ }^{7}$,

$$
\begin{align*}
\frac{d N_{\theta_{\gamma}}}{d \cos \theta_{\gamma}} & =N \int_{\theta_{h}} d P_{\theta_{h}, \theta_{\gamma}} \\
& =\frac{N}{1+\mathrm{x}^{2}+\mathrm{y}^{2}}\left[\mathrm{x}^{2} \frac{3}{4} \sin ^{2} \theta_{h}+\left(1+\mathrm{y}^{2}\right) \frac{3}{8}\left(1+\cos ^{2} \theta_{\gamma}\right)\right] \\
\frac{d N_{\theta_{h}}}{d \cos \theta_{h}} & =N \int_{\theta_{\gamma}} d P_{\theta_{h}, \theta_{\gamma}}  \tag{3}\\
& =\frac{N}{1+\mathrm{x}^{2}+\mathrm{y}^{2}}\left[\frac{5}{8}\left(3 \cos ^{2} \theta_{h}-1\right)^{2}+\mathrm{x}^{2} \frac{15}{16} \sin ^{2} 2 \theta_{h}+\mathrm{y}^{2} \frac{15}{16} \sin ^{4} \theta_{h}\right]
\end{align*}
$$

Where $N$ corresponds to the number of events. Using the fits to the data (see Figures 8 and 9 ) we measure the following helicity production ratios,

$$
\begin{array}{ll}
\mathrm{x}_{f_{2}(1270)}^{2}=0.00_{-0.00}^{+0.02}, & \mathrm{y}_{f_{2}(1270)}^{2}=0.09_{-0.07}^{+0.08}, \\
\mathrm{x}_{f_{2}^{\prime}(1525)}^{2}=0.00_{-0.00}^{+0.10}, & \mathrm{y}_{f_{2}^{\prime}(1525)}^{2}=0.30_{-0.17}^{+0.22} .
\end{array}
$$

Where only statistical errors are included. These results indicate that both resonances are predominantly produced with helicity 0 . They are in agreement with the predictions of [2], and in good agreement with the twist-two order predictions of [3]: no $\lambda=1$ production, and $\lambda=2$ production suppressed by a factor of $\left(m_{X} / m_{b}\right)^{2}$ with respect to $\lambda=0$ production, where $m_{X}$ is the mass of the tensor meson and $m_{b}$ is the mass of the $b$ quark.

[^4]

FIG. 9: Distributions of $\left|\cos \theta_{K}\right|$ (top) and $\left|\cos \theta_{\gamma}\right|$ (bottom) for the signal events in the $f_{2}^{\prime}(1525)$ invariant mass region. The solid lines correspond to a simultaneous fit to the $J=2$ helicity formalism prediction (Equation 3).

We use the results from fitting the angular distributions to correct the Monte Carlo simulation efficiencies, which are calculated using flat distributions in the relevant angles, by a factor of $0.78 \pm 0.02$ for the $f_{2}(1270), 0.90 \pm 0.01$ for the $f_{2}^{\prime}(1525)$, and $0.88_{-0.01}^{+0.03}$ for the significant excess in the $2-3 \mathrm{GeV} / \mathrm{c}^{2}$ region of the di-kaon invariant mass. The large correction in the pion mode is due to the necessarily stronger muon suppression requirement. The measured branching ratios of the significant resonances are,

$$
\begin{aligned}
& \mathcal{B}\left(\Upsilon(1 S) \rightarrow \gamma f_{2}(1270)\right)=(10.2 \pm 0.8 \pm 0.7) \times 10^{-5} \\
& \mathcal{B}\left(\Upsilon(1 S) \rightarrow \gamma f_{2}^{\prime}(1525)\right)=\left(3.7_{-0.7}^{+0.9} \pm 0.8\right) \times 10^{-5},
\end{aligned}
$$

and the measured branching ratio of the significant excess of events in $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma K^{+} K^{-}$ with a di-kaon invariant mass between $2-3 \mathrm{GeV} / \mathrm{c}^{2}$ is,

$$
\mathcal{B}\left(\Upsilon(1 \mathrm{~S}) \rightarrow \gamma K^{+} K^{-}\right)=(1.14 \pm 0.08 \pm 0.10) \times 10^{-5}
$$

