

7.5

STANDARD CANDLE \Rightarrow LUMINOSITY KNOWN
 STANDARD YARDSTICK \Rightarrow PROPER LENGTH KNOWN

$$\text{ie. (7.28) } d_L = S_K(r)(1+z)$$

$$(7.36) \quad d_A = S_K(r)(1+z)^{-1}$$

$$\Rightarrow (7.37) \quad d_A = d_L (1+z)^{-2}$$

DEFINITIONS

$$(7.21) \quad d_L = \left(\frac{L}{4\pi f} \right)^{1/2}$$

$$(7.33) \quad d_A = \frac{\rho}{8\theta}$$

$$\Sigma \sim \frac{f}{(8\theta)^2} = \left(\frac{L}{4\pi d_L^2} \right) \left(\frac{d_A}{\rho} \right)^2$$

$$= \frac{L}{4\pi d_L^2} \cdot \frac{d_L^2}{\rho^2 (1+z)^4}$$

$$= \boxed{\frac{L}{4\pi \rho^2 (1+z)^4}}$$

7.6 from (3.44) & (3.45) $\delta t_e = \delta t_o (1+z)^{-1}$

$$\Rightarrow \boxed{\delta t_e = 12 \text{ hrs}}$$

$$\boxed{R_{\text{max}} = c \delta t_e = 87 \text{ Au}}$$

$$(7.38) \quad d_A = \frac{R_{\text{max}}}{8\theta} = \frac{d_p(t_o)}{1+z} \quad (7.38)$$

$$(7.13) \quad d_p(t_o) = c \int_{t_o}^0 \frac{dt}{a(t)} = c \int_{a(t_o)}^1 \left(\frac{dt}{da} \right) \frac{da}{a(t)}$$

$$= c \int_{1/6}^1 \frac{da}{a^2}$$

Simplify using Friedmann equation for benchmark model

$$\frac{H(t)^2}{H_0^2} = \frac{\Omega_{r,0}}{a^4} + \frac{\Omega_{m,0}}{a^3} + \Omega_{\Lambda,0}$$

$$\dot{a} = H_0 \sqrt{\Omega_{r,0} a^{-2} + \Omega_{m,0} a^{-1} + \Omega_{\Lambda,0} a^2}$$

flws

$$d_p(t_0) = c \int_{116}^1 \frac{da}{H_0 \sqrt{\Omega_{r,0} + \Omega_{m,0} a + \Omega_{\Lambda,0} a^4}}$$

\uparrow \uparrow \uparrow
 v_0 $.3$ $.7$

$$d_p(t_0) \approx 7800 \text{ Mpc}$$

$$d_A = d_p / 1.2 = 1300 \text{ Mpc}$$

$$\delta\theta = \frac{R_{max}}{d_A} = 3.2 \times 10^{-13} \text{ rad}$$

8.2

$$L = (1.8 \pm 0.8) \times 10^5 L_\odot$$

\uparrow SPHERE OF RADIUS $r_s = 120 \pm 12 \text{ pc}$ CONTAINS $\frac{1}{2} L$

$$(8.38) \quad \sigma_r = \langle (v_r - \langle v_r \rangle)^2 \rangle^{1/2}$$

$$\text{ISOTROPIC} \Rightarrow \langle v^2 \rangle = 3\sigma_r^2$$

$$(8.34) \quad M = \frac{\langle v^2 \rangle r_s}{2G} = \boxed{2.3 \times 10^7 M_\odot}$$

\uparrow $d \approx .4$ (R135)

$$\boxed{\frac{M}{L} = 128 M_\odot / L_\odot}$$

8.3 (8.48) $\alpha = \frac{4GM}{c^2 b}$

EARTH :	$\alpha = 2.8 \times 10^{-9}$ RAD	=	5.7×10^{-4} ARCSEC
U-DWARF :	$\alpha = 3.95 \times 10^{-4}$ RAD	=	81.5°
N. STAR :	$\alpha = 0.78$ RAD	=	42.5°

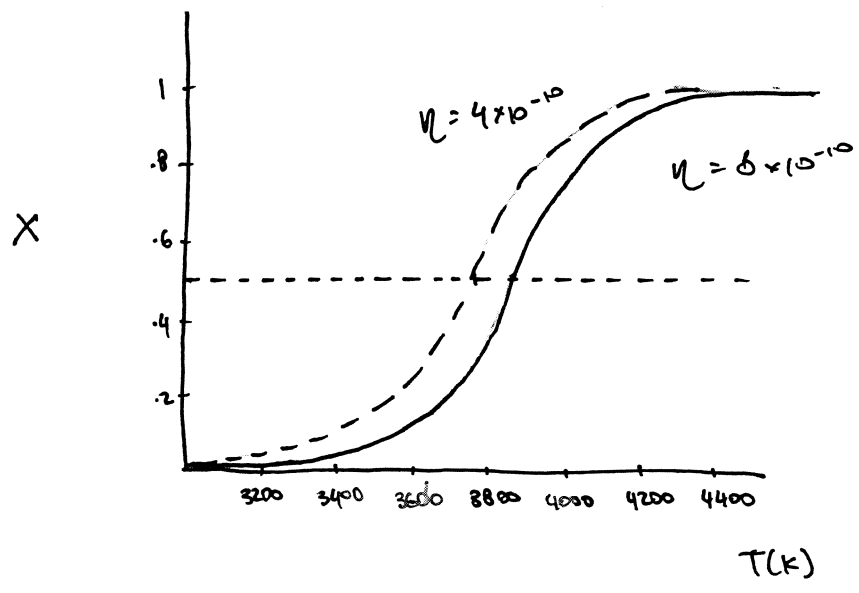
8.5 $n = \frac{N}{\frac{4}{3}\pi r^3} = \boxed{70.74 \text{ gal/Mpc}^3}$