where the first error is statistical and the second includes the systematic uncertainty. The sources of systematic uncertainty are $1 \%$ from the number of $\Upsilon(1 \mathrm{~S})$ decays, $2 \%$ from the Monte Carlo simulation of the track reconstruction, $3 \%$ ( $8 \%$ ) from the Monte Carlo efficiency modeling of the event requirements in the pion (kaon) mode, and $1 \%$ to $3 \%$ from the uncertainty in the angular distribution measurements. We also assign a $15 \%$ systematic uncertainty from possible interference with $\gamma f_{2}(1270)$ to $\gamma f_{2}^{\prime}(1525)$, and less than a $1 \%$ systematic uncertainty from high-momentum neutral pions faking photons in the decay $\Upsilon(1 S) \rightarrow \rho \pi$ based on the upper limit in [72]. Finally, we include the uncertainties in the $f_{2}(1270)$ and $f_{2}^{\prime}(1525)$ hadronic branching ratios [40] in the systematic uncertainty. For our less significant signal candidates, the branching fraction central value, along with its significance and $90 \%$ confidence level upper limit, is shown in Table I.

TABLE I: Branching fraction central value (BF), its statistical significance, and its $90 \%$ confidence level upper limit (UL), for each signal candidate with a significance $<5 \sigma$. In the branching fraction central value, the first uncertainty is statistical and the second is systematic. The first three table entries are product branching fractions.

| Channel | BF $\times\left(10^{-5}\right)$ | Significance | UL $\times\left(10^{-5}\right)$ |
| :---: | :---: | :---: | :---: |
| $\gamma f_{0}(980) \times f_{0}(980) \rightarrow \pi^{+} \pi^{-}$ | $1.8_{-0.7}^{+0.8} \pm 0.1$ | $4.3 \sigma$ | $<3$ |
| $\gamma f_{4}(2050) \times f_{4}(2050) \rightarrow \pi^{+} \pi^{-}$ | $0.37 \pm 0.14 \pm 0.03$ | $2.6 \sigma$ | $<0.6$ |
| $\gamma f_{0}(1710) \times f_{0}(1710) \rightarrow K^{+} K^{-}$ | $0.38 \pm 0.16 \pm 0.04$ | $3.2 \sigma$ | $<0.7$ |
| $\gamma p \bar{p}, 2 \mathrm{GeV} / \mathrm{c}^{2}<m_{p \bar{p}}<3 \mathrm{GeV} / \mathrm{c}^{2}$ | $0.41 \pm 0.08 \pm 0.10$ | $4.8 \sigma$ | $<0.6$ |

## V. DETERMINATION OF UPPER LIMITS FOR $f_{J}(2220)$ AND $X(1860)$ PRODUCTION AND DECAY

To measure upper limits of the product branching ratio for the decays $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma f_{J}(2220)$ with $f_{J}(2220) \rightarrow h^{+} h^{-}$, we fit the $h^{+} h^{-}$invariant mass plots, shown in Figure 10, using a Breit-Wigner with a peak mass and width fixed at $2.234 \mathrm{GeV} / \mathrm{c}^{2}$ and $0.017 \mathrm{GeV} / \mathrm{c}^{2}$, respectively. These are the values from the possible $f_{J}(2220)$ signal reported by the BES experiment [49], which is considered a candidate for a glueball. To model the general excess of events between 2.0 and $2.5 \mathrm{GeV} / \mathrm{c}^{2}$ we also use a flat background function in the fit. Although the highest bin in the $p \bar{p}$ plot is indeed in the region of the $f_{J}(2220)$, the excess ( $12 \pm 5$ events) is not significant, and there are no significant signals anywhere in these three plots. To find upper limits for $f_{J}(2220) \rightarrow h^{+} h^{-}$decays, we fix the area of the Breit-Wigner to different values, minimize the chi-squared from the fit, and give that area a probability proportional $e^{-\chi^{2} / 2}$. These probability distributions are then used to obtain the following $90 \%$ confidence level upper limits on the product branching ratio for $f_{J}(2220)$ production and decay to each mode:

$$
\begin{aligned}
\mathcal{B}(\Upsilon(1 S) & \left.\rightarrow \gamma f_{J}(2200)\right) \times \mathcal{B}\left(f_{J}(2200)\right. \\
\mathcal{B}(\Upsilon(1 \mathrm{~S}) & \left.\rightarrow \gamma \pi^{+} \pi^{-}\right)<8 \times 10^{-7} \\
\mathcal{B}(\Upsilon(1 \mathrm{~S}) & \left.\left.\rightarrow \gamma f_{J}(2200)\right) \times \mathcal{B}\left(f_{J}(2200)\right) \times K^{+} K^{-}\right)<6 \times 10^{-7} \\
\mathcal{B}\left(f_{J}(2200)\right. & \rightarrow p \bar{p})<11 \times 10^{-7}
\end{aligned}
$$

The systematic uncertainties on the branching ratios where added in quadrature with the statistical errors in forming the above limits. Using the $X(1860)$ parameters measured in [52], and proceeding in a similar manner as described above, we obtain,

$$
\mathcal{B}(\Upsilon(1 S) \rightarrow \gamma X(1860)) \times \mathcal{B}(X(1860) \rightarrow p \bar{p})<5 \times 10^{-7}
$$

with a $90 \%$ confidence level.

## VI. SUMMARY AND CONCLUSION

We have confirmed CLEO's previous observation of the $f_{2}(1270)$ in radiative $\Upsilon(1 S)$ decays and made a new observation of the $f_{2}^{\prime}(1525)$, obtaining factors of $0.07 \pm 0.01$ and $0.08_{-0.03}^{+0.04}$ for the ratio of the $\Upsilon(1 \mathrm{~S})$ branching fraction with respect to the one measured in $J / \psi$ radiative


FIG. 10: Invariant mass of $\pi^{+} \pi^{-}$(top) $K^{+} K^{-}$(middle) and $p \bar{p}$ (bottom) from $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma h^{+} h^{-}$ with the fit that determines the $90 \%$ C.L. upper limit overlaid. The upper limit is obtained from fits to a $f_{J}(2220)$ signal Breit-Wigner function (dashed) and flat background function (solid). The flat background function is used to model the excess of events from $\Upsilon(1 \mathrm{~S}) \rightarrow \gamma h^{+} h^{-}$in this invariant mass region.
decays, respectively. These values are larger than, but the same order of magnitude as, the ratio of 0.04 expected from naive scaling arguments. The observed $f_{2}(1270)$ production is in agreement with the prediction in [3] and somewhat lower than the prediction in [2]. In both of the measured modes we can confirm by fits to the angular distributions of the photon and charged particles that the two daughter hadrons are indeed produced by a spin- 2 parent. We find that this parent is produced mostly with helicity 0 , in good agreement with the predictions in $[2,3]$. No structure is seen in the $p \bar{p}$ invariant-mass distribution. In particular, we do not observe a near-threshold enhancement as in [52]. Finally, stringent limits have been put on the production of the glueball candidate $f_{J}(2220)$ in radiative $\Upsilon(1 \mathrm{~S})$ decays. Glueball production is expected to be enhanced in $\Upsilon(1 \mathrm{~S})$ radiative decays [22, 73, 74], but we find that, within our experimental sensitivity, tensor meson states, believed to be composed only of quarks, dominate the di-gluon spectrum.