$l = \frac{1}{n^2} = \boxed{14.14 \text{ Mpc}}$

$\tau = \frac{l}{v} = \frac{l}{\sqrt{3}c} = \boxed{9 \text{ Gyr}}$

9.1 (9.30) $X = \frac{-1 + \sqrt{1+4S}}{2S}$

(9.31) $S(T, \eta) = 3.84 \eta \left(\frac{kT}{m_0 c^2}\right)^{3/2} e^{a/kT}$



9.2 (2.25) $\epsilon(v) dv = \frac{8\pi h}{c^3} \frac{v^3 dv}{\exp(hv/kT) - 1}$

$n(v) dv = \frac{1}{hv} \epsilon(v)$
 $= \frac{8\pi}{c^3} \frac{v^2 dv}{\exp(hv/kT) - 1}$

for $kT \ll Q$

This is a very important integral in physics!

$N_{\text{ionize}} = \int_Q^{\infty} n(v) dv$
 $\approx \frac{8\pi}{c^3} \int_Q^{\infty} \frac{v^2 dv}{\exp(hv/kT)}$
 $= \frac{8\pi}{c^3} \int_Q^{\infty} v^2 \exp(-hv/kT) dv$
 $= \frac{8\pi}{c^3} \left(\frac{kT}{h}\right)^3 \int_{hQ/kT}^{\infty} u^2 e^{-u} du$
 $= 8\pi \left(\frac{kT}{ch}\right)^3 \left\{ -u^2 e^{-u} \Big|_{hQ/kT}^{\infty} + \int_{hQ/kT}^{\infty} 2ue^{-u} du \right\}$
 $\quad \quad \quad \uparrow$
 $2 \left\{ -ue^{-u} \Big|_{hQ/kT}^{\infty} + \int_{hQ/kT}^{\infty} e^{-u} du \right\}$
 $= 8\pi \left(\frac{kT}{ch}\right)^3 \left\{ \left(\frac{hQ}{kT}\right)^2 e^{-hQ/kT} + 2\left(\frac{hQ}{kT}\right) e^{-hQ/kT} + e^{-hQ/kT} \right\}$

$N_{\text{tot}} = \int_0^{\infty} n(v) dv$
 $= \beta T^3 \quad (2.28)$

$\uparrow \beta = \frac{2.404}{\pi^2} \frac{k^3}{15c^3} = 2.03 \times 10^7 \text{ m}^{-3} \text{ K}^{-3} \quad (2.29)$

$f = \frac{N_{\text{ionize}}}{N_{\text{tot}}}$

9.4 $z_{LS} = 1100$

USING THE SAME MANIPULATIONS AS IN QUESTION 7.6

$$d_p(t_0) = c \int_{1/1101}^1 (H_0 \sqrt{\Omega_{r,0} + \Omega_{M,0} a + \Omega_{\Lambda,0} a^4})^{-1} da$$

\uparrow \uparrow \uparrow
 ~ 0 $\sim .8$ $\sim .7$

$$d_c(t_0) = (1+z) d_p(t_0)$$

$$= 1101 \cdot d_p(t_0)$$

9.5 (9.61) $\tau = \int_{t_*}^{t_0} n_{\text{e}}(t) \sigma_e c dt$

$$\uparrow$$

$$n_{\text{e}}(t) = \frac{n_{\text{bary},0}}{a^3}$$

$$= \int_{t_*}^{t_0} a(t)^3 \underbrace{n_{\text{bary},0}}_{4.4 \times 10^{-21} \text{ s}^{-1}} \sigma_e c dt$$

from (9.34)

FOR A FLAT MATTER-ONLY UNIVERSE
(559) $t_0 = \frac{2}{3H_0}$ $a(t) = (t/t_0)^{2/3}$

$$\Rightarrow \tau = n_{\text{bary},0} \sigma_e c t_0^2 \left(\frac{1}{t_*} - \frac{1}{t_0} \right)$$

$$1 = (4.4 \times 10^{-21} \text{ s}^{-1}) \left(\frac{2}{3H_0} \right)^2 \left(\frac{1}{t_*} - \left(\frac{2}{3H_0} \right)^{-1} \right)$$

$$\Rightarrow \boxed{t_* = 3.9 \times 10^{14} \text{ s}} \\ \boxed{= 13 \text{ Myr}}$$

$$1+z_* = a_*^{-1} = (t_*/t_0)^{3/2}$$

$$\boxed{z_* \approx 83}$$

50 SHEETS EYE-EASE: 8 SQUARE
100 SHEETS EYE-EASE: 9 SQUARE
200 SHEETS EYE-EASE: 9 SQUARE