We gratefully acknowledge the effort of the CESR staff in providing us with excellent luminosity and running conditions. This work was supported by the National Science Foundation and the U.S. Department of Energy.
[1] G.T. Bodwin, E. Braaten and G.P. Lepage, Phys. Rev. D 51, 1125 (1995).
[2] S. Fleming, C. Lee and A.K. Leibovich, Phys. Rev. D 71, 074002 (2005).
[3] J.P. Ma, Nucl. Phys. B 605, 625 (2001).
[4] V.N. Baier and A.G. Grozin, Sov. J. Nucl. Phys. 35, 596 (1982).
[5] V.N. Baier and A.G. Grozin, Z. Phys. C 29, 161 (1985).
[6] T. Barnes, Z. Phys. C 10, 275 (1981).
[7] J.M. Cornwall and A. Soni, Phys. Lett. B 120, 431 (1983).
[8] W.S. Hou and G.G. Wong, Phys. Rev. D 67, 034003 (2003).
[9] C.J. Morningstar and M.J. Peardon, Phys. Rev. D 60, 034509 (1999).
[10] A. Vacarino and D. Weingarten, Phys. Rev. D 60, 114501 (1999).
[11] C. Liu, Chin. Phys. Lett. 18, 187 (2001).
[12] D.Q. Liu, J.M. Wu and Y. Chen, High Energy Phys. Nucl. Phys. 26, 222 (2002).
[13] P.G.O. Freund and Y. Nambu, Phys. Rev. Lett. 34, 1645 (1975).
[14] R.L. Jaffe and K. Johnson, Phys. Lett. B 60, 201 (1976).
[15] T. Barnes, F.E. Close and S. Monaghan, Nucl. Phys. B 198, 380 (1982).
[16] C.E. Carlson, T.H. Hanson, and C. Peterson, Phys. Rev. D 27, 1556 (1983).
[17] N. Isgur and J. Paton, Phys. Rev. D 31, 2910 (1985).
[18] S. Narison, Z. Phys. C 26, 209 (1984).; Nucl. Phys. B 509, 312 (1998); Nucl. Phys. Proc. Suppl. 96, 244 (2001).
[19] M.H. Thomas, M.Lust and H.J. Mang, J. Phys. G 18, 1125 (1992).
[20] J.Y. Cui, J.M. Wu and H.Y. Jin, Phys. Lett. B 424, 381 (1998).
[21] S.J. Brodsky, A.S. Goldhaber and J. Lee, Phys. Rev. Lett. 91, 112001 (2003).
[22] X.G. He, H.Y. Jin and J.P. Ma, Phys. Rev. D 66, 074015 (2002).
[23] M. Melis, F. Murgia and J. Parisi, Phys. Rev. D 70, 034021, (2004).
[24] J.C. Su and J.X. Chen, Phys. Rev. D 69, 076002 (2004).
[25] V.V. Anisovich et al. (Crystal Ball Collaboration), Phys. Lett. B 323, 233 (1994).
[26] V.V. Anisovich, D.V. Bugg, A.V. Sarantsev and B.S. Zou, Phys. Rev. D 50, 1972 (1994).
[27] C. Amsler et al. (Crystal Barrel Collaboration), Phys. Lett. B 355, 425 (1995).
[28] C.A. Meyer, "Proceedings of the Workshop on Gluonic Excitations," Newport News, Virginia 2003, AIP Conf. Proc. 698, 554 (2004).
[29] K.K. Seth, Nucl. Phys. B (Proc. Suppl.) 96, 205 (2001).
[30] Bing-Song Zou, Nucl. Phys. A644, 41c (1999).
[31] C. Amsler and F.E. Close, Phys. Lett. B 353, 385 (1995).
[32] D. Weingarten, Nucl. Phys. Proc. Suppl. 53, 232 (1997).
[33] D.V. Bugg, M.J. Peardon and B.S. Zou, Phys. Lett. B 486, 49 (2000).
[34] J. Sexton, A. Vaccarino and D. Weingarten, Phys. Rev. Lett. 75, 4563 (1995).
[35] V.V. Anisovich, Phys. Lett. B 364, 195 (1995).
[36] F. Giacosa, T. Gutsche and A. Faessler, Phys. Rev. C 71, 025202 (2005).
[37] M. Chanowitz, hep-ph/0506125.
[38] A.H. Fariborz, Int. J. Mod. Phys. A 19, 5417 (2004).
[39] N. Brambilla et al., hep-ph/0412158.
[40] S. Eidelman et al., (Particle Data Group Collaboration), Phys. Lett. B 592, 1 (2004).
[41] G. Alexander et al. (Pluto Collaboration), Phys. Lett. B 76, 652 (1978).
[42] D.L. Scharre, " $10^{\text {th }}$ International Symposium on Lepton and Photon Interactions at High Energy", Bonn (1981).
[43] C. Edwards et al. (Crystal Ball Collaboration), Phys. Rev. D 25, 3065 (1982).
[44] J.E. Augustin et al. (DM2 Collaboration), Z. Phys. C 36, 369 (1987).
[45] R.M. Baltrusaitis et al. (Mark III Collaboration), Phys. Rev. D 35, 2077 (1987).
[46] J.Z. Bai et al. (BES Collaboration), Phys. Rev. D 68, 052003 (2003).
[47] M. Krammer, Phys. Lett. B 74, 361 (1978).
[48] J.G. Korner, J.H. Kuhn, M. Krammer and H. Schneider, Nucl. Phys. B 229, 115 (1983).
[49] J.Z. Bai et al., (BES Collaboration), Phys. Rev. Lett. 76, 3502 (1996).
[50] R.M. Baltrusaitis et al. (MARK-III Collaboration), Phys. Rev. Lett. 56, 107 (1986).
[51] J.E. Augustin et al. (DM2 Collaboration), Phys. Rev. Lett. 60, 2238 (1988).
[52] J.Z. Bai et al. (BES Collaboration), Phys. Rev. Lett. 91, 022001 (2003).
[53] A. Sibirtsev, J. Haidenbauer, S. Krewald, U.G. Meissner and A.W. Thomas, Phys. Rev. D 71, 054010 (2005).
[54] B. Kerbikov, A. Stavinsky and V. Fedotov, Phys. Rev. C 69, 055205 (2004).
[55] D.V. Bugg, Phys. Lett. B 598, 8 (2004).
[56] C.S. Gao and S.L. Zhu, Commun. Theor. Phys. 42, 844 (2004).
[57] B.S. Zou and H.C. Chiang, Phys. Rev. D 69, 034004 (2004).
[58] B. Loiseau and S. Wycech, hep-ph/0501112.
[59] G.J. Ding and M.L. Yan, hep-ph/0502127.
[60] A. Anastassov et al., (CLEO Collaboration), Phys. Rev. Lett. 82, 286 (1999).
[61] G. Masek et al., (CLEO Collaboration), Phys. Rev. D 65, 072002 (2002).
[62] L. Breva-Newell, Ph.D. Thesis, University of Florida, hep-ex/0412075.
[63] G. Viehhauser, Nucl. Inst. Meth. A 462, 146 (2001); D. Peterson et al., Nucl. Inst. Meth. A 478, 142 (2002); A. Warburton et al., Nucl. Inst. and Meth. A 488, 451 (2002); M. Artuso et al., physics/0506132.
[64] R.A. Briere et al., (CLEO Collaboration), Phys. Rev. D 70, 072001 (2004).
[65] $Q Q$ - The CLEO Event Generator, http://www.lns.cornell.edu/public/CLEO/soft/QQ (unpublished).
[66] R. Brun et al., Geant 3.15, CERN Report No. DD/EE/84-1 (1987).
[67] T. Sjöstrand, Comput. Phys. Commun. 82, 74 (1994); T. Sjöstrand and M. Bengtsson ibid 43, 367 (1987); T. Sjöstrand, ibid 39, 347 (1986).
[68] M.A. Selen, R.M. Hans and M.J. Haney, IEEE Trans. Nucl. Sci. 48, 562 (2001).
[69] J.D. Richman CALT-68-1148 (1995) (unpublished).
[70] P.K. Kabir and A.J.G. Hey, Phys. Rev. D 13, 3161 (1976).
[71] M. Ablikim et al. (BES Collaboration), Phys. Rev. D 70, 092004 (2004)
[72] S.A. Dytman (CLEO Collaboration), hep-ex/0307035.
[73] S. Godfrey and J. Napolitano, Rev. Mod. Phys. 71, 1411 (1999).
[74] F.E. Close, G.R. Farrar and Z.p. Li, Phys. Rev. D 55, 5749 (1997).


[^0]:    *Submitted to the XXII International Symposium on Lepton and Photon Interactions at High Energies, June 30-July 5, 2005, Uppsala, Sweden

[^1]:    ${ }^{1}$ Several authors have studied meson production in $\Upsilon(1 \mathrm{~S})$ radiative decays, giving predictions for branching and helicity production ratios. The heavy-quarkonia system is usually described by non-relativistic QCD [1], while the gluonic hadronization has been treated using soft collinear effective theory [2], gluon distribution amplitudes [3], and perturbative QCD [4, 5].
    ${ }^{2}$ Glueballs are a natural consequence of QCD, and predictions of their properties have been made using different approaches, such as, potential models [6-8], lattice QCD calculations [9-12], bag models [13-16], flux-tube models [17], the QCD sum rule [18], the Bethe-Salpeter (B-S) equation [19, 20], QCD factorization formalism models [21, 22], weakly-bound-state models [23], and a three-dimensional relativistic equation [24]. However, despite intense experimental searches [25-30], there is no conclusive experimental evidence of their observation, although there are strong indications that glueballs contribute to the rich light scalar spectrum [31-38].
    ${ }^{3}$ We refer to the Particle Data Group [40] for a summary of $J / \psi$ radiative decays.

[^2]:    ${ }^{4}$ We obtain this factor, $f$, from the integrated luminosities of the the $\Upsilon(1 \mathrm{~S})$ and $\sqrt{s}=10.56 \mathrm{GeV}$ datasets, and the assumption that, to first order, the cross sections of the continuum processes in each run are proportional to $1 / s$. This factor is roughly equal to the factor obtained by using the average energy of

[^3]:    ${ }^{6}$ Some authors use a probability distribution that also depends on a third angle, $\phi_{h}$ [45]. However, extreme care must be taken when using this angle because it makes the probability distribution sensitive to the relative phases of the helicity amplitudes. Thus, two new free parameters need to be introduced in such a probability distribution, as was noted by [48] and correctly implemented by [44, 45]. Otherwise, the measurement of the helicity amplitudes rests on the assumption that their relative phases are 0 [41, 43, 70, 71].

[^4]:    ${ }^{7}$ We choose to use a simultaneous fit to these two distributions instead of a two-dimensional fit using Equation 2 because of our limited statistics.

